

CRM tests of PBL-based cumulus mass flux closures

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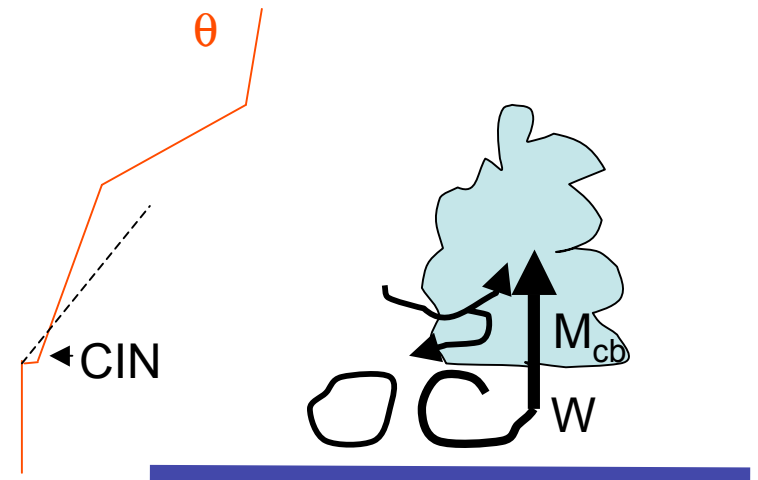
UW Shallow Cu CIN-based mass flux closure

- Cu form above the moistest, strongest boundary-layer updrafts, which have some vertical eddy velocity scale W .
- Convective inhibition (CIN) regulates fraction of these updrafts rising into buoyant Cu:

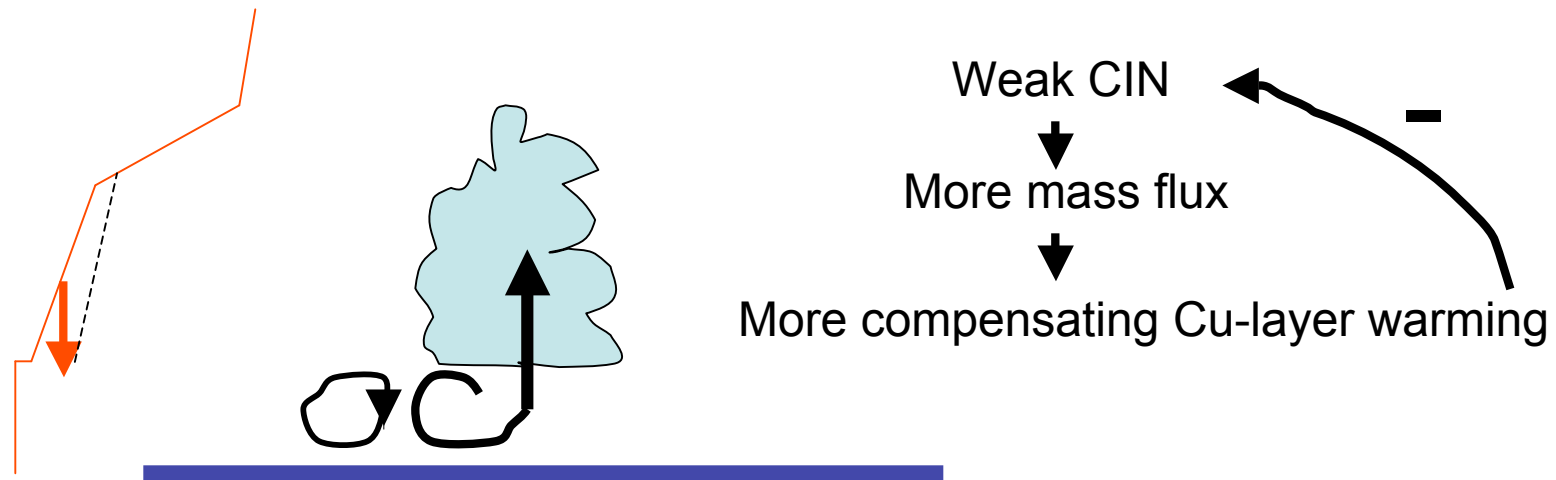
$$M_{cb} = c_1 W \exp(-c_2 \text{CIN}/W^2)$$

(Bretherton et al. 2004; Mapes 2000)

- Cu vertical structure is separately predicted from Cu-layer thermodynamic profiles.



CIN feedback in cumulus convection

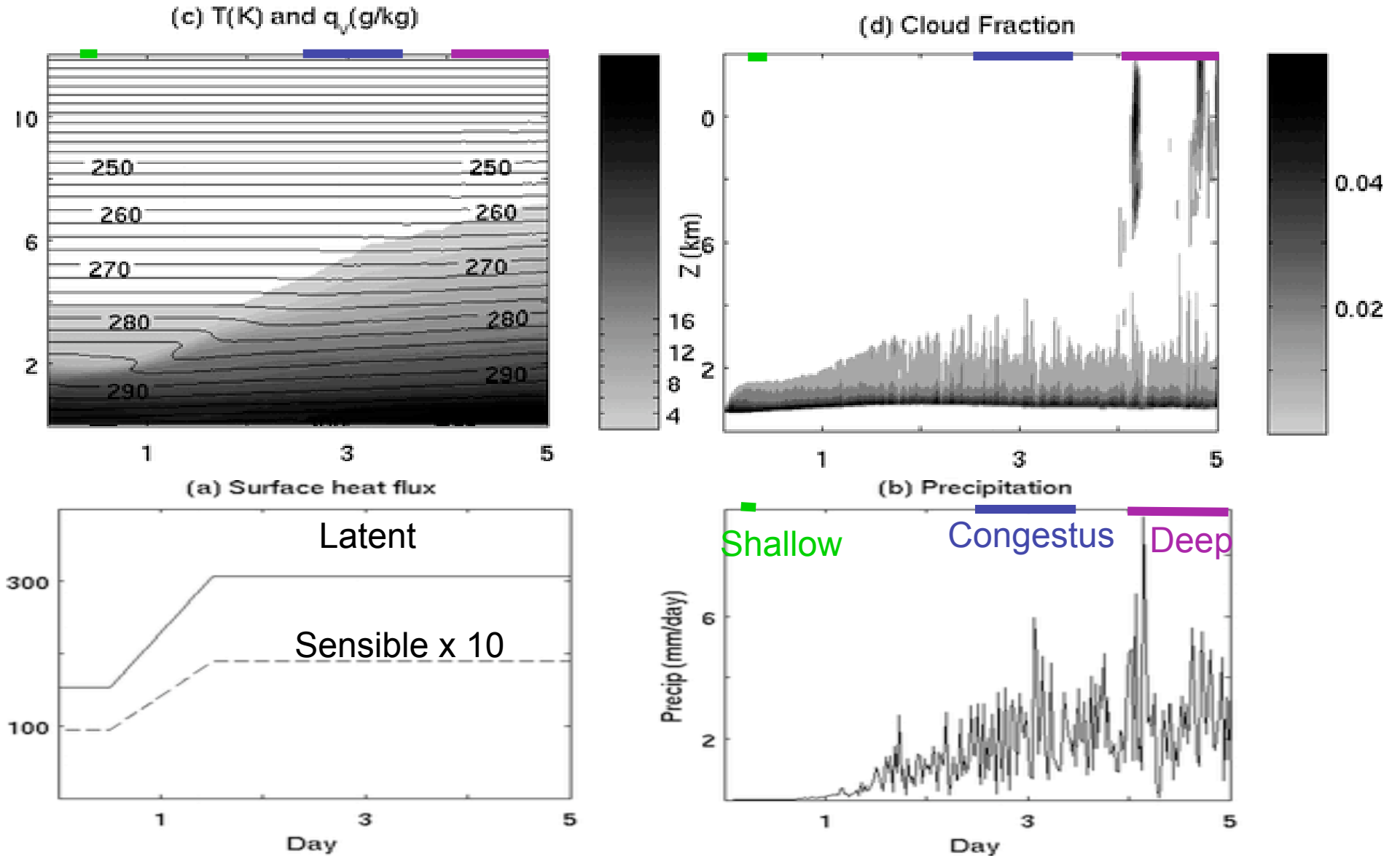


CIN acts as a fast ‘valve’ adjusting cloud base mass flux to keep Cu base near the subcloud layer top.

With slowly-varying surface/large-scale forcings, the Cu-layer thermodynamic and mass flux profiles evolve naturally into a quasi-equilibrium dependent on the cumulus plume model: No need for CAPE closure.

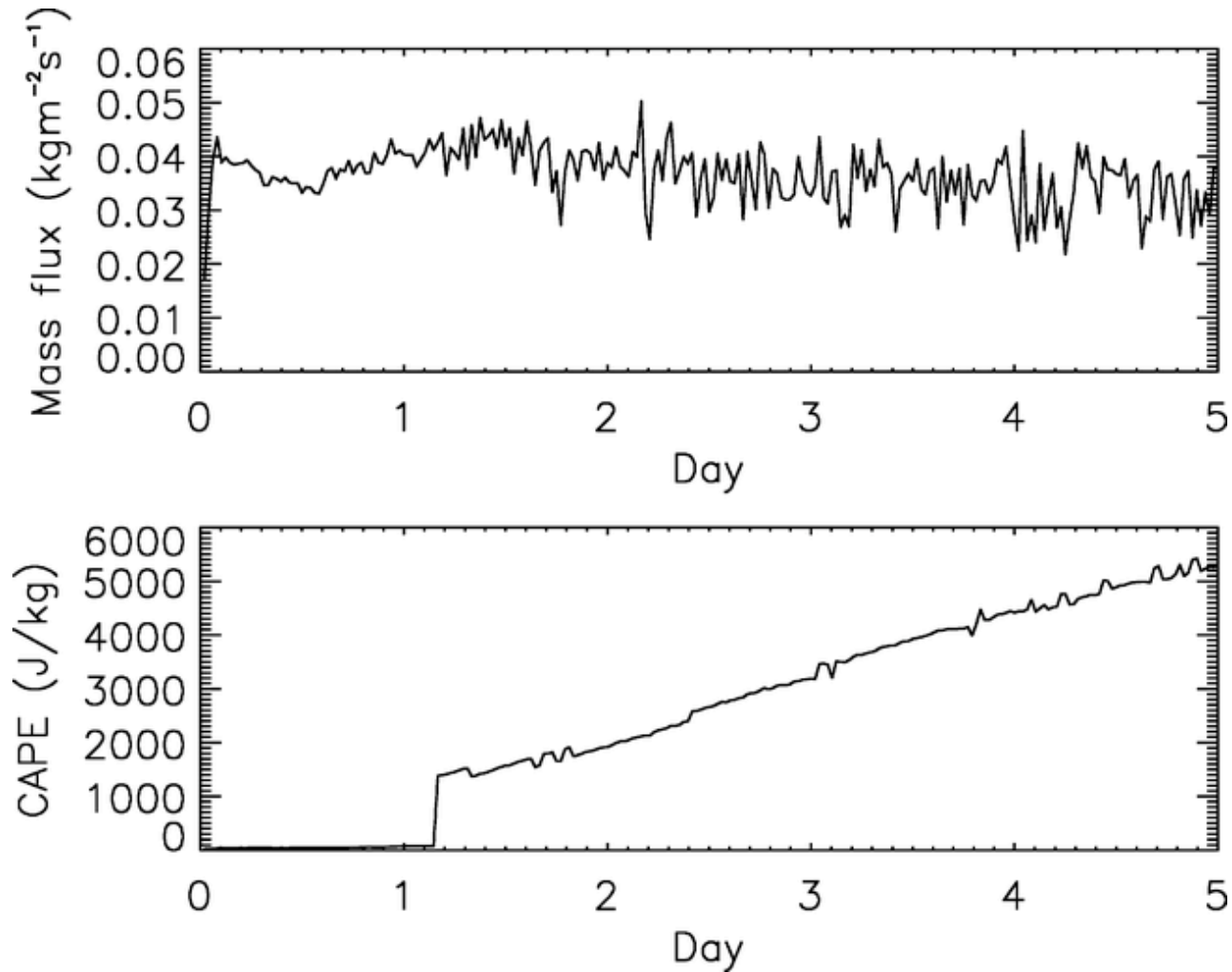
Does this conception work for deep convection?

Kuang and Bretherton (2006): CRM-simulated shallow-to-deep Cu transition



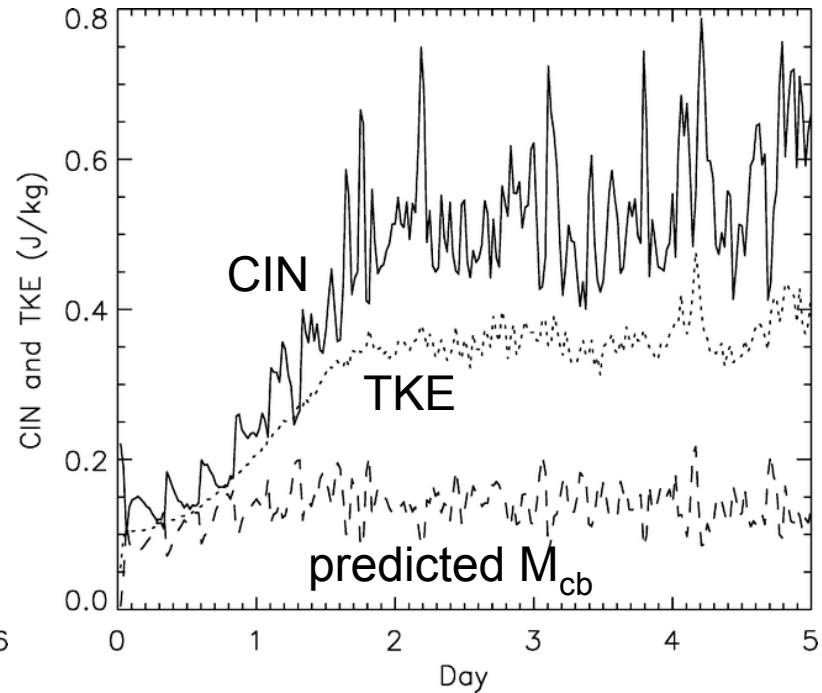
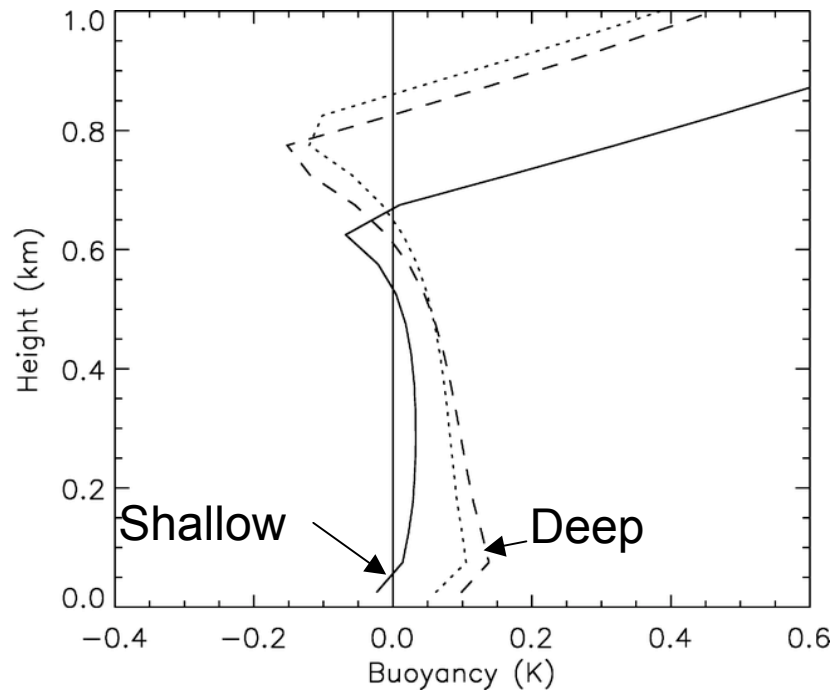
Cloud base mass flux closure

- Mass flux quasi-constant through simulation, but not CAPE.



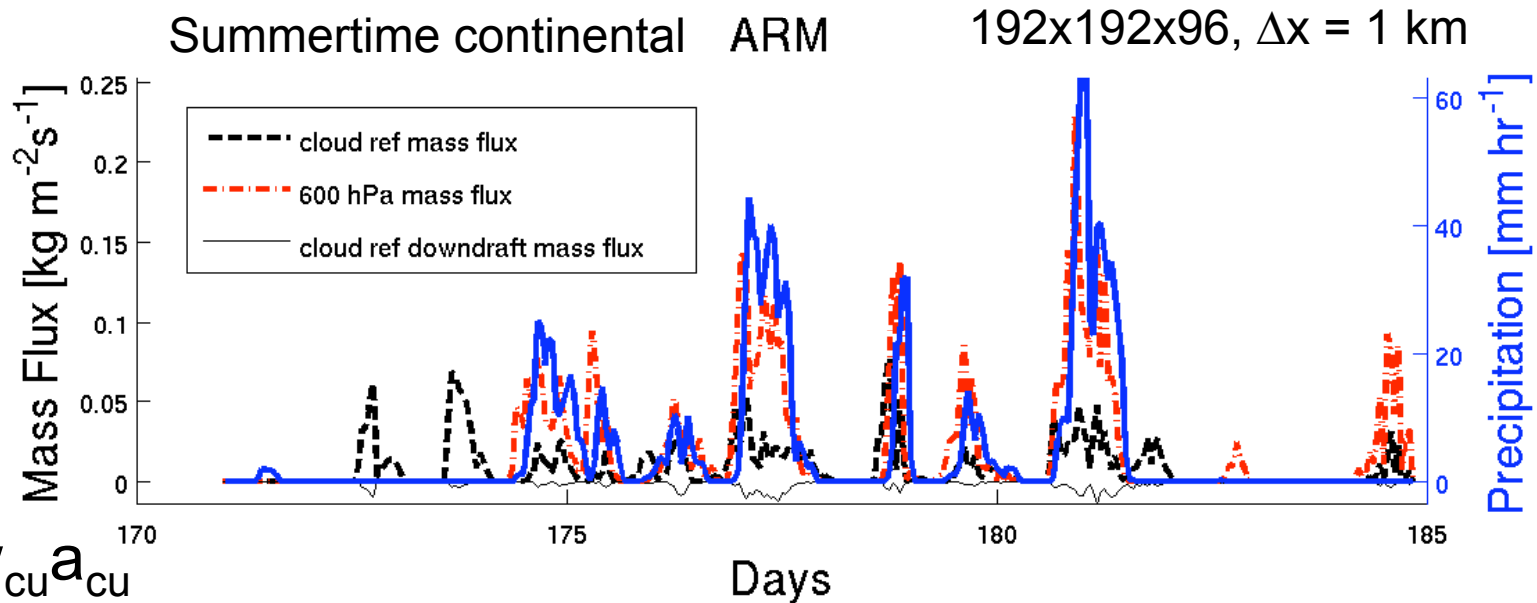
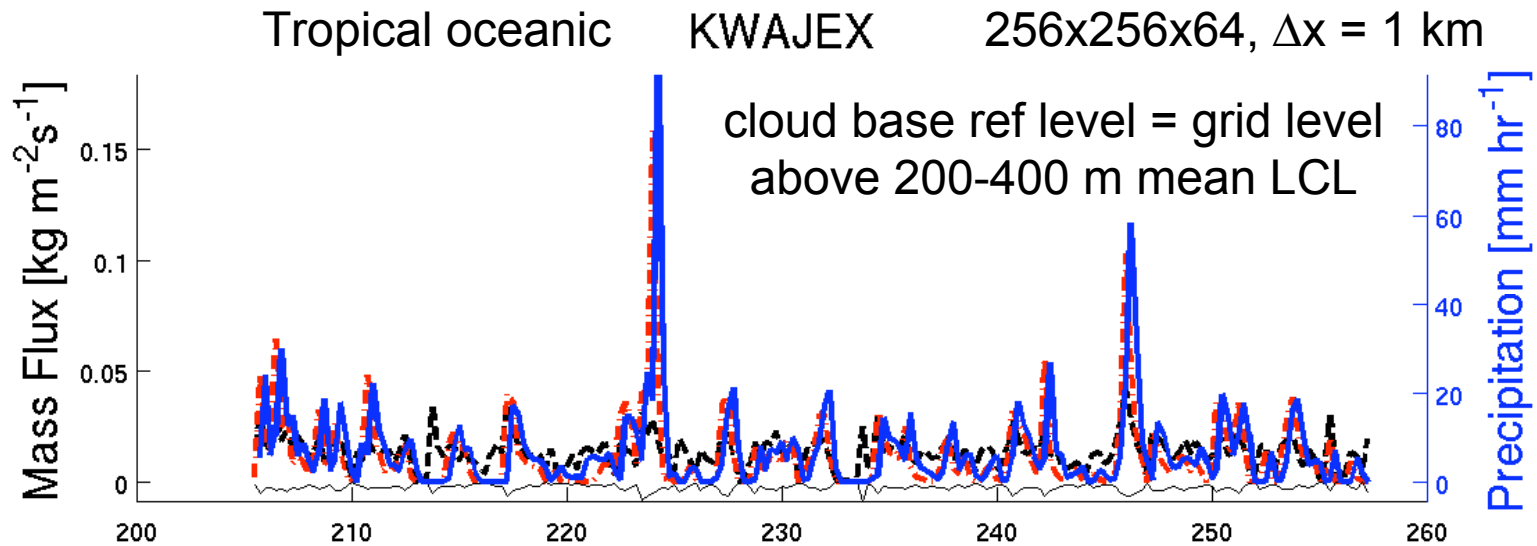
CIN regulation of mass flux

- CIN present in all 3 regimes (based on mean sounding and adiabatically displacing mean cloud base air parcel).
- Nicely follows $M_{cb} = c_1 W \exp(-CIN/W^2)$, $W^2 = TKE_{PBL}$

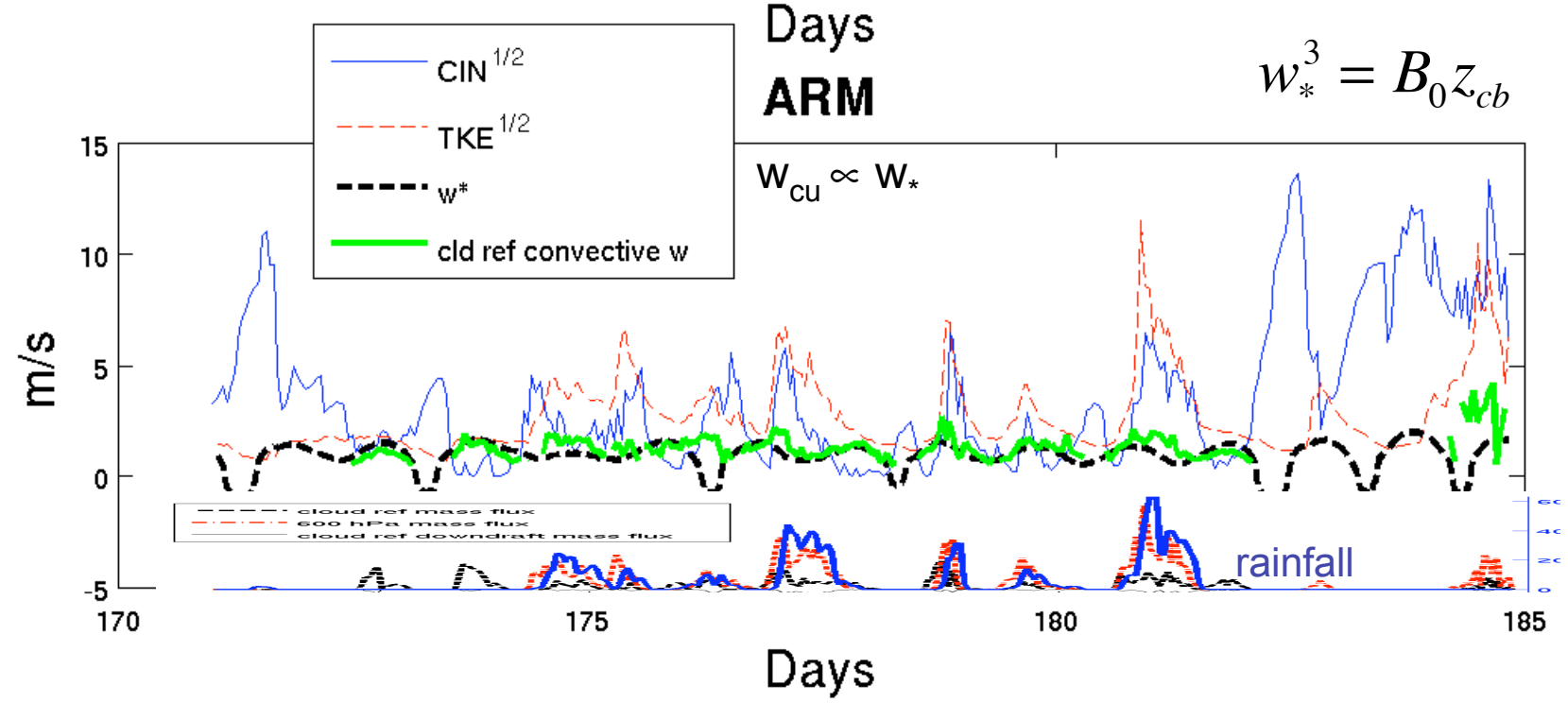
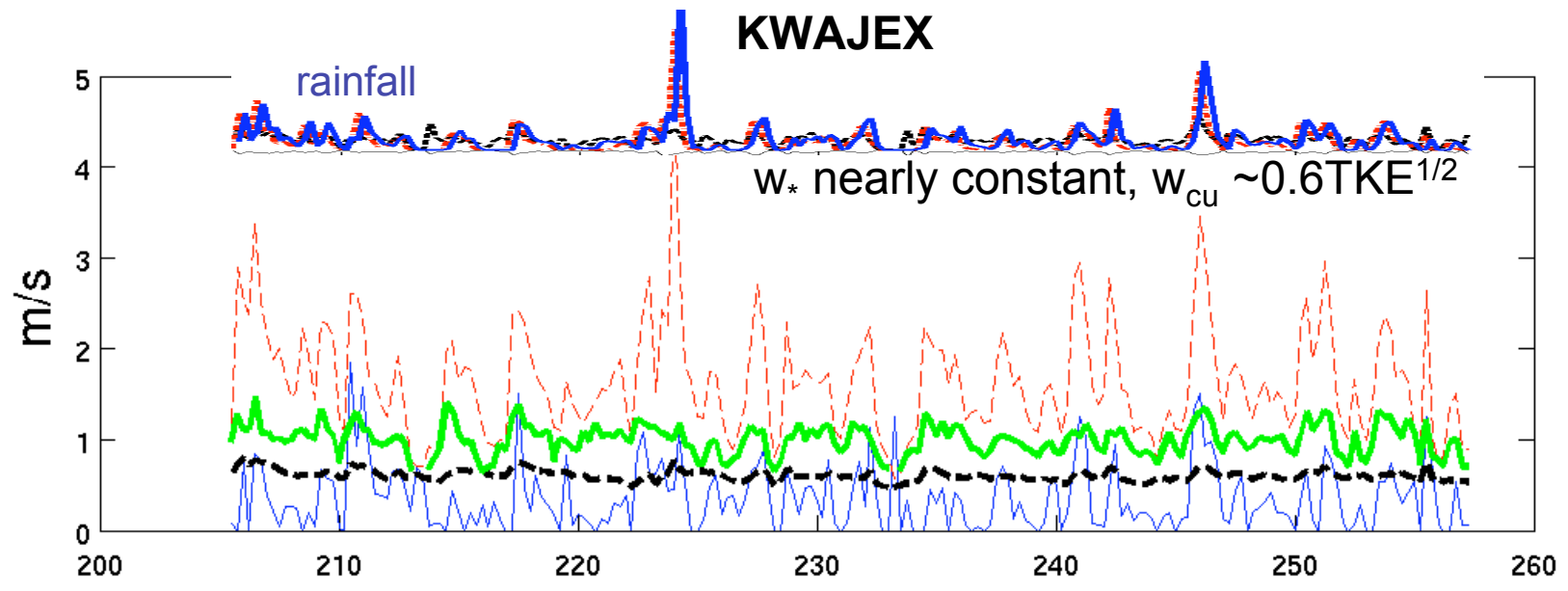


...but convection not that deep or rainy, and case is idealized.

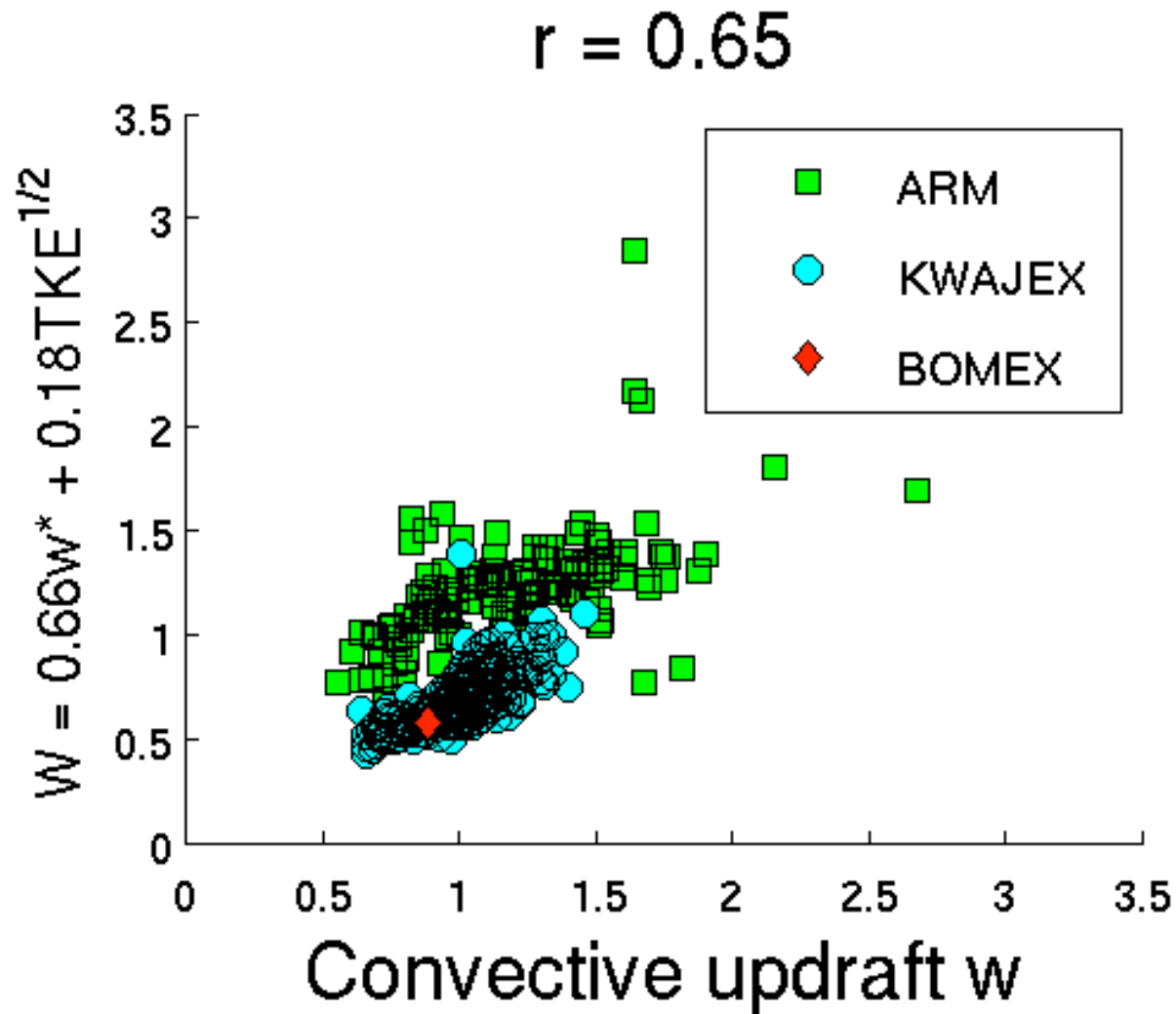
Two CRM-simulated 'real' deep Cu cases



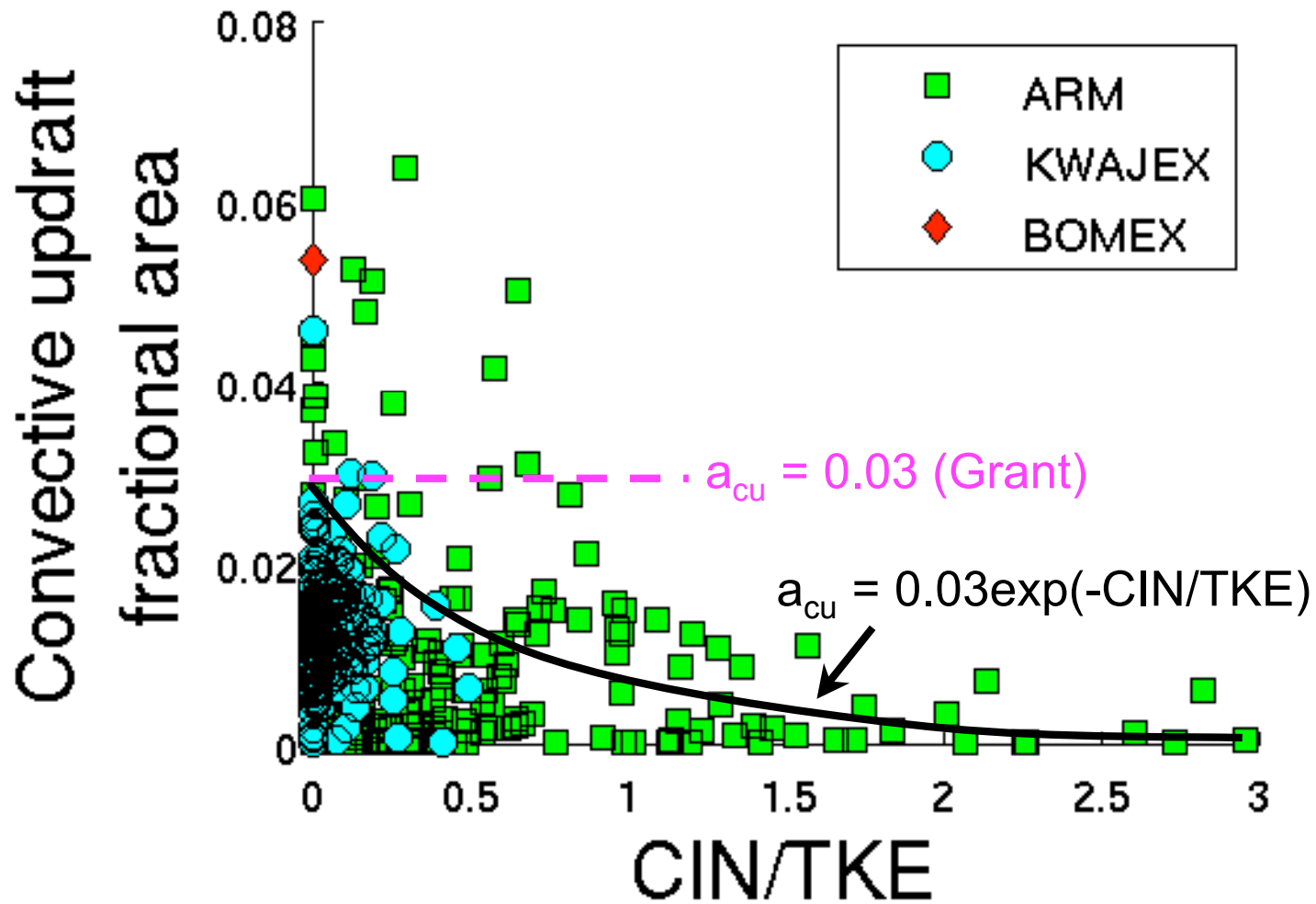
$$M_{\text{cu}} = w_{\text{cu}} a_{\text{cu}}$$



Updraft W best predicted using both w_* and $\text{TKE}^{1/2}$



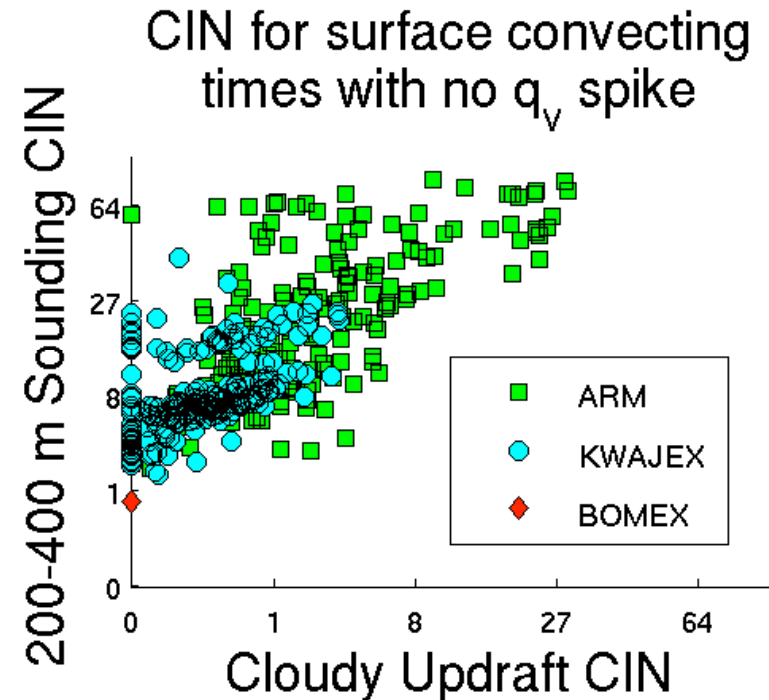
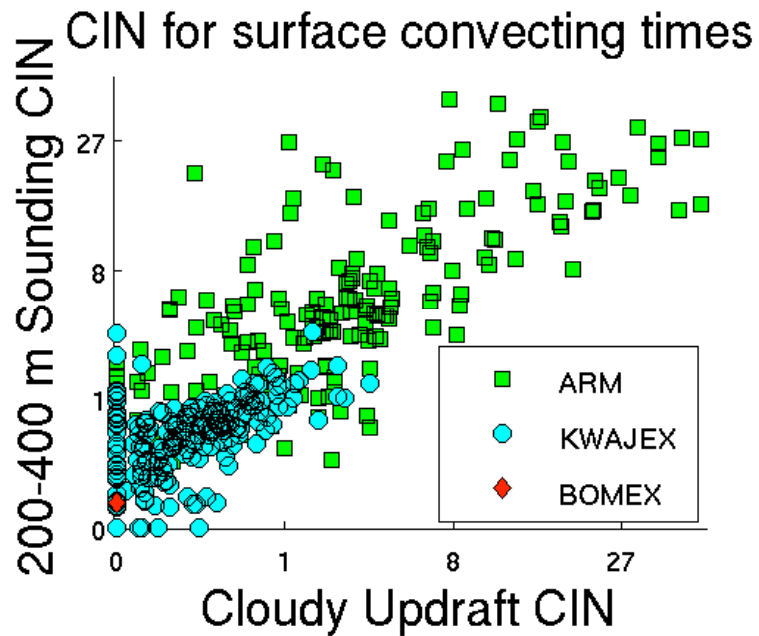
Surface Convection times only



Predicting cloudy updraft CIN

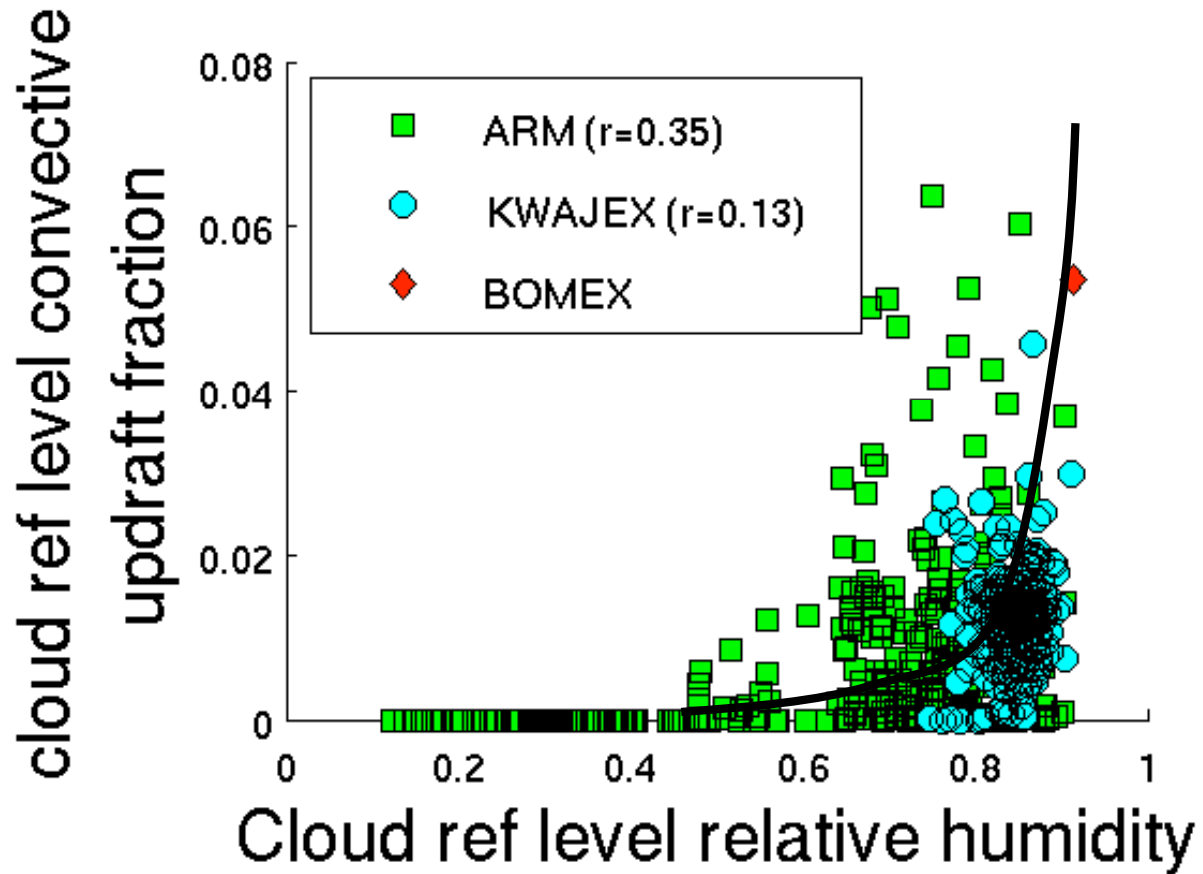
- Empirically, 200-400 m level works better than near-surface air for predicting cloudy updraft properties:

$$\theta_{cu} \sim \theta_{200-400}, \quad q_{cu} \sim q_{200-400} + \sigma_q$$



Cloud fraction consistency

Simulated cases show a consistent relation between RH and cloudy updraft area fraction near cloud base.



A typical diagnostic cloud scheme coupled to a CIN-based mass-flux scheme may not enforce this consistency.

Conclusions

- CIN-based closure appears viable for deep convection.
- A plausible approach is:

$$M_{\text{cu}} = w_{\text{cu}} a_{\text{cu}}$$

$$W_{\text{cu}} = b_1 w_* + b_2 \text{TKE}^{1/2}$$

$$a_{\text{cu}} = 0.03 \exp(-c_2 \text{CIN}_{\text{cu}} / W_{\text{cu}}^2) \text{ or } f(\text{RH}_{\text{cb}}) \exp(-c_2 \text{CIN}_{\text{cu}} / W_{\text{cu}}^2)$$

UWShCu uses the first of these, with $b_1=0$, $b_2=1$, $c_2=1$.

- Care is required to ensure the computation of cloud fraction is consistent with the computation of cumulus updraft fraction near the cumulus base.