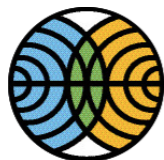


WACCM studies at FMI

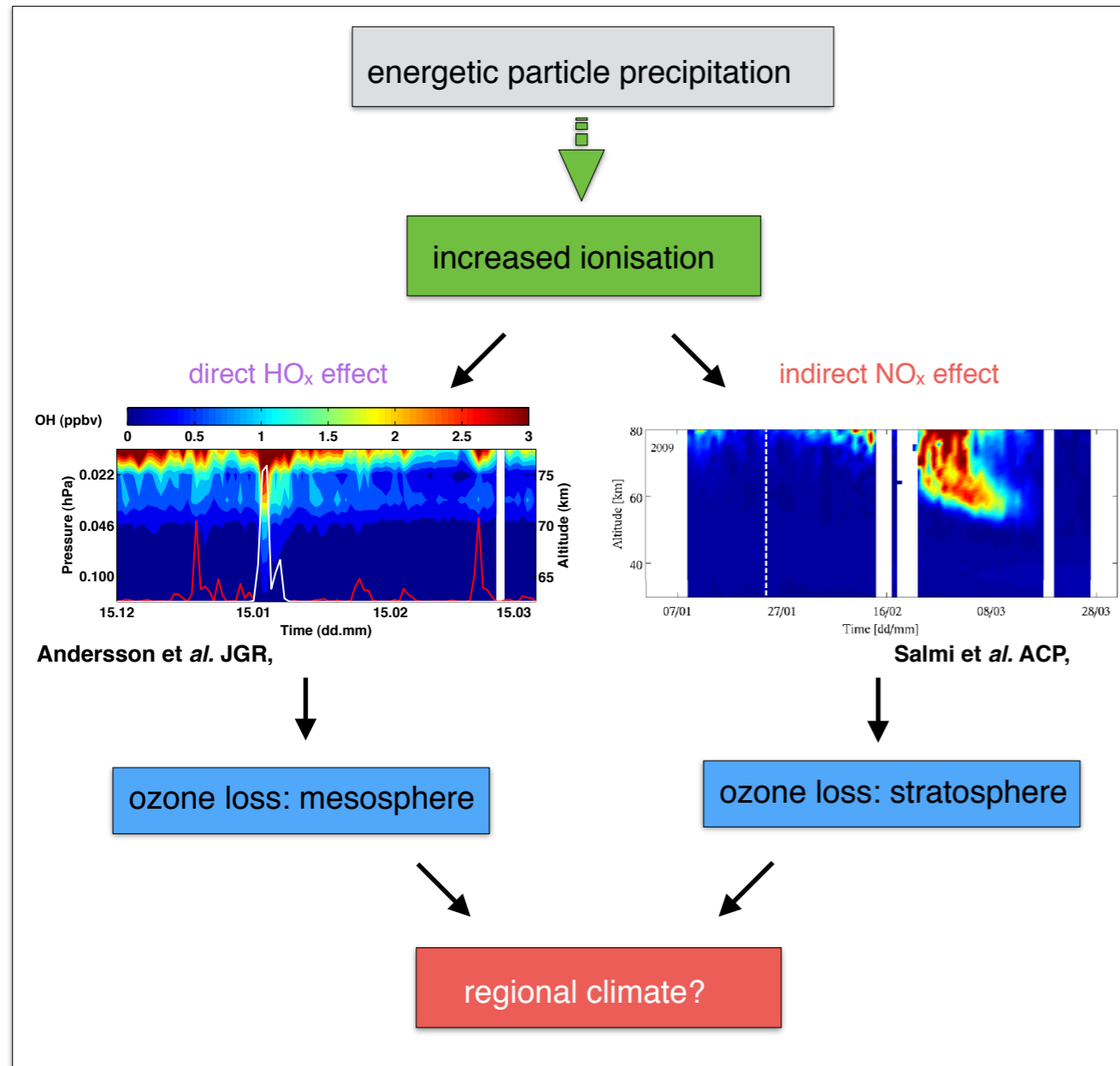
M. Andersson, P.T. Verronen, D.R. Marsh, Annika Seppälä, Niilo Kalakoski, Sanna-Mari Päivärinta, J.M.C. Plane, T. Kovacs,



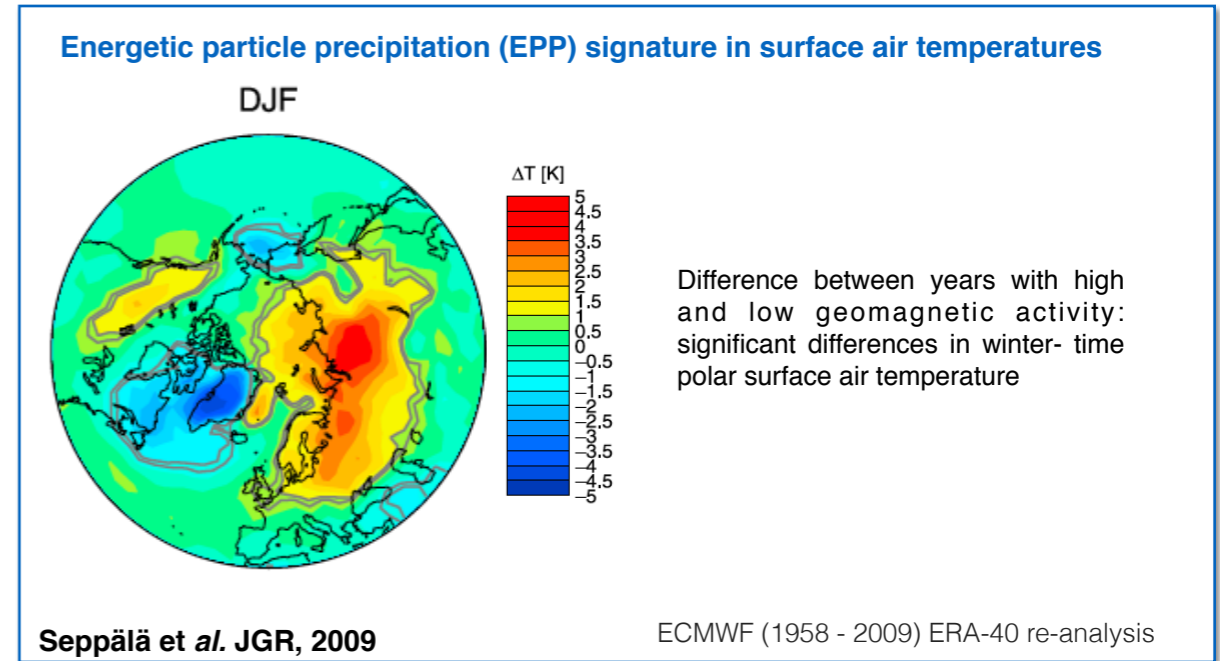
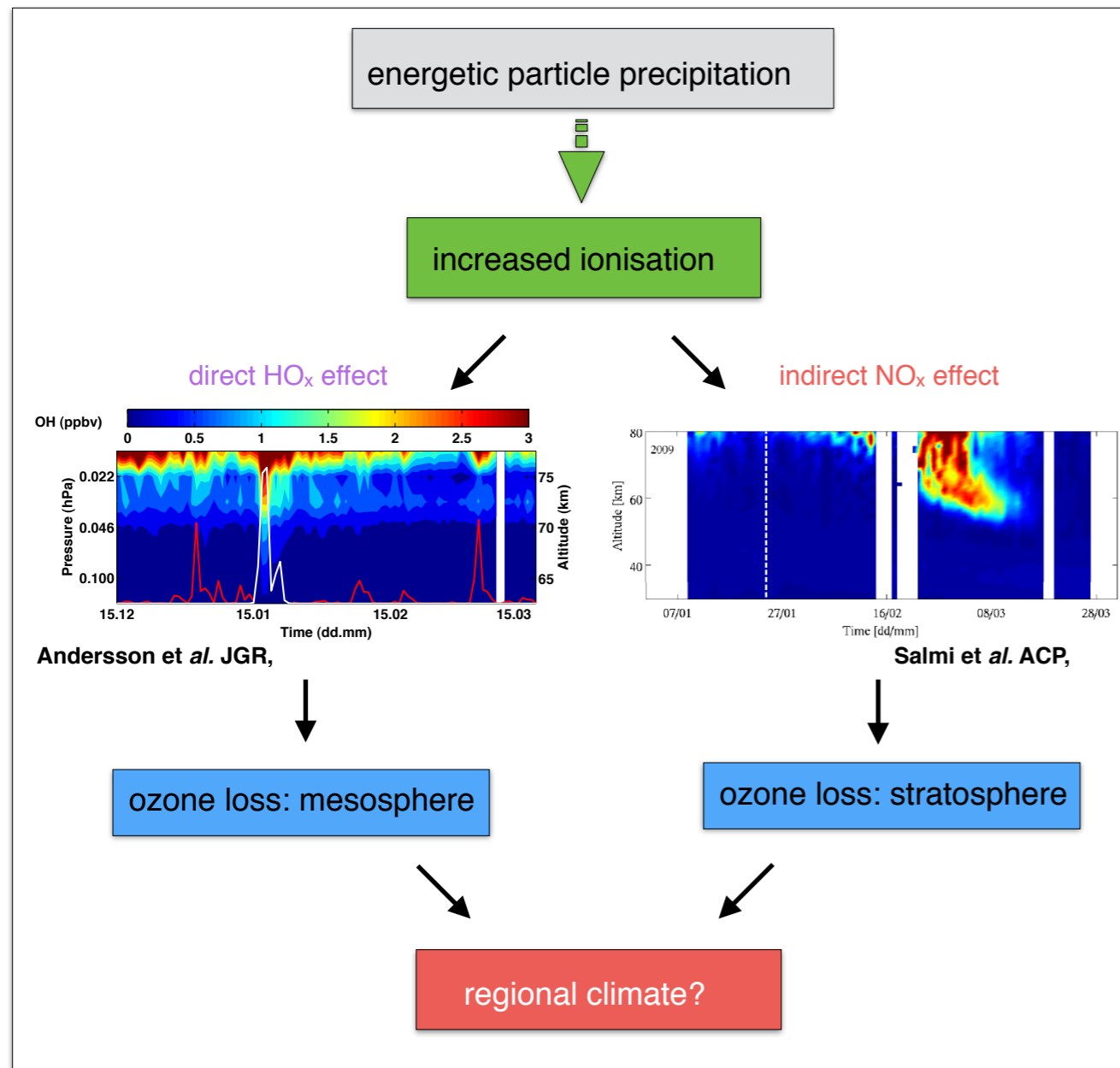
Finnish Meteorological Institute



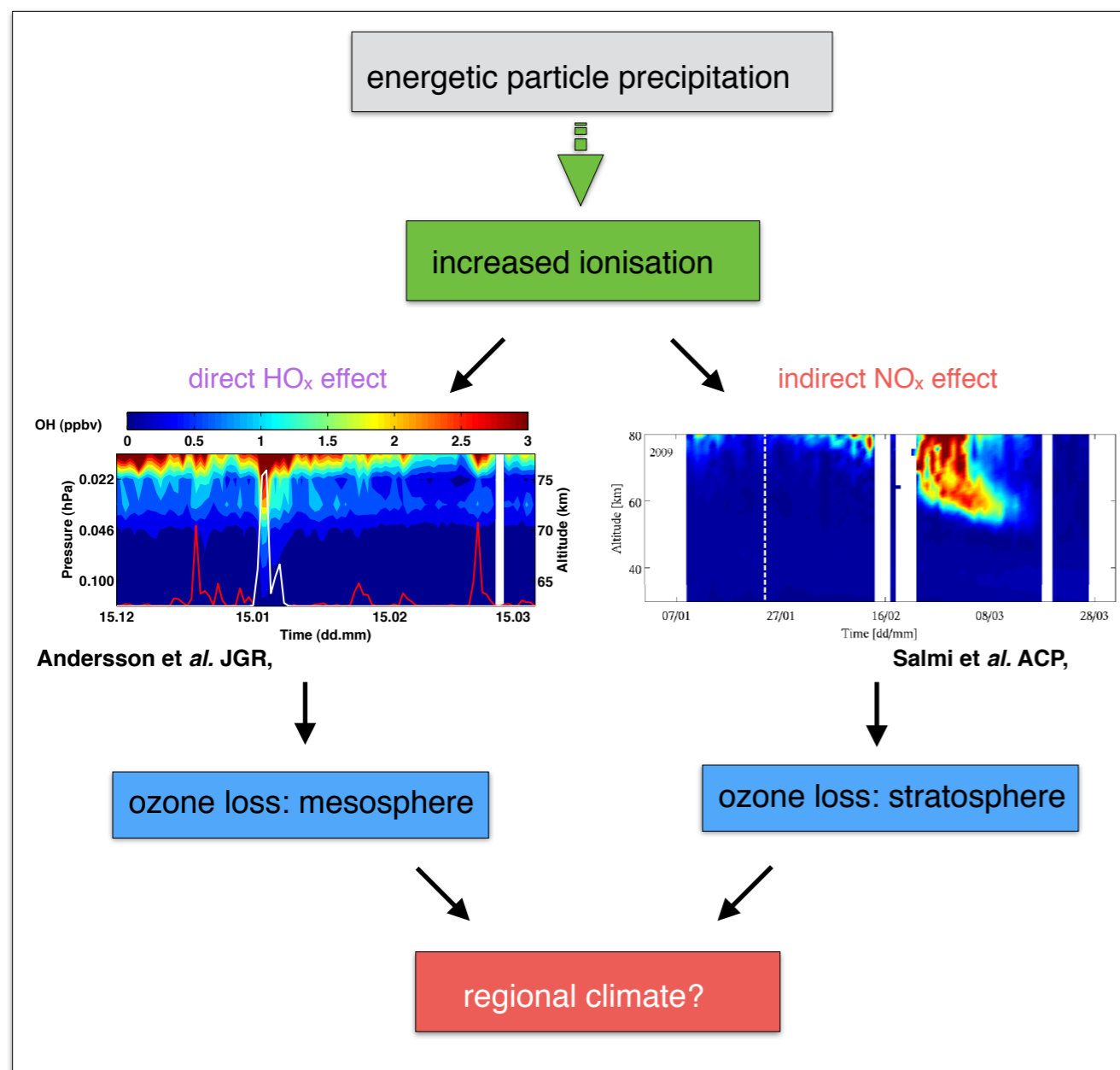
Energetic Particle Precipitation (EPP) mechanism



Energetic Particle Precipitation (EPP) mechanism

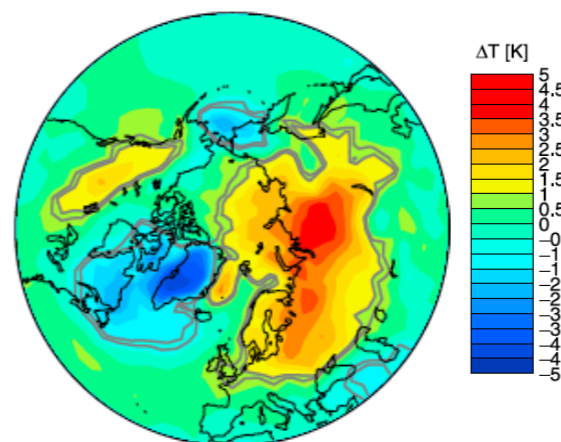


Energetic Particle Precipitation (EPP) mechanism



Energetic particle precipitation (EPP) signature in surface air temperatures

DJF



Difference between years with high and low geomagnetic activity: significant differences in winter-time polar surface air temperature

Seppälä et al. JGR, 2009

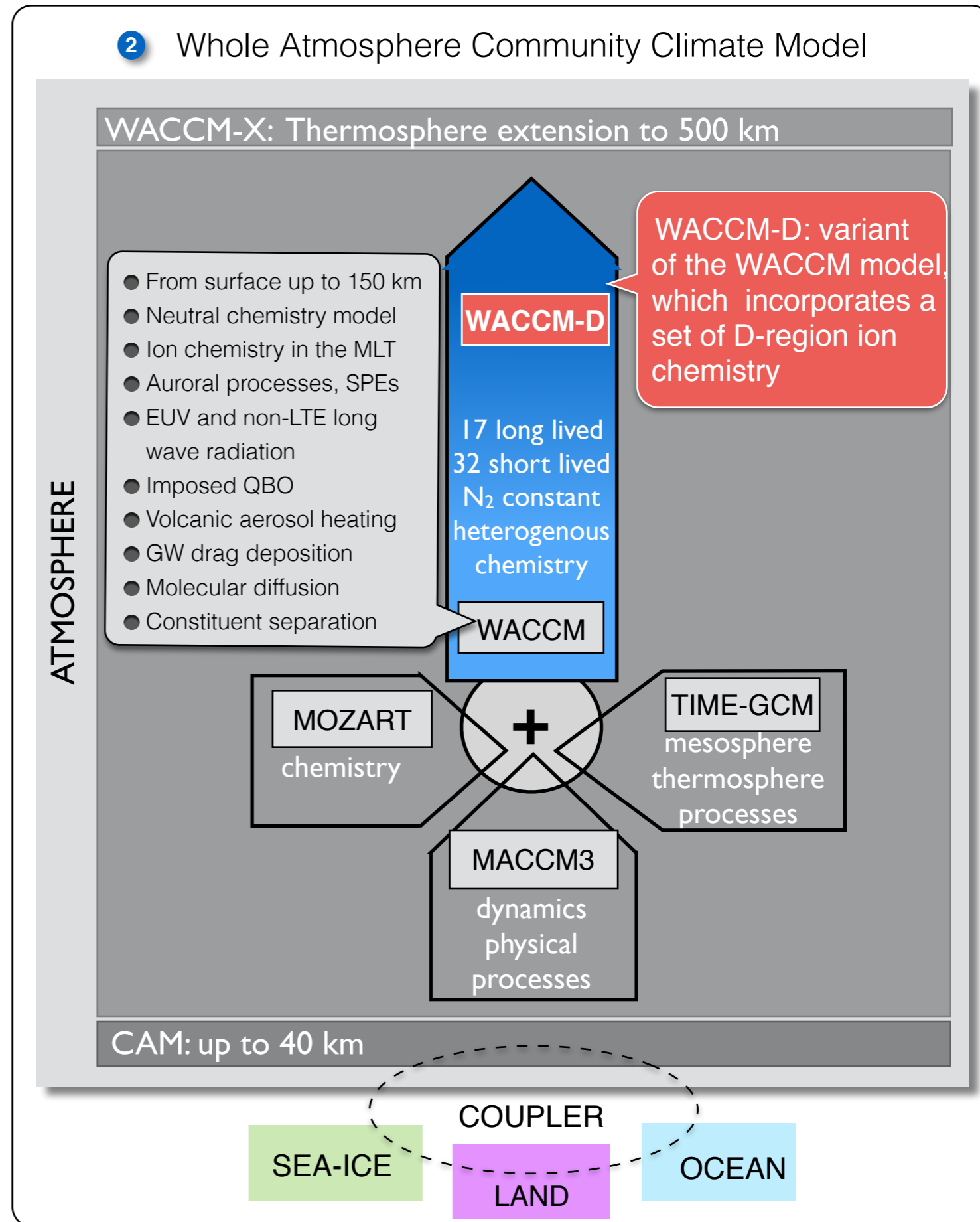
ECMWF (1958 - 2009) ERA-40 re-analysis

However:

1. The ion-neutral connection in the middle atmosphere is missing from the climate models today and the observed EPP effects in some important species such as HNO₃ cannot be reproduced using the commonly used parameterisation of HO_x and NO_x production
2. EEP ionisation is not included in the models

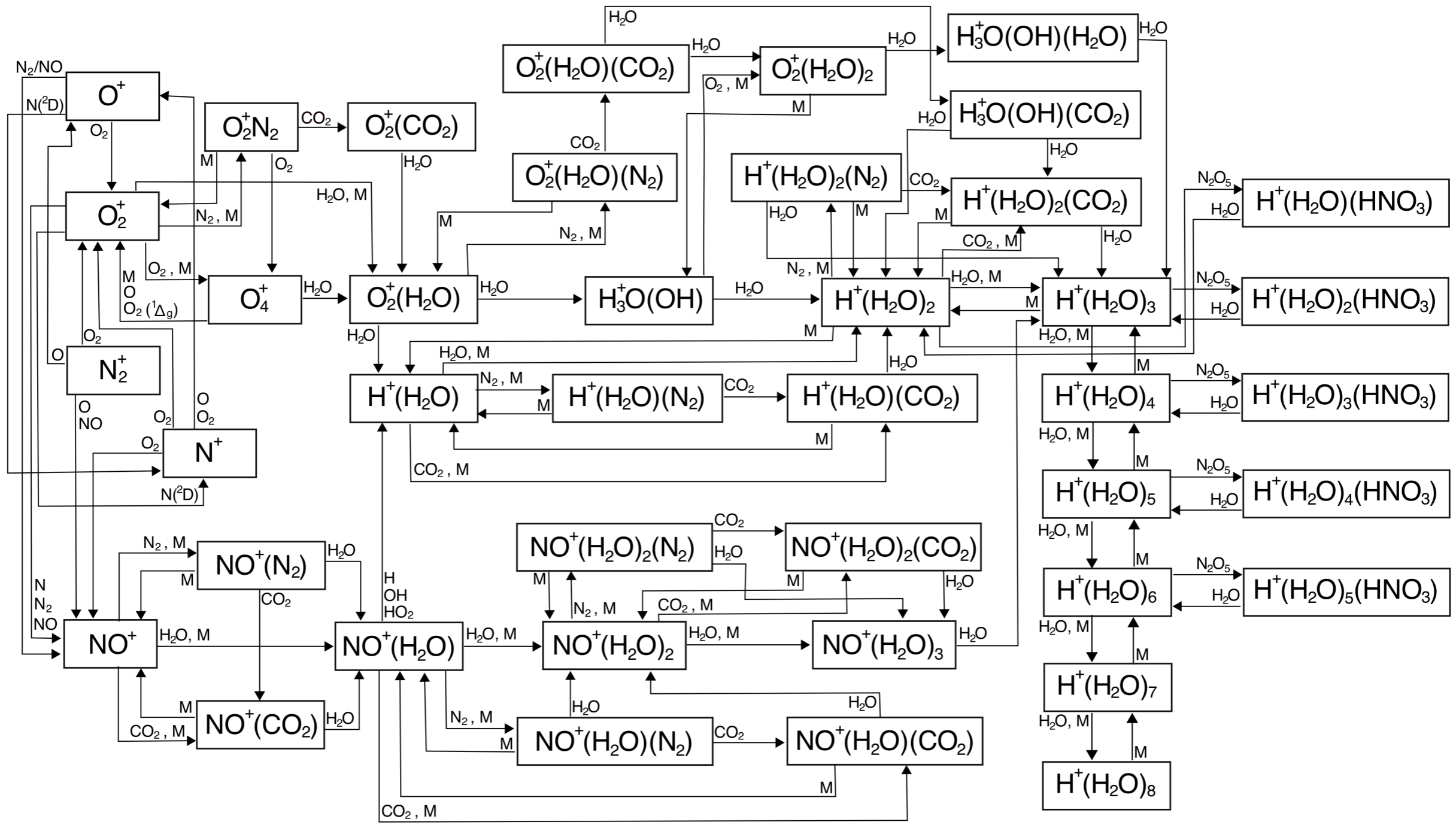
WACCM-D

- incorporates D-region ion chemistry with the aim to produce the observed EPP effects
- parameterised HO_x and NO_x production replaced by initial production rates of ions and neutrals due to particle impact
- reaction rate coefficients for the WACCM-D ion chemistry from SIC model, original WACCM 6-constituent ion chemistry model for the E-region retained
- ion reaction schemes based on SIC model, significant number of ions and reactions have been excluded



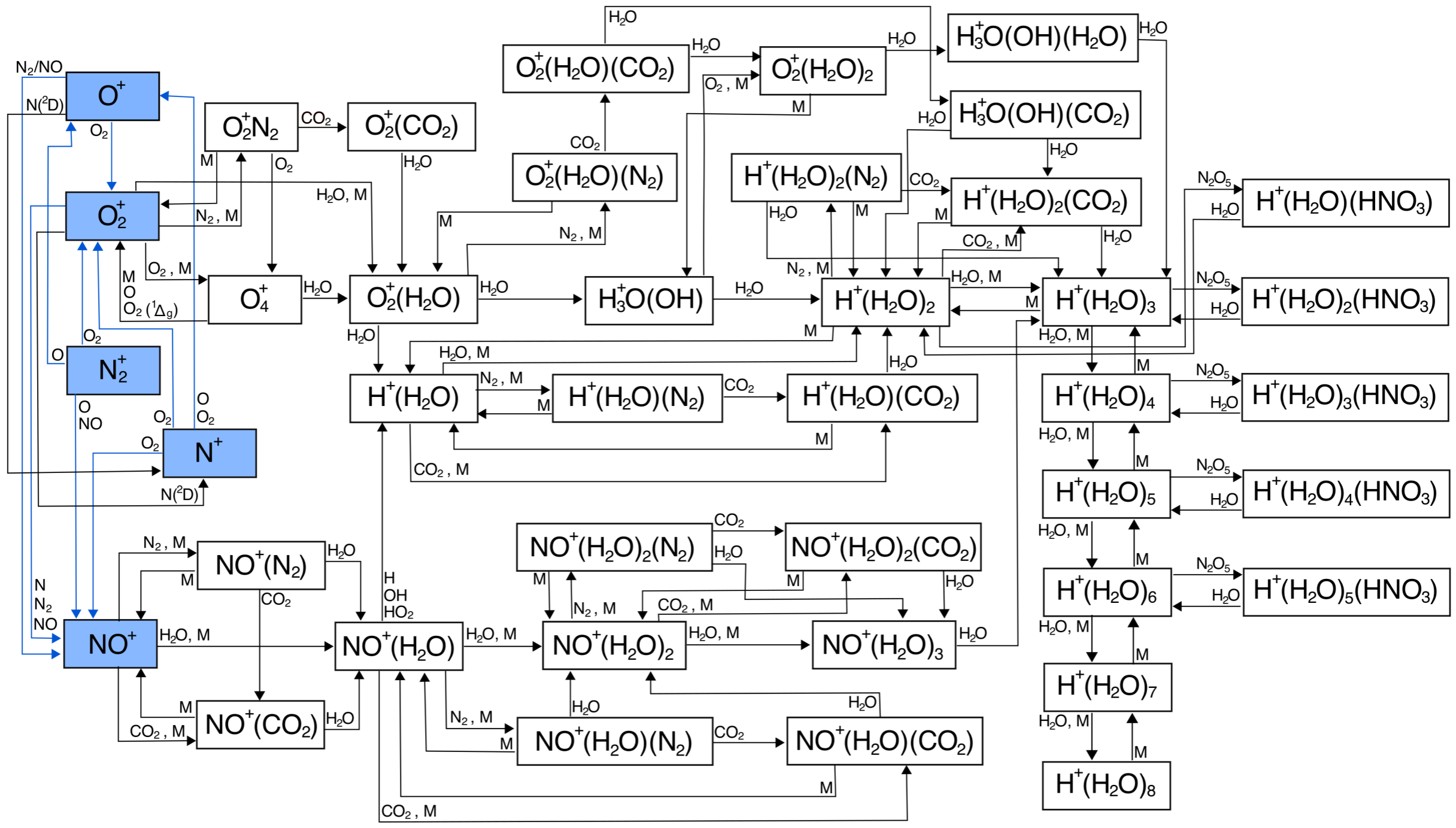
Positive ions

□ SIC

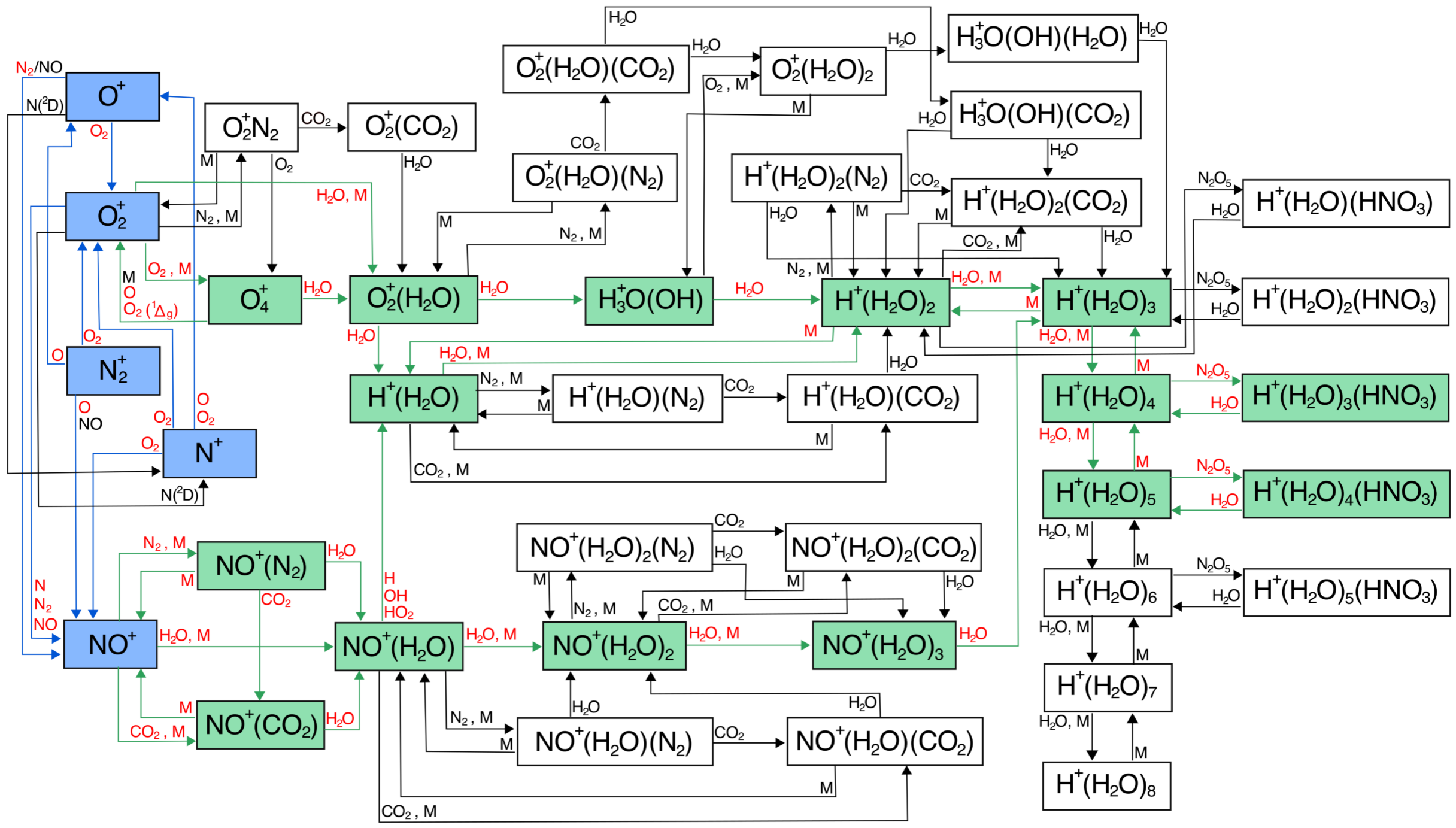


Positive ions

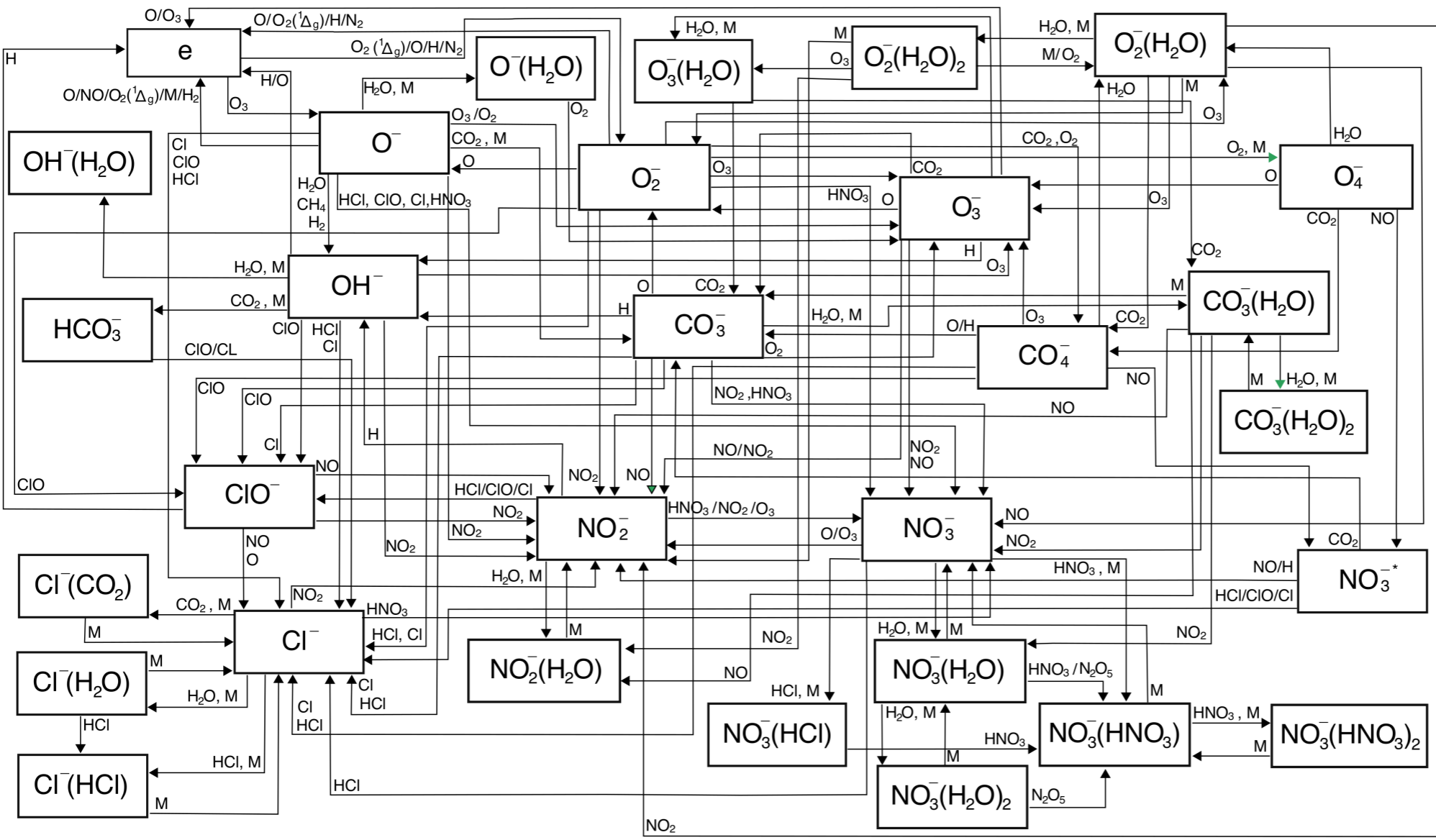
SIC WACCM



Positive ions

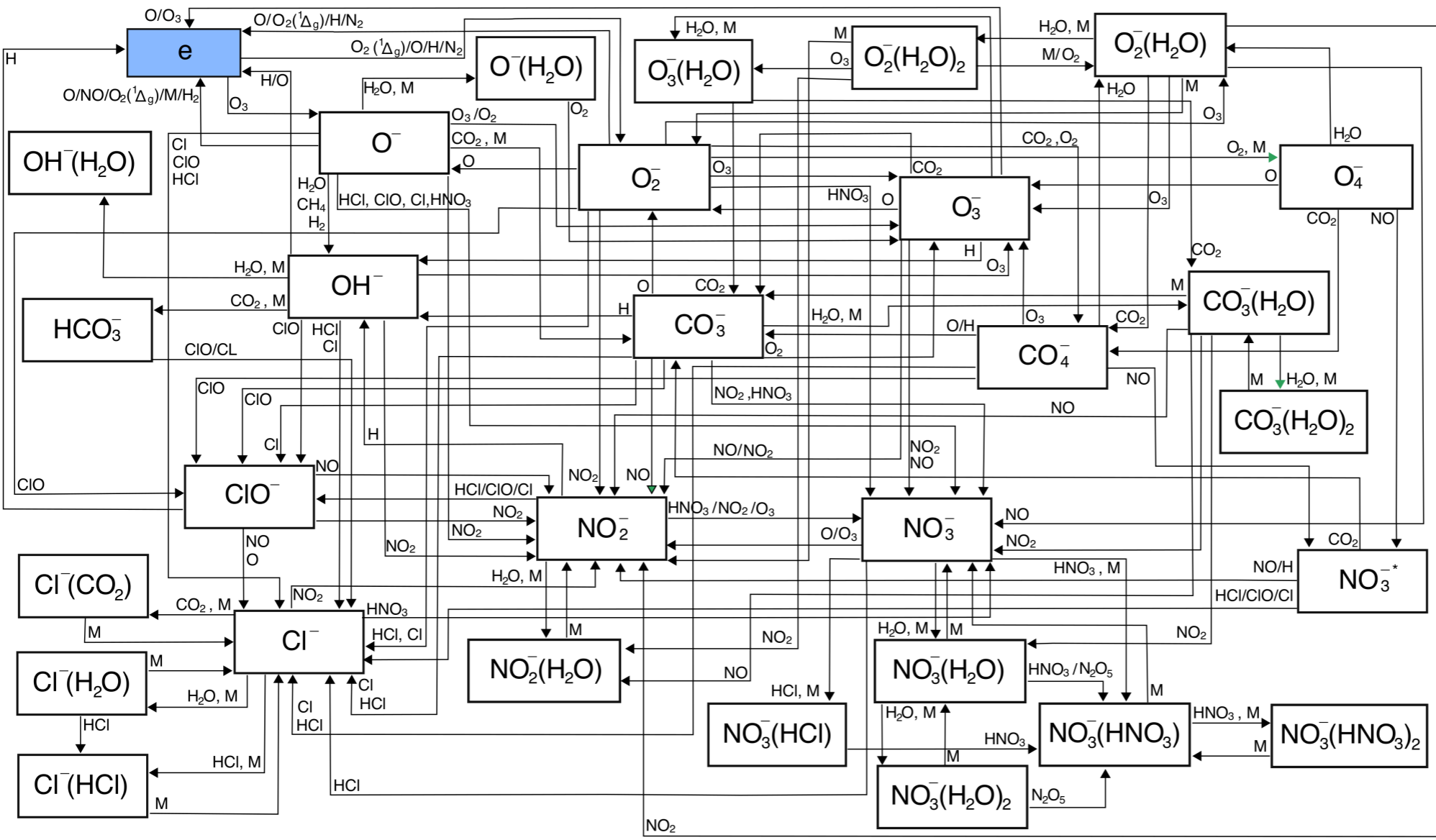


Negative ions



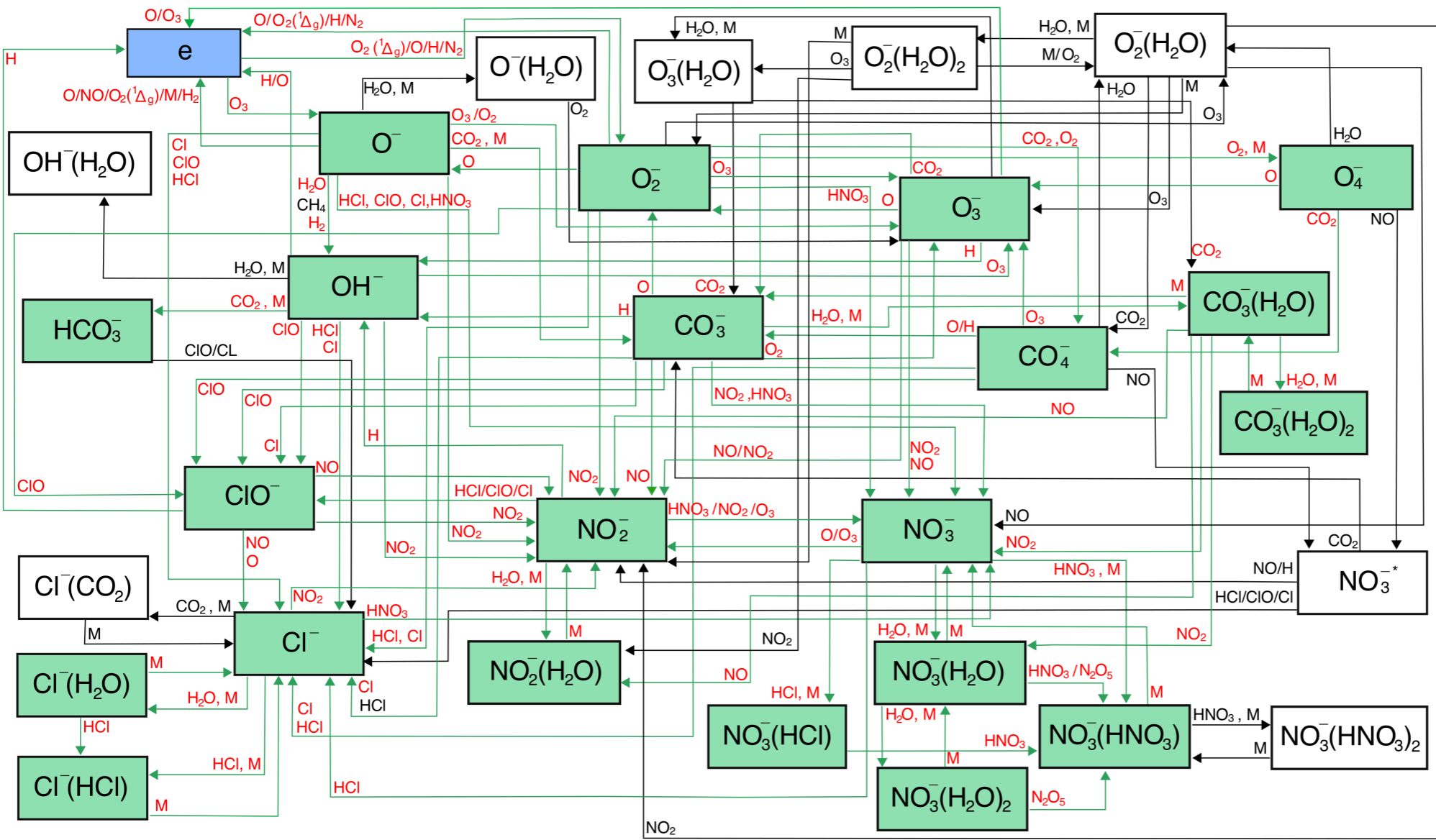
□ SIC

Negative ions

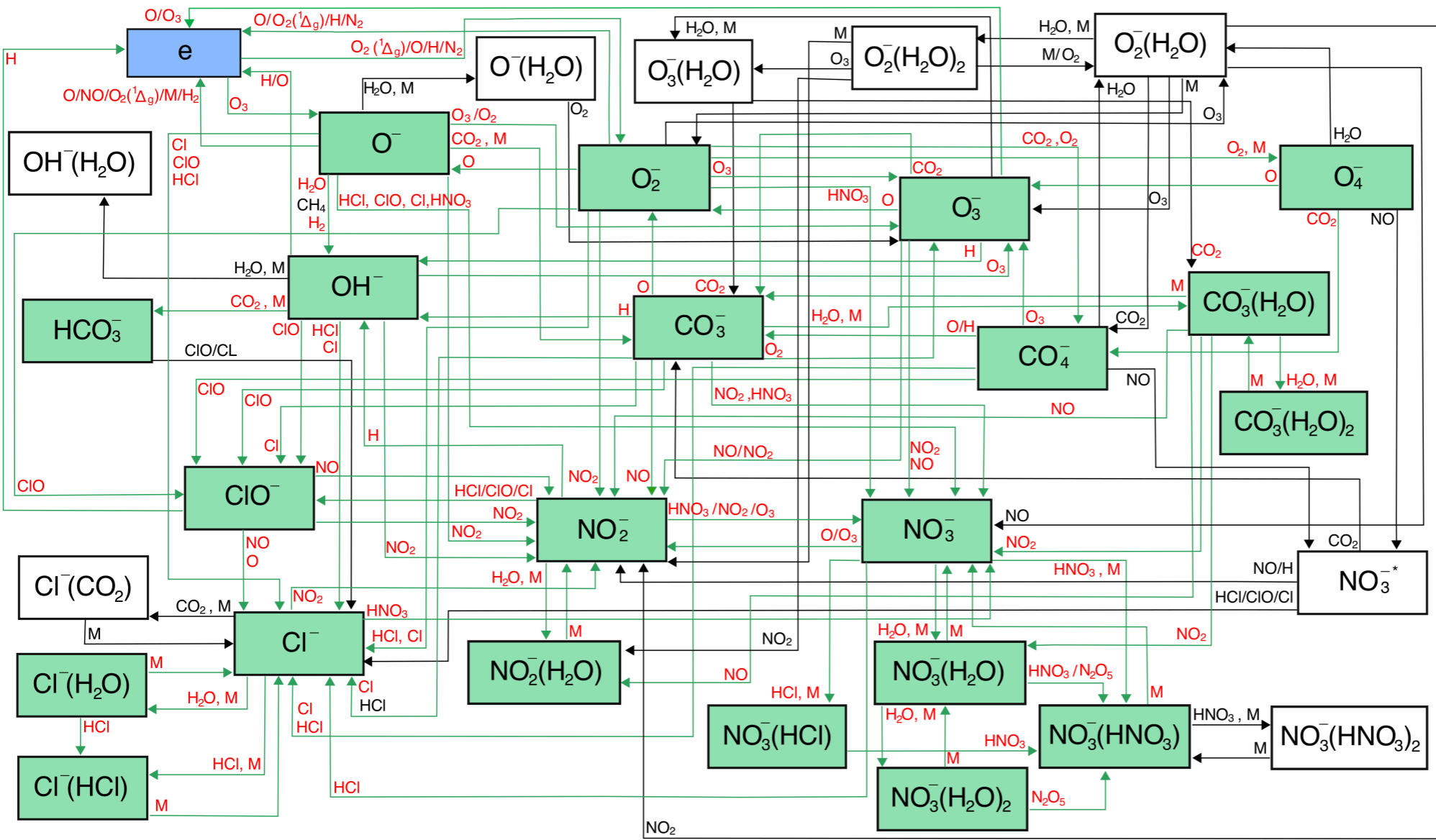


□ SIC
■ WACCM

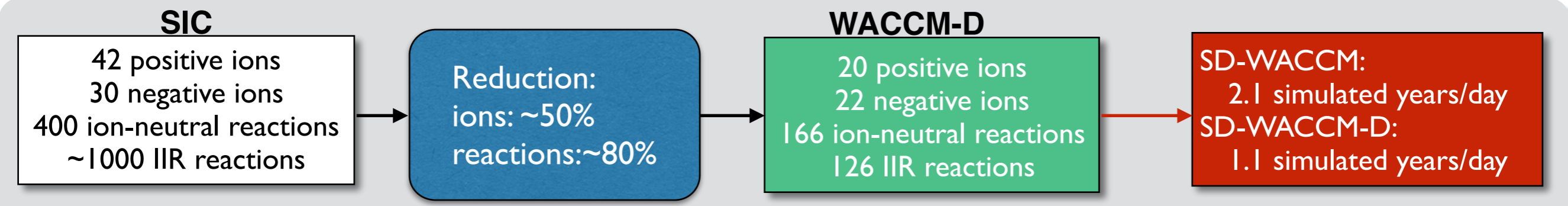
Negative ions



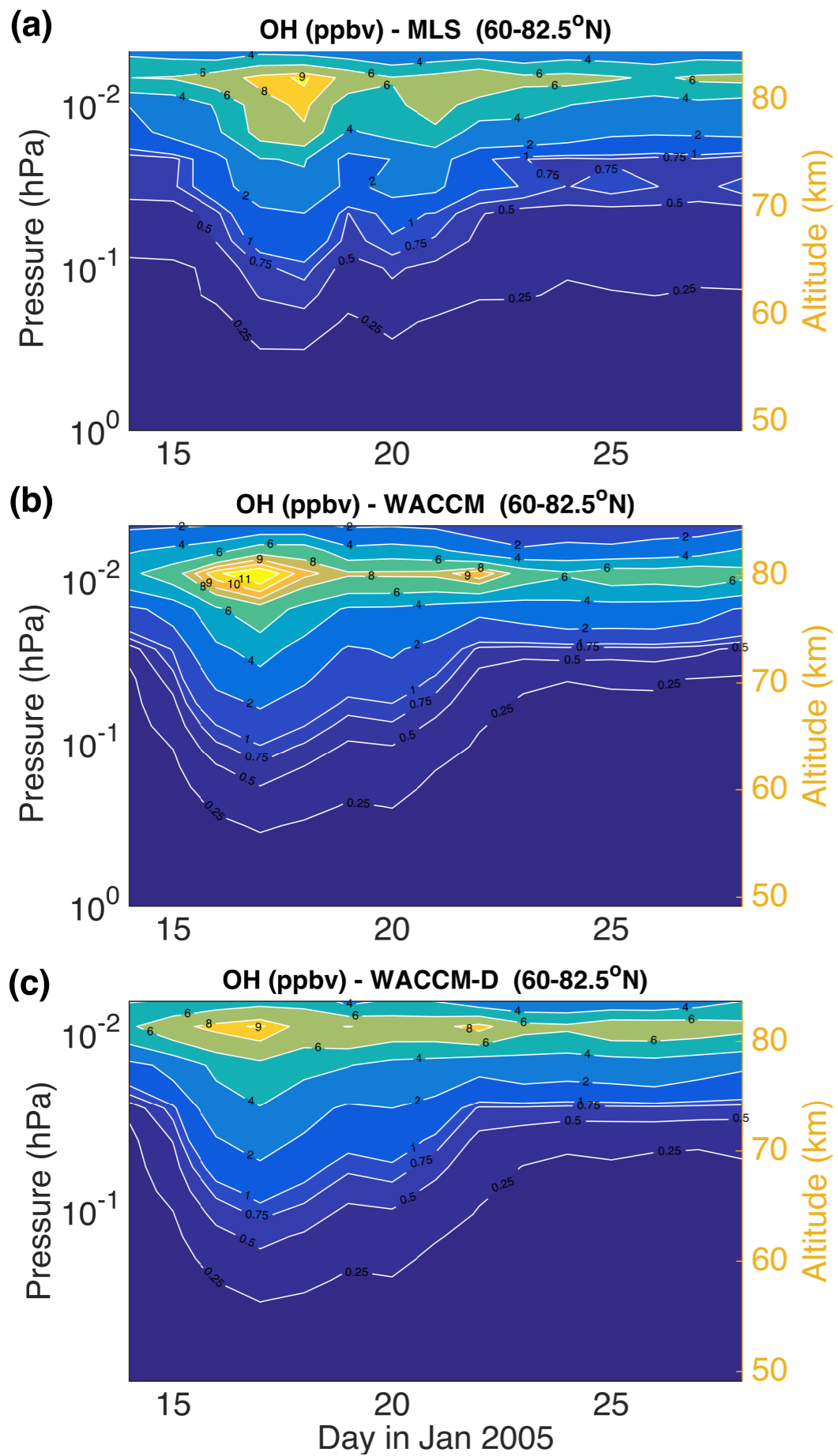
Negative ions



- SIC
- WACCM
- WACCM-D

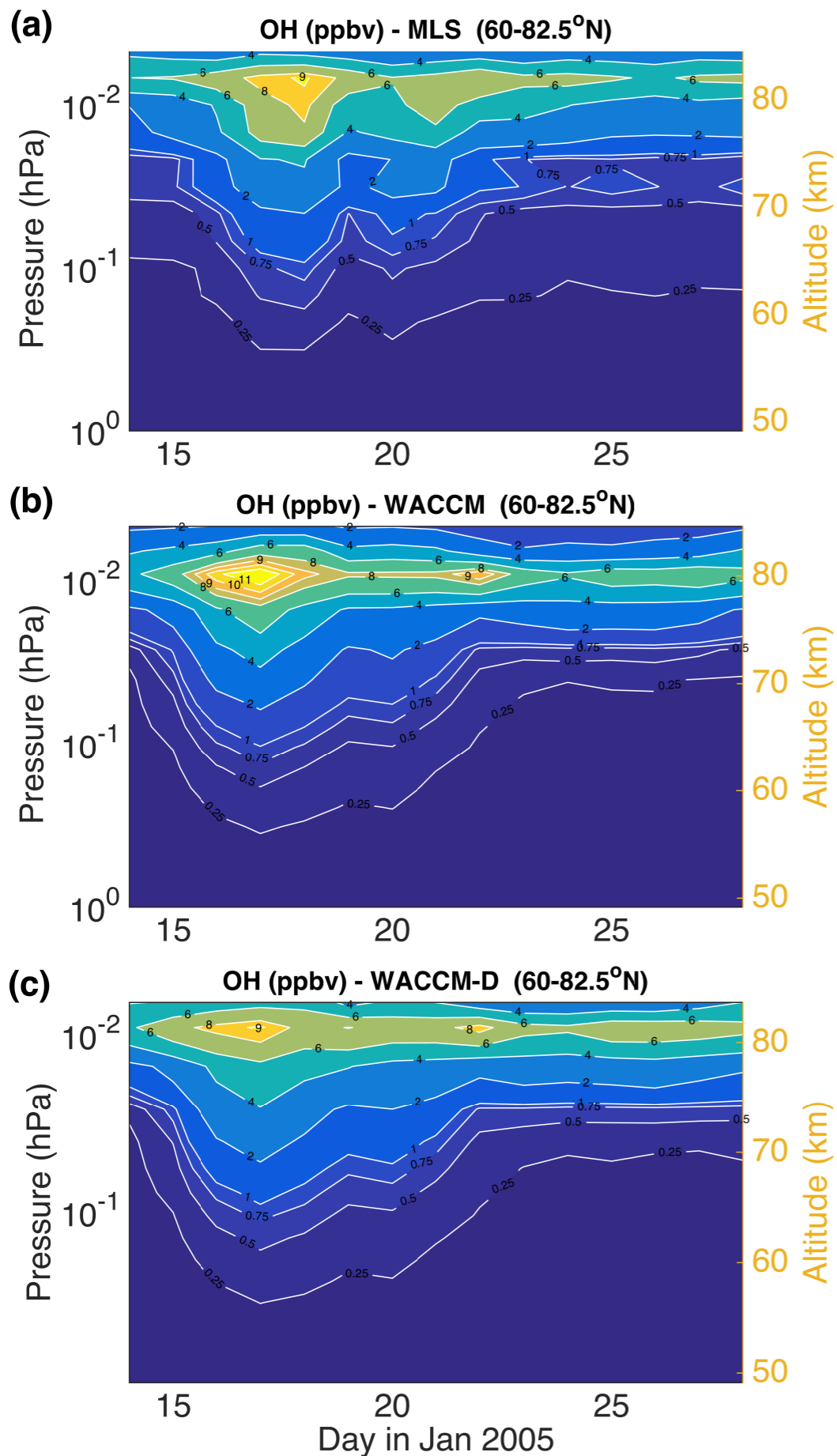


WACCM-D vs MLS: OH



2 model runs: SD WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations.

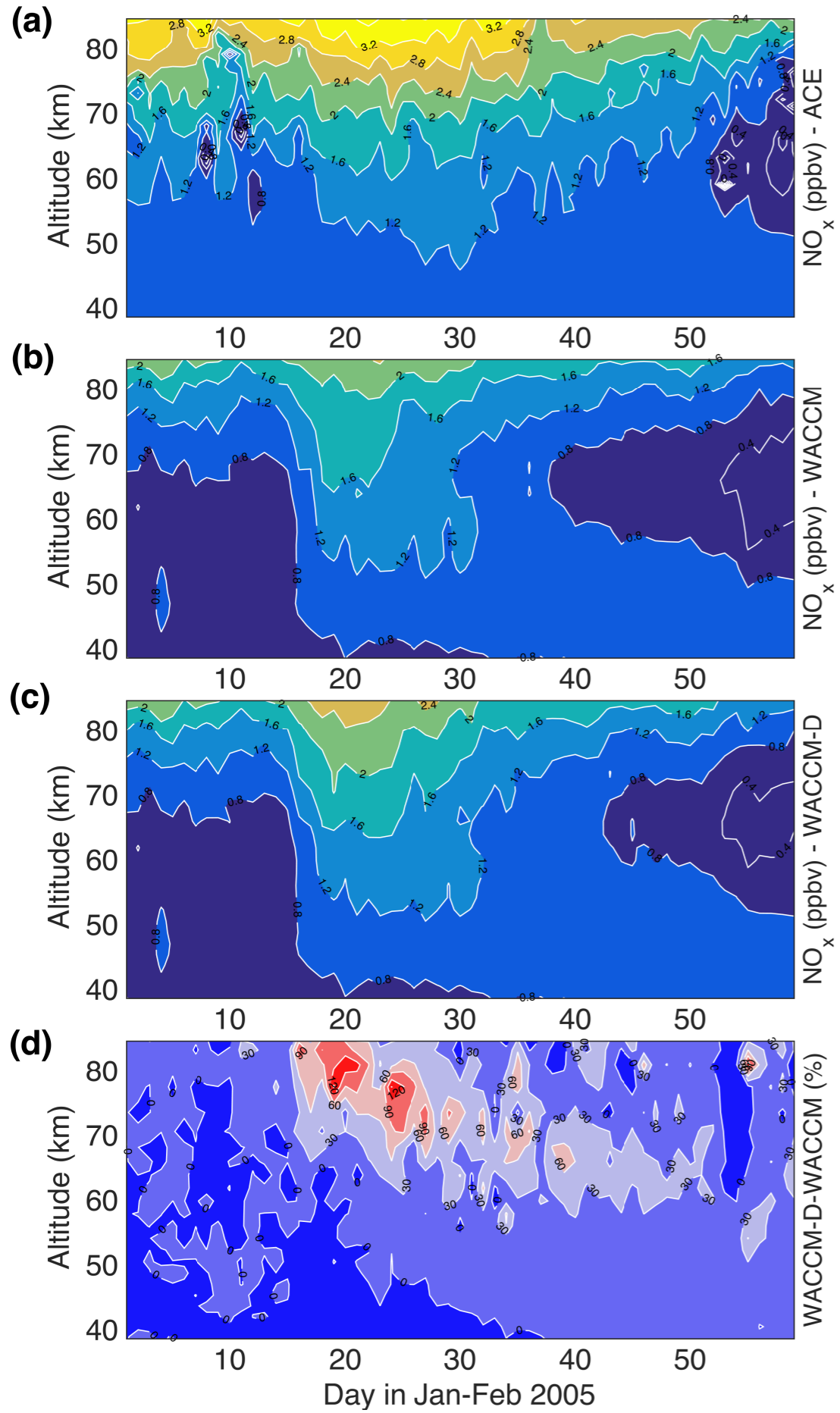
WACCM-D vs MLS: OH



2 model runs: SD WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations.

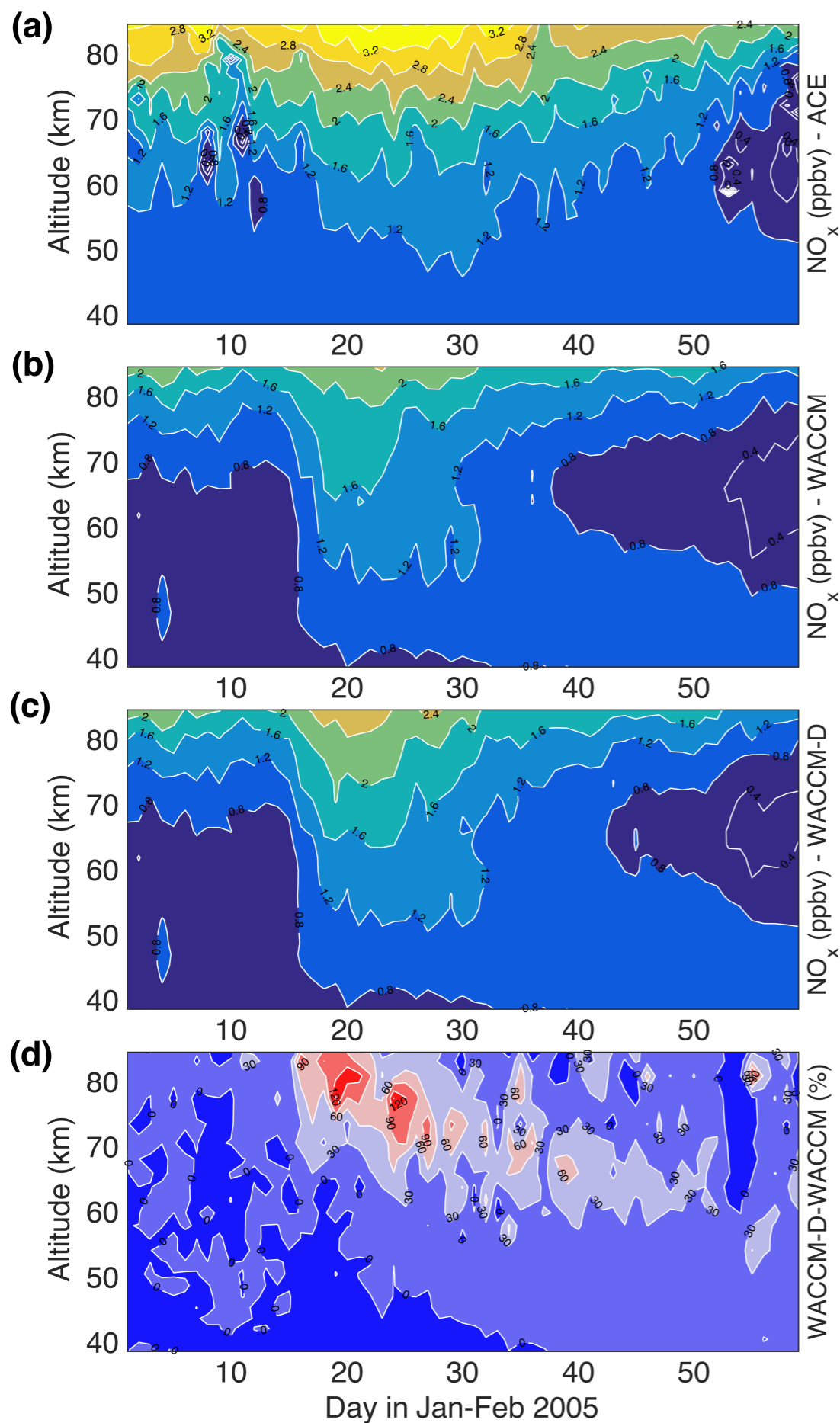
- significant perturbation observed in HO_x as a consequence of SPE
- MLS data and model predictions, show large OH enhancement during the SPE at altitudes between 58–84 km
- the largest observed and modelled increase occurred on 17–18 January at altitudes 80–82 km
- WACCM predictions overestimate OH values by 25-50% at altitudes between 70-80 km
- WACCM-D OH enhancement are in better agreement with observations than predictions from WACCM.

WACCM-D vs ACE-FTS: $\text{NO}_x = \text{NO} + \text{NO}_2$



2 model runs: SD WACCM and SD-WACCM-D for 1 January - 28 February 2005. WACCM output at ACE-FTS times and locations.

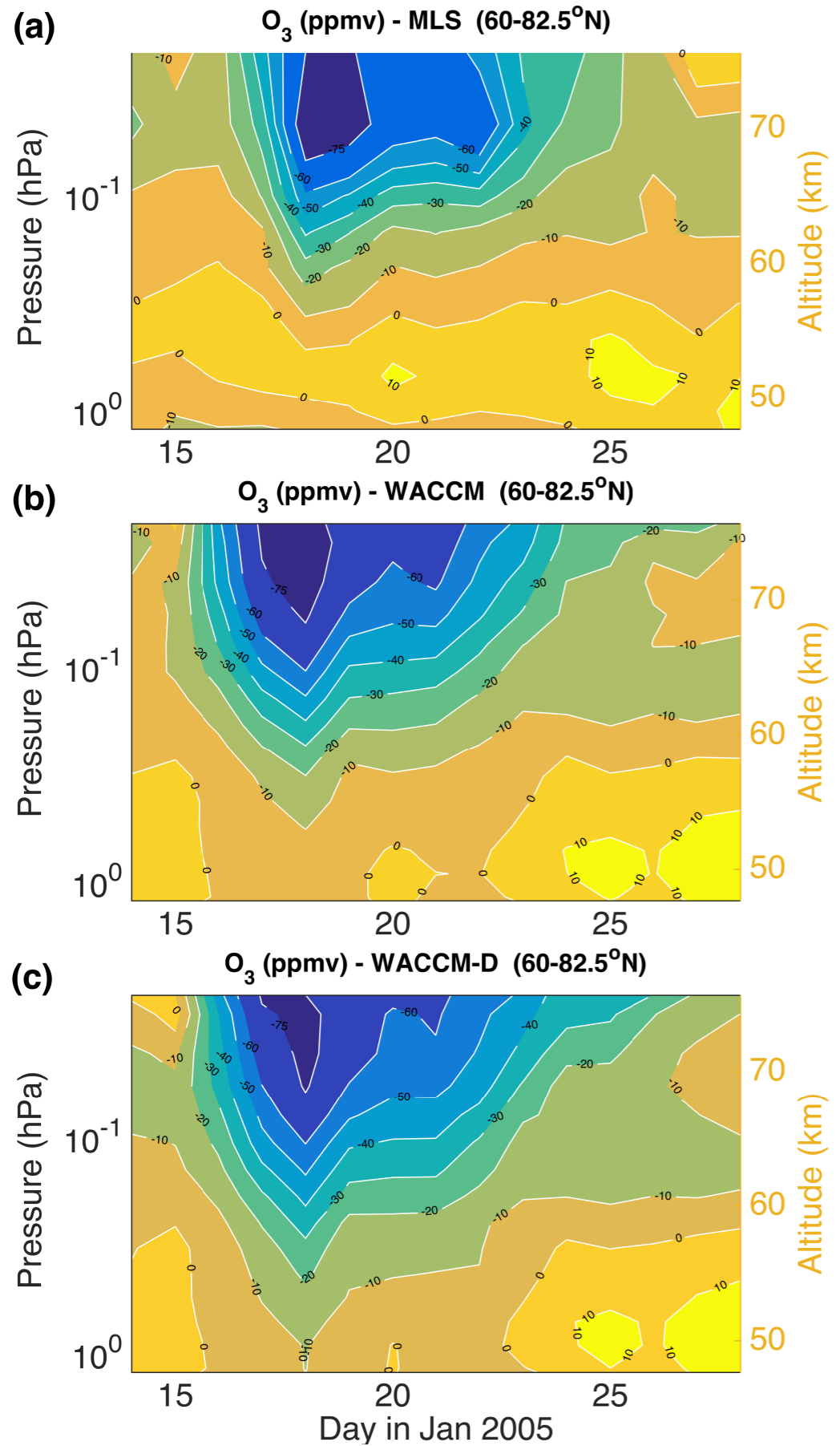
WACCM-D vs ACE-FTS: $\text{NO}_x = \text{NO} + \text{NO}_2$

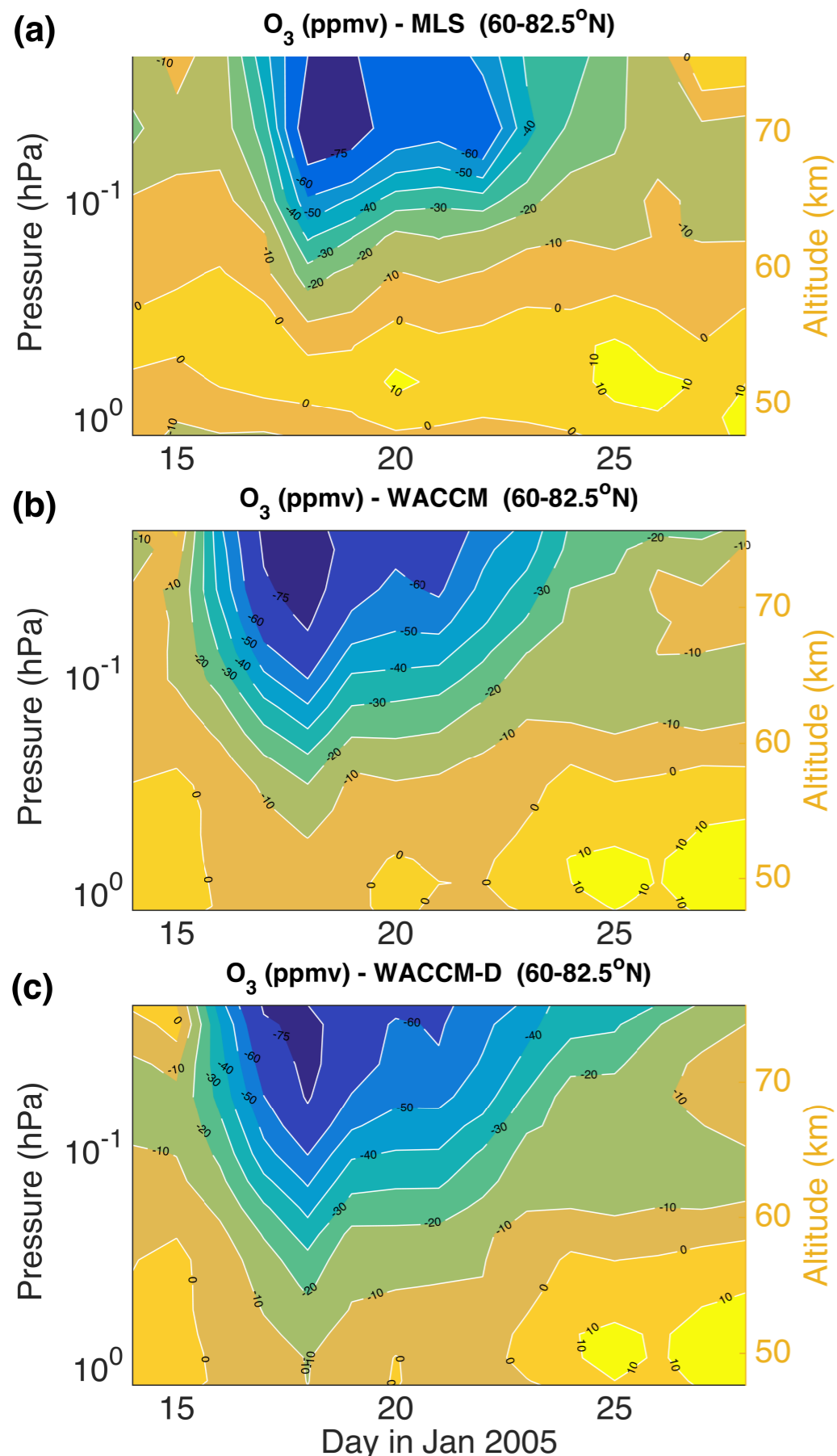


2 model runs: SD WACCM and SD-WACCM-D for 1 January - 28 February 2005. WACCM output at ACE-FTS times and locations.

- during the whole January, large amounts of NO_x were observed with rather moderate response during and after SPE impact between 16-31 of January
- predicted NO_x shows distinctive peak which corresponds to the SPE event
- during the whole considered period NO_x values predicted by WACCM and WACCM-D are significantly smaller than ACE-FTS measurements
- WACCM-D improvement of NO_x of about 60–130% between day 16 and 40 is still not enough to reach the observed levels

2 model runs: SD WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations.

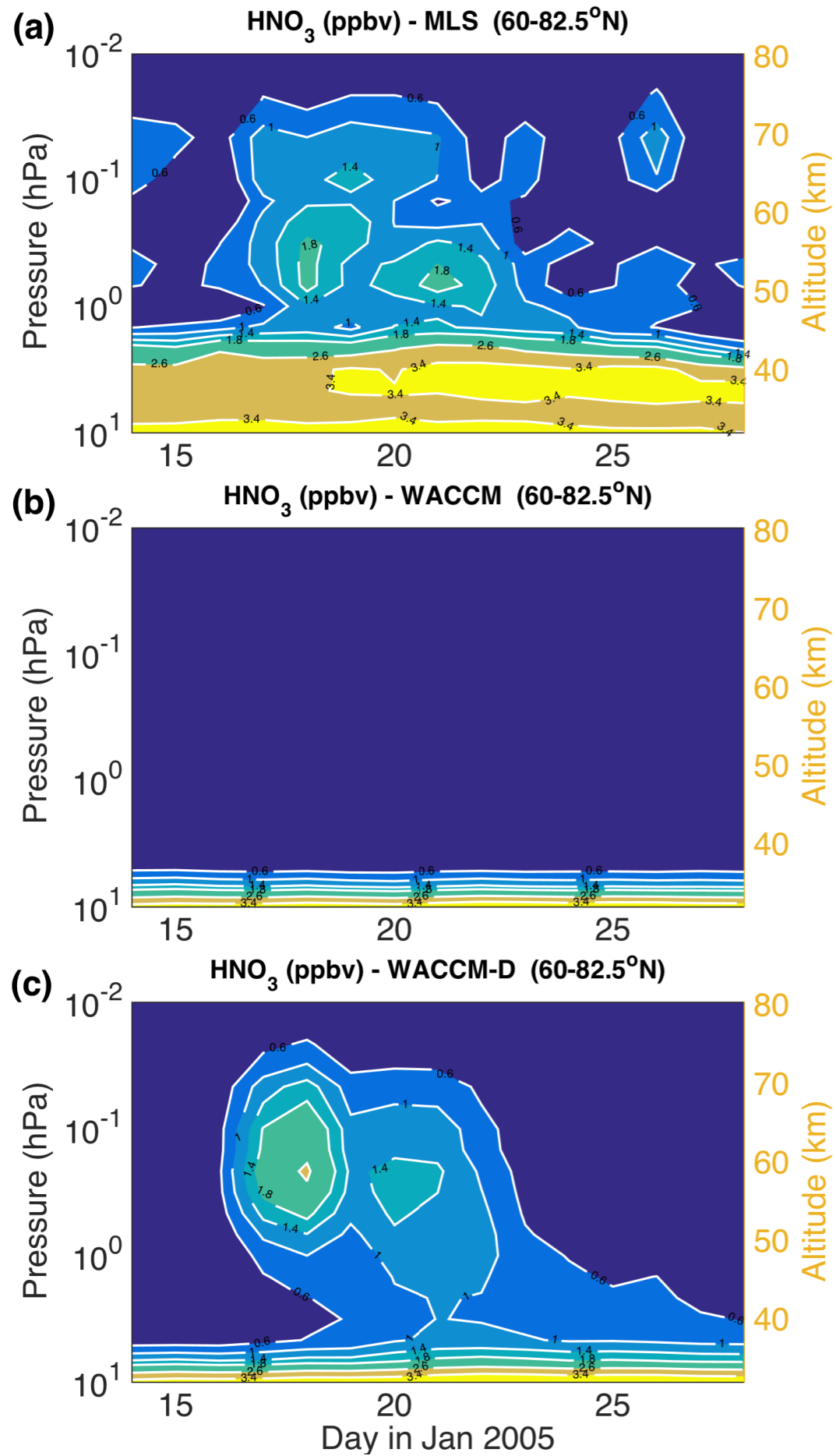




2 model runs: SD WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations.

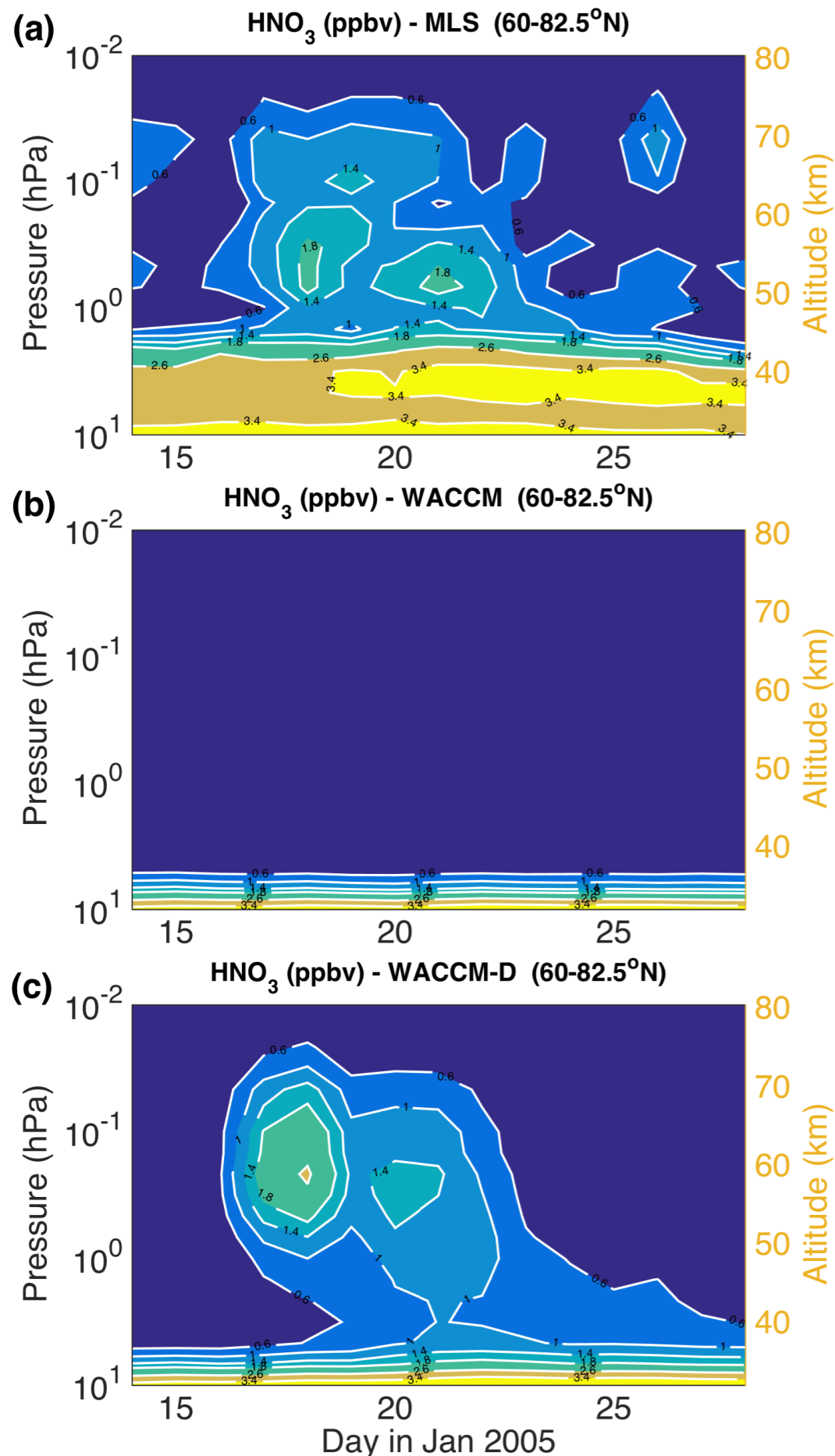
- during the SPE, ozone decreases up to about 80% are observed by MLS and predicted by WACCM and WACCM-D
- WACCM and WACCM-D are in reasonably good agreement with the observations considering the magnitude of the ozone changes
- The short-term ozone response does not change much but there is evidence of longer-term impact of ion chemistry on NO_x and ozone which can not be modelled with the ion chemistry parametrisation used in WACCM

WACCM-D vs MLS: HNO_3



2 model runs: SD-WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations. Aura MLS averaging kernels have been applied to WACCM output.

WACCM-D vs MLS: HNO_3



2 model runs: SD-WACCM and SD-WACCM-D for 1-31 January 2005. WACCM output at Aura MLS times and locations. Aura MLS averaging kernels have been applied to WACCM output.

- during the SPE, significant enhancement of MLS HNO_3 at altitudes between about 40–75 km
- elevated HNO_3 values (0.6–1.8 ppbv) lasted for about 10 days with maximum increase on January 17 and 21
- predicted HNO_3 values by WACCM were ~100 times lower (0.03–0.04 ppbv) than seen by MLS observations
- WACCM-D reproduce observed HNO_3 with a very good agreement:
 - the magnitude of the observed HNO_3 changes (0.6– 2.5 ppbv)
 - duration of the enhancement (~10 days)
 - affected altitudes (40–75 km)

APEEP

- precipitation model for radiation belt electrons created by van de Kamp et al., [2015]
- Ap index as the driving input parameter, defining the level of magnetospheric disturbance
- reconstructed Ap record used to create an electron precipitation time series for the whole CMIP6 period
- atmospheric ionisation data set calculated based on an Ap-dependent electron flux model

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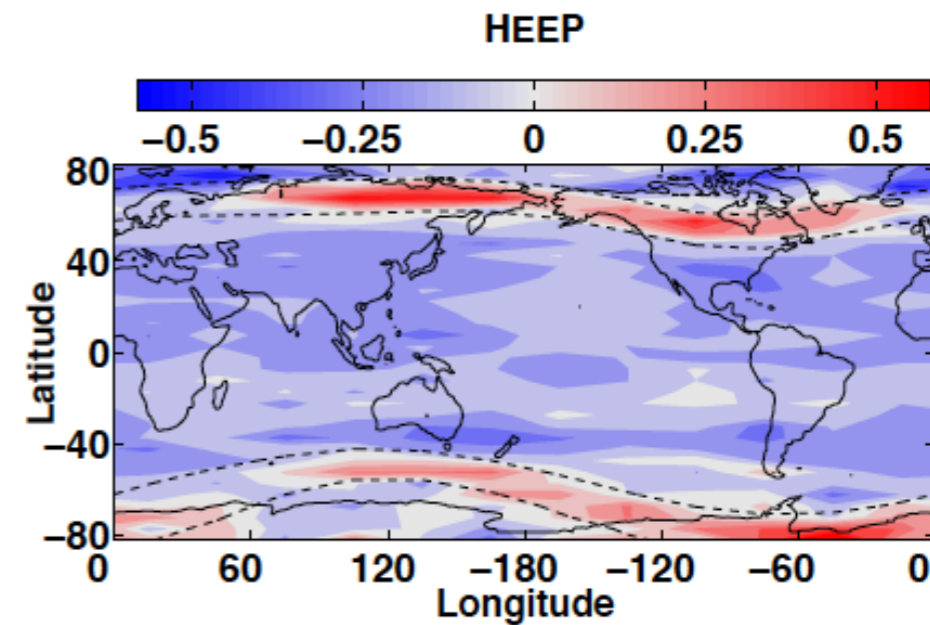
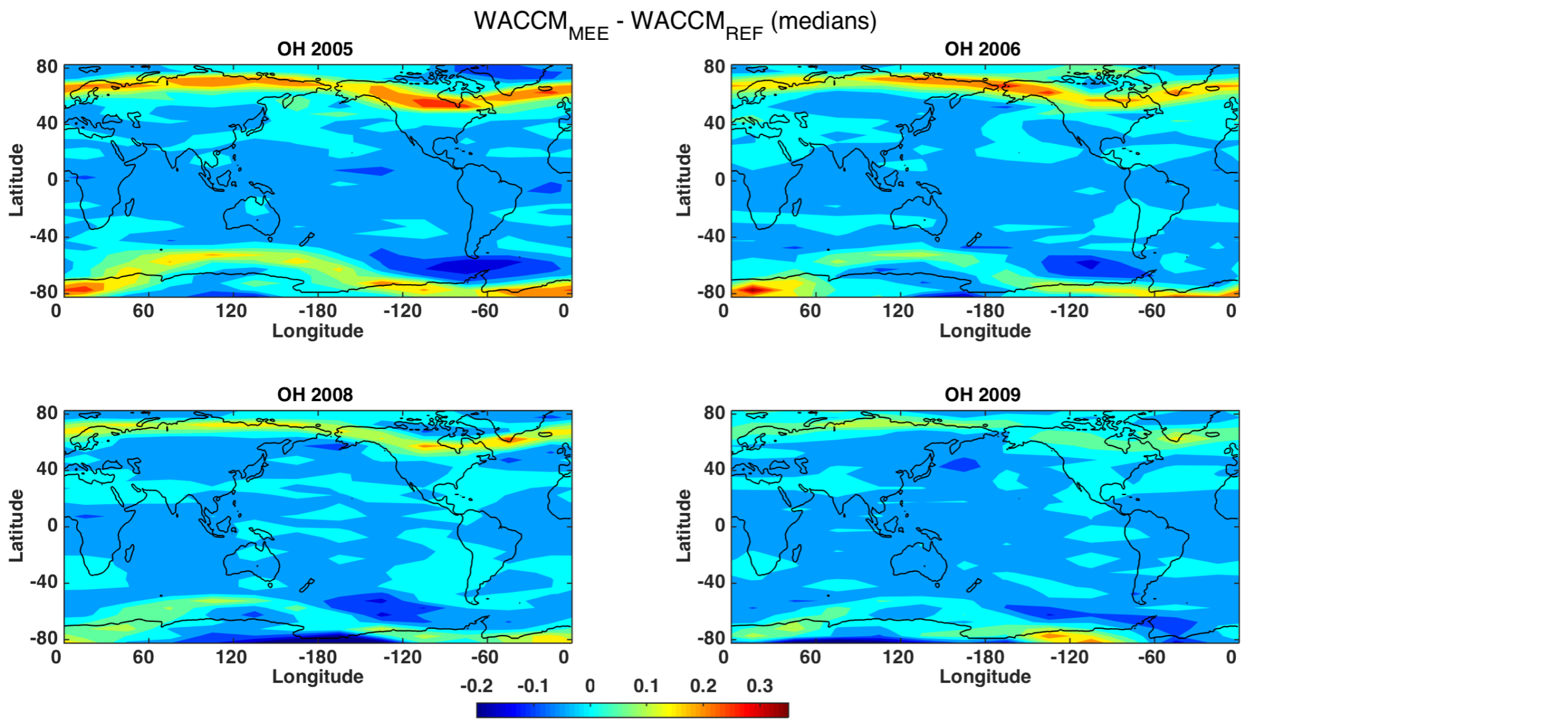
WEEP_CMIP5

compare APEEP output against original CMIP5 runs without EEP in order to show the significance of APEEP

WEEP_DYN

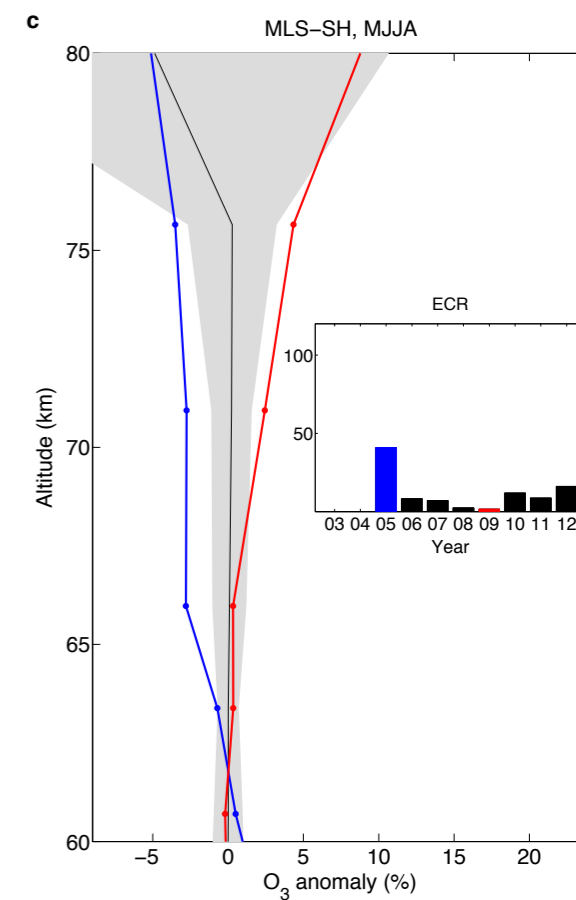
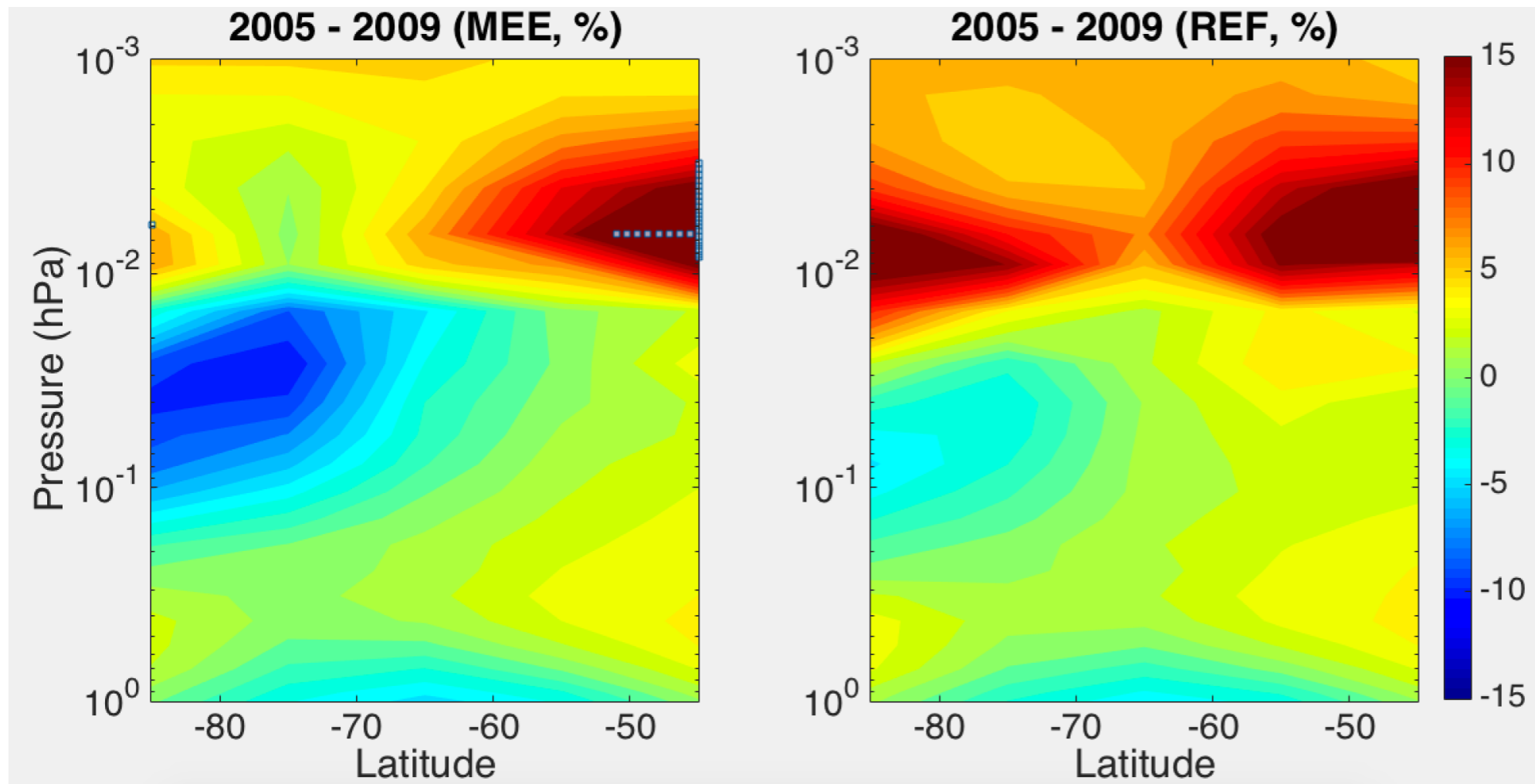
compare high and low years of EEP to determine sources and mechanism of dynamical changes in the atmosphere

Proof of concept - OH example



Andersson et al. 2014, ACP

Ozone relative changes % (MJJA, SH)



Andersson et al. 2014, Nature Comm.

- **WACCM-D** incorporates D-region ion chemistry to reproduce the neutral atmospheric effects caused by EPP in the polar region
- Including ion chemistry scheme in the model significantly improved the response of important neutral species to the energetic particle precipitation
- Including EEP source of ionisation will give us a complete picture of the importance of ion chemistry in the middle atmosphere in order to improve our understanding of solar influence on the atmosphere and climate.