WACCM Studies at CU/LASP

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Jeff France, WACCM Working Group Meeting, 9 Feb 2016

Outline

Ongoing EPP work using 300 years of FR-WACCM - Randall, Peck, France

The mesospheric polar vortices in FR-, SD-WACCM vs. Reanalyses - Harvey

CU/LASP assisting with the "Sathist" capability - Harvey, Pettit

RAISE - Response of the Atmosphere to Impulsive Solar Events

Simulation Specifications

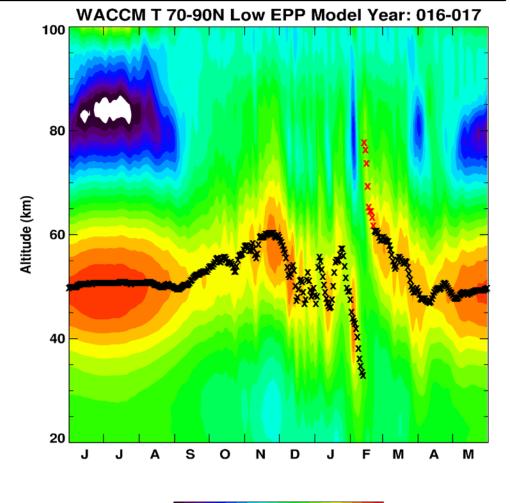
- CESM 1.2.2
- BWCN Compset
- 1.9° x 2.5° Horizontal Grid

Name	Run Length (spinup)	Solar Flux (f10.7)	Ap Index
Low EEP	300 years (10 years)	128	3
300 High EEP (10 years)		128	27

Elevated Stratopause Events in WACCM

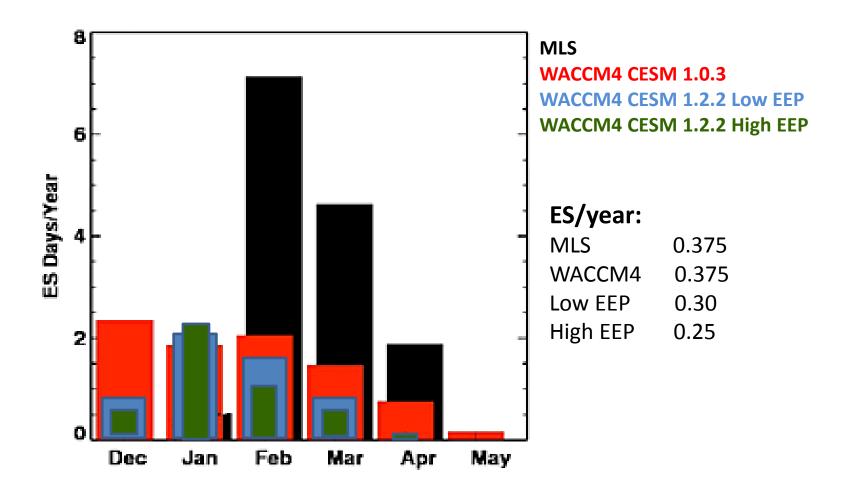
Method for identifying ES events based on *France and Harvey* [2013].

- Stratopause height at onset date must be at least 1 standard deviation above the climatological daily mean.
- 2. Must have "jump" of 25 km to onset height.



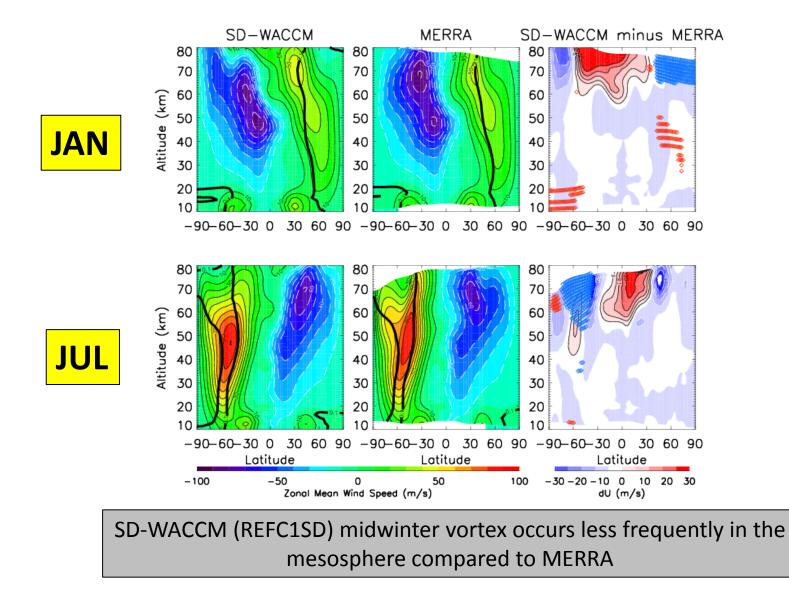


WACCM ES frequency



France and Harvey [2013]

Largest Differences in **Ū** are in the Summer Mesosphere



Updated Instruments for "Sathist" Output Option

- WACCM is output at observation locations and times.
- Original functionality created by Matthias Brakebusch. Current archives and scripts maintained by Chuck Bardeen and CU/LASP.
- >38 observation sources from 1978-Present
- Create files containing geo-location and profile information: DATE, TIME, LATITUDE, LONGITUDE, ORBIT_NUMBER, PROFILE_NUMBER
- Ongoing observations need updating: ACE, SOFIE, EOS-MLS, SABER, SBUV/2
- Update with newest data versions: MLS v4
- Add additional observations: SNOE

Response of the Atmosphere to Impulsive Solar Events (RAISE)

- Four main science questions:
- (1) How well do models simulate effects of recent ISEs?
- (2) What factors control the atmospheric response to ISEs?
- (3) What is the range and sensitivity of the atmospheric response?
- (4) Are there long-term, cumulative effects of ISEs on the atmosphere and climate, and with what certainty can these effects be modeled?
- We will use WACCM, SD-WACCM, and WACCM-X to simulate recent ISEs (CMEs, SPEs, flares) and quantify the atmospheric response.

Josh Pettit is the LASP point-of-contact for WACCM simulations in support of RAISE

He is using SD-WACCM to answer the following 3 questions:

- 1. What is the medium/high energy electron influence on middle atmosphere chemistry?
- 2. Can WACCM's middle atmosphere chemistry be improved by implementing medium and high energy electrons?

	ISE Time Period	Reasons to Study	
	14-16 July 2000	3 rd largest solar proton event (SPE) in past 50	
		years, X5 flare. Southern Hemisphere winter.	
	19 Oct to 5 Nov 2003	4th largest SPE in past 50 years. X17 flare plus	
		overall high flare activity. Many atmos. obs.	
		Early Northern Hemisphere (NH) winter.	
	15-17 Jan 2005	13th largest SPE in past 50 years and hard	
		spectrum. X2 flare. Many atmospheric	
		observations. NH winter.	
	7-11 Sep 2005	Moderate SPE. X17 flare. ISR electron density	
		data. Many atmospheric observations.	
У	15 Feb to 9 Mar 2011	X1,X2 and many M-class flares. Very small (50	
-		pfu) SPE. SDO data.	
	4-9 Aug 2011	X6 flare. Very small (<100 pfu) SPEs. SDO data.	
	7-13 Mar 2012	10th largest SPE in past 50 years. X5 flare. SDO	
		data.	

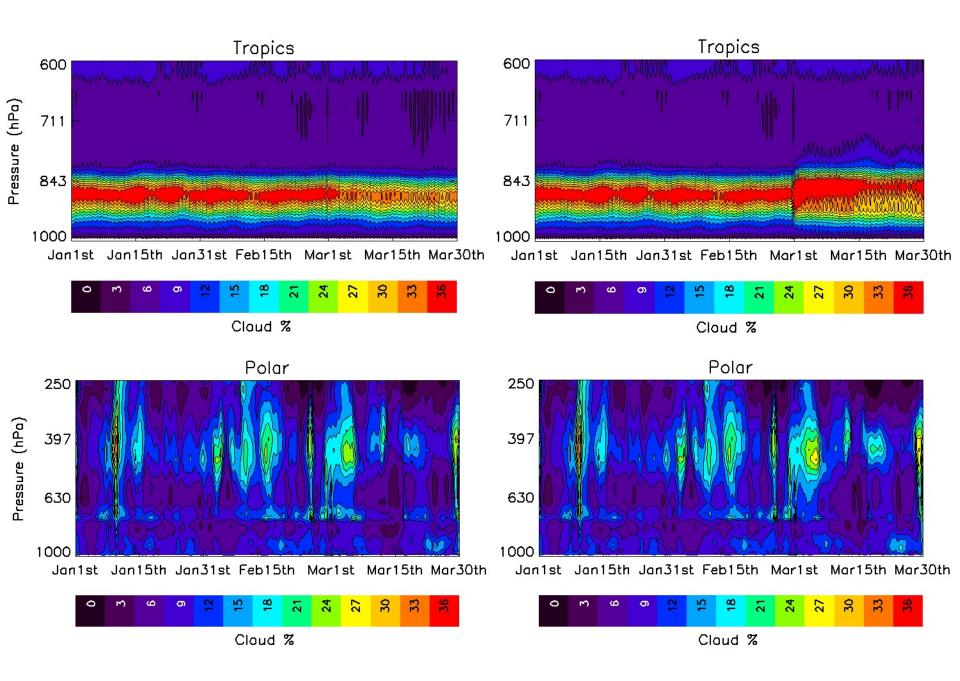
- 3. What electron data set would be the best for the implementation into WACCM? (contamination issues, etc.)
- **PhD Dissertation**
- Medium and higher energy electrons effects on the atmosphere in WACCM

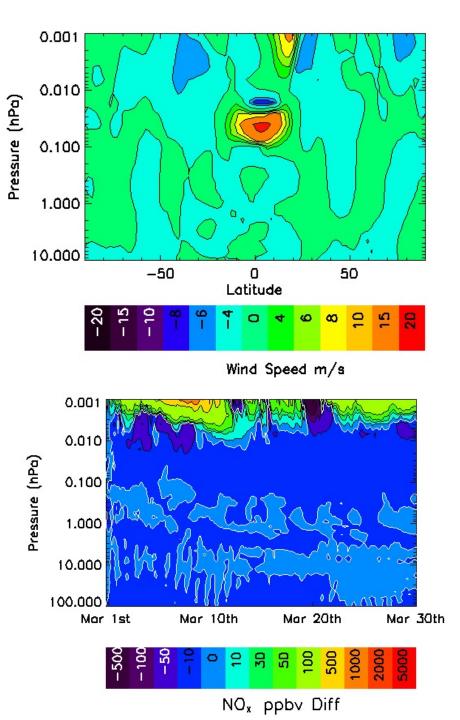
Time-Step simulations for RAISE

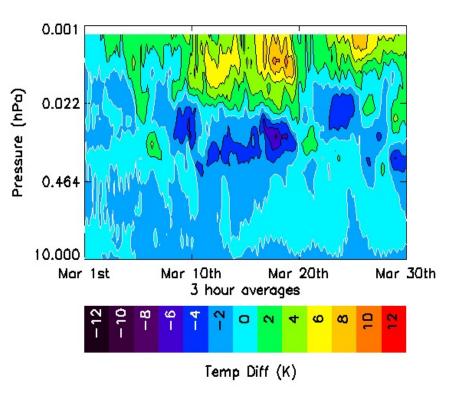
Goal: Test the sensitivity of WACCM by decreasing the time-step

Purpose:

- 1) More accurately capture solar flare effects
- 2) Potential use in WACCM-X







• Currently working on code modifications to 'super-cycle' the clouds at both 5 minute time-steps and 10 minute time-steps

Summary

Elevated Stratopause in 300 years of FR-WACCM - Peck and France

Polar vortices in WACCM and Reanalyses - Harvey

EPP – Randall

We will continue to keep the Sathist datasets and the SPE files up-to-date.

Stay tuned for more exciting results from Josh Pettit.

Thanks!

Energetic Particle Precipitation (EPP) Ionization & Dissociation NO_x and HO_x

NO_x and HO_x Destroy Ozone

