# Using TUV for inline photolysis in CESM

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# Current Photolysis Algorithm in CESM2

- Provides photolysis rates and chemical heating rates
- Uses 100 wavelength bands from 121-750 nm
- Two methods for actinic flux:
  - jshort (121-200 nm, 33 bands)
    - Calculates actinic flux as O<sub>2</sub> and O<sub>3</sub> transmission
  - jlong (201-750 nm, 67 bands)
    - Calculates actinic flux using a lookup table
- Calculates photolysis rates using a table of cross-sections and quantum yields compiled by Doug Kinnison
- Clouds: effective albedo (clouds below) and transmission factor (clouds above) based upon cloud fraction and cloud water content

## Actinic Flux Lookup Table

- Calculated off-line using TUV
  - 4-stream calculation
  - Fixed O<sub>3</sub> profile that scales with column amount
  - No aerosols, NO<sub>2</sub>, or SO<sub>2</sub>
- /glade/p/cesmdata/cseg/inputdata/atm/wacc m/phot/RSF\_GT200nm\_v3.0\_c080811.nc

Dimension	Count	Minimum Value	Maximum Value
Albedo	6	0.05	1.0
Column O <sub>3</sub>	20	30 DU	600 DU
Solar Zenith Angle	24	0°	97°
Altitude	151	0 km	150 km
Wavelength	67	201 nm	725 nm



# Tropospheric Ultraviolet and Visible (TUV) Radiation Model

- Calculates the following in the UV and visible:
  - actinic flux
  - photolysis rates
  - action spectra
- Has its own set of cross-sections and quantum yields
- Stand-alone version created by Sacha Madronich (Madronich & Flocke, 1997; Lee-Taylor & Madronich, 2007)
- Available for download: <u>https://www2.acom.ucar.edu/modeling/tropospheric-ultraviolet-and-visible-tuv-radiation-model</u>
- Ported to WRF by Stacy Walters
- Ported to CESM & MICM by Francis Vitt

# New CESM/TUV inline implementation

- Inline TUV currently only used for the actinic flux calculation
  - 2-stream calculation
  - Uses CESM wavelength grid (100 bands from 121-750 nm)
  - Uses actual gas profiles (O<sub>2</sub>, O<sub>3</sub>, NO<sub>2</sub>, and SO<sub>2</sub>)
  - Includes combined aerosol profile
    - optics interpolated from optical properties calculated for RRTMG
  - Clouds can be handled in TUV following Liao et al. (1999) or in chemistry as we do now
  - Provides output for actinic flux and some spectral integrals
- Uses cross-sections and quantum yields from CESM2 (Kinnison)
- Used for photolysis and heating rate calculations

### SD-WACCM: CESM2 & TUV give similar O<sub>3</sub>

Year 2011



But TUV up to 22 DU (~7%) higher in some regions

#### Using TUV Increases Ozone column



# Aerosols reduce tropical $O_3$ , increase polar $O_3$



Perhaps removing the fixed O<sub>3</sub> profile affects high latitude?

# Nuclear weapons ignite mass fires injecting smoke into upper troposphere/stratosphere

- Flash with thermal pulse
- Followed by blast wave
- Can ignite fires in large area and generate secondary ignition
- Firestorms may develop which modify local meteorology and inject smoke into the upper troposphere
- Area burned scales with weapon yield
- Hiroshima
  - 15 kt weapon
  - 13 km<sup>2</sup> burn area



Hiroshima after blast and fires

# WACCM 150 Tg Smoke: CESM vs. TUV US/Russia Global Nuclear War

 $O_3$  Column: TUV – Table

O<sub>3</sub> Column: TUV Table



# WACCM 150 Tg Smoke: Ozone and UV Index US/Russia Global Nuclear War

O<sub>3</sub> Column: TUV

max(UV Index): TUV



Can now get this from CESM, no postprocessing!





5-25% reductions are seen in calculated OH when inline TUV is included in WACCM calculations.

OH oxidizes  $SO_2$ .

Reduced OH leads to longer SO<sub>2</sub> lifetimes.



# Summary

- Inline calculations using TUV allow for more precise calculations of photolysis rates including the effects of actual profiles of O<sub>3</sub>, SO<sub>2</sub>, NO<sub>2</sub>, aerosols, and clouds.
- ~1.6% change in model throughput in WACCM
- Allows new output fields:
  - Action Spectra: UV-A, UV-B, UV-B\*, PAR, UV Index, max(UV Index)
  - Profiles of actinic flux: Lyman alpha (121 nm), Schuman Runge Bands (180 nm), O3 Hartley Band (210, 240, 308 nm), Chappuis Bands (400, 600 nm)

# Questions/Discussion

- Still using existing lookup table for cross-sections and quantum yields, will want to unify values from Kinnison, TUV (Madronich), and other sources (e.g. HARP, Hall)
- How do we want to handle cloud effects (cloud fraction, cloud overlap, ...)? We currently have 2 different approaches.
- TUV 4-stream (discrete ordinate) code will also be ported to WRF, CESM and MICM as an option. It will be more expensive (maybe 2.25x the cost of 2-stream), but perhaps not too expensive.
- Are there other photolysis reactions that people are concerned about that should be evaluated?
- Do we want any other output? Other action spectra?
- Rare computational errors resulting in high max(UV Index) values require further investigation.
- We are planning to port FastJ. Are there any other approaches people would like to see?