



Impacts of Solar Irradiance and Auroral EPP in WACCM

Presented by,
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WACCM Simulations

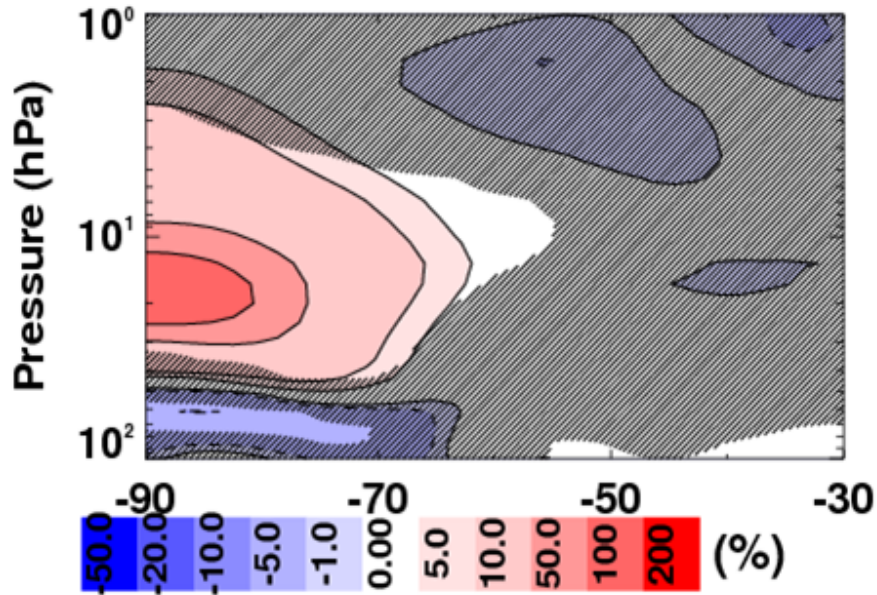
- ▶ CESM1.0.3
- ▶ FW Compset
- ▶ 1.9x2.5 degree horizontal resolution.

Name	Run Length (spinup)	Solar Flux f10.7	Ap Index	Auroral EPP
Low EPP Low Solar	42 years (2 years)	72	3	Low
High EPP Low Solar	42 years (2 years)	72	27	High
Low EPP High Solar	42 years (2 years)	210	3	Low
High EPP High Solar	42 years (2 years)	210	27	High

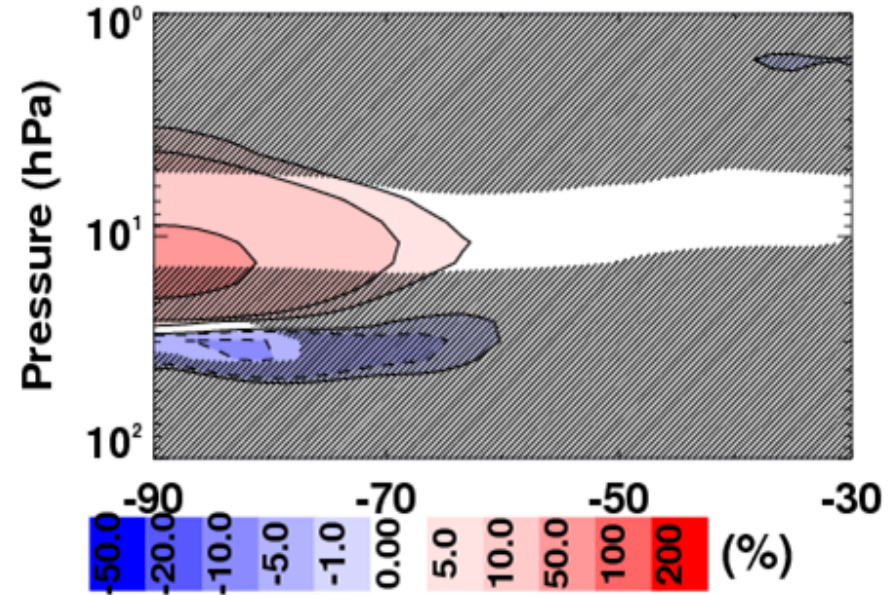
WACCM Simulation Comparisons

High EPP–Low EPP (December)

ΔNO_x High Solar
High EPP - Low EPP

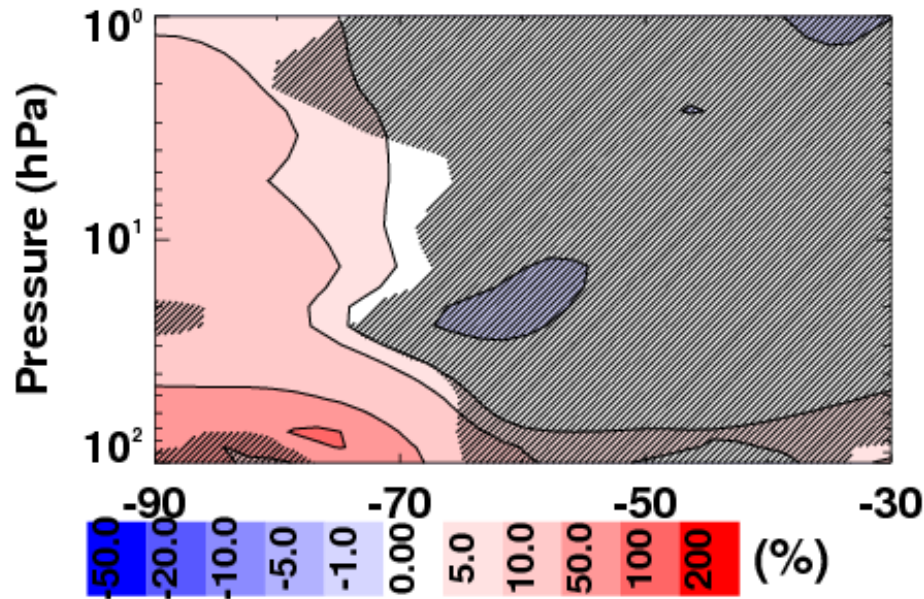


ΔNO_x Low Solar
High EPP - Low EPP

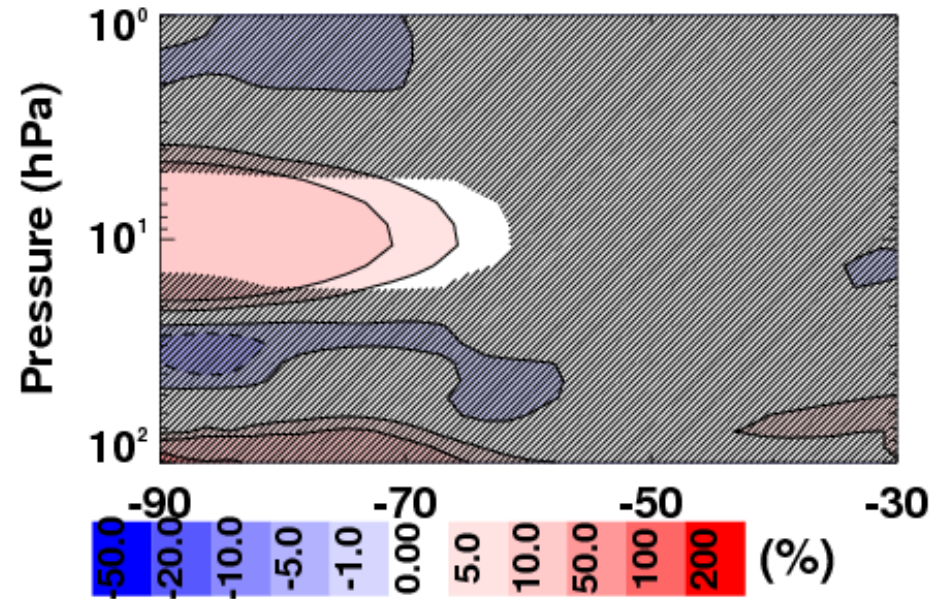


High EPP–Low EPP (December)

Δ CIONO₂ High Solar
High EPP - Low EPP

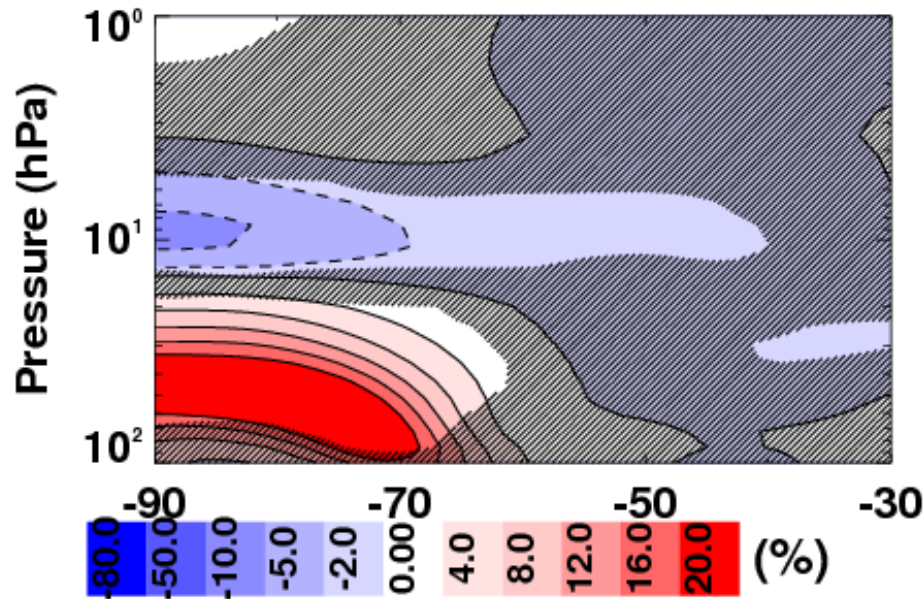


Δ CIONO₂ Low Solar
High EPP - Low EPP

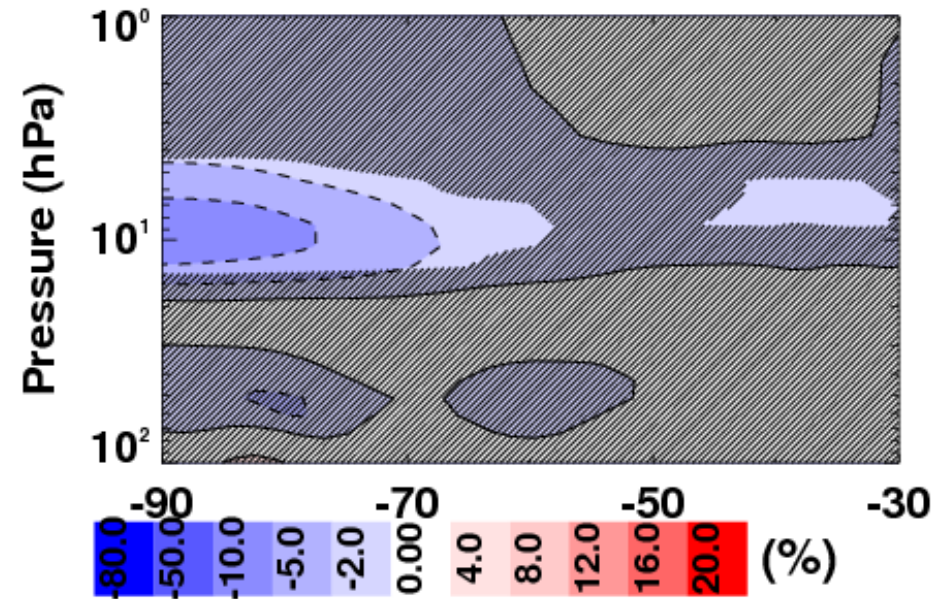


High EPP–Low EPP (December)

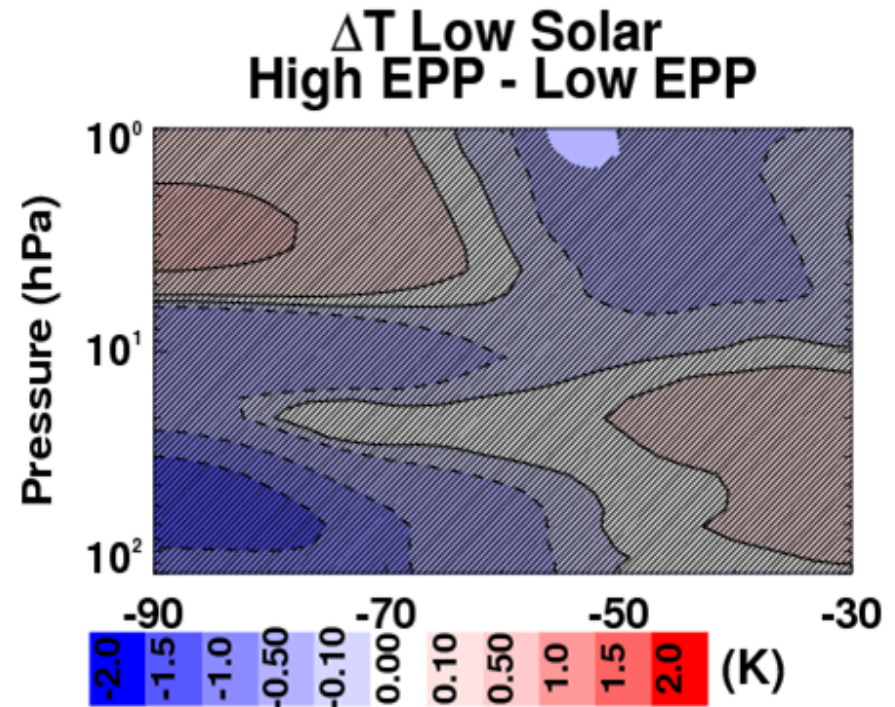
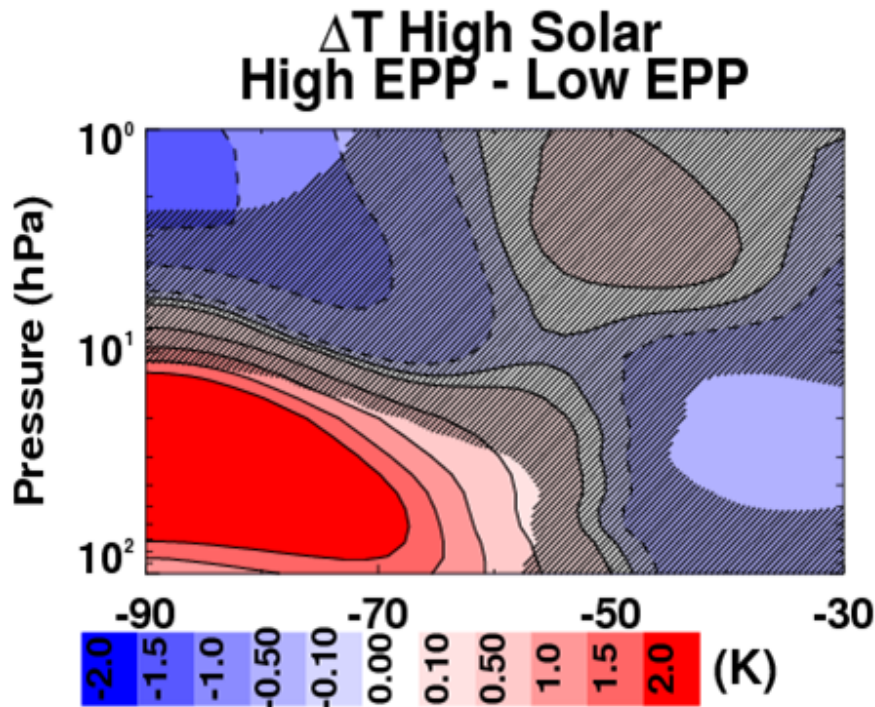
ΔO_3 High Solar
High EPP - Low EPP



ΔO_3 Low Solar
High EPP - Low EPP

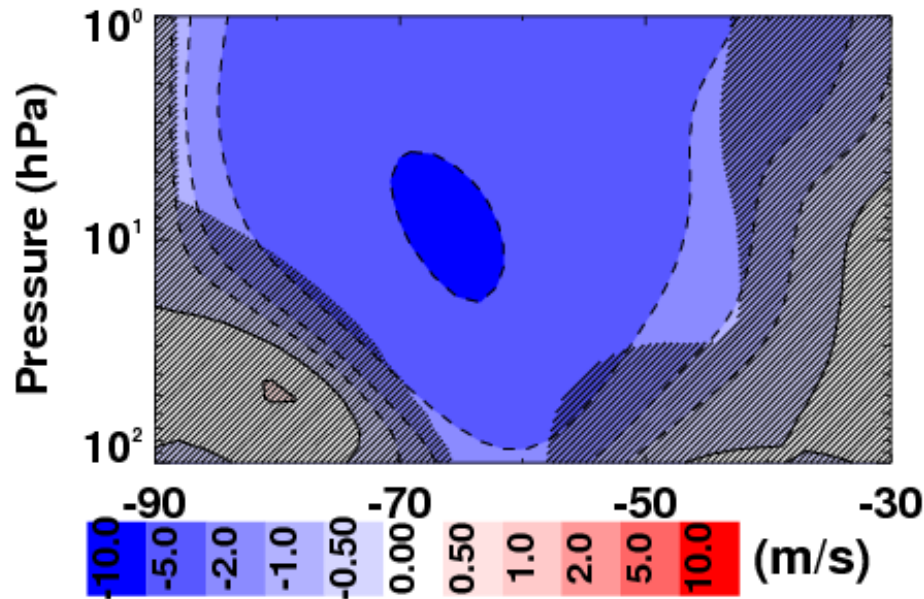


High EPP–Low EPP (December)

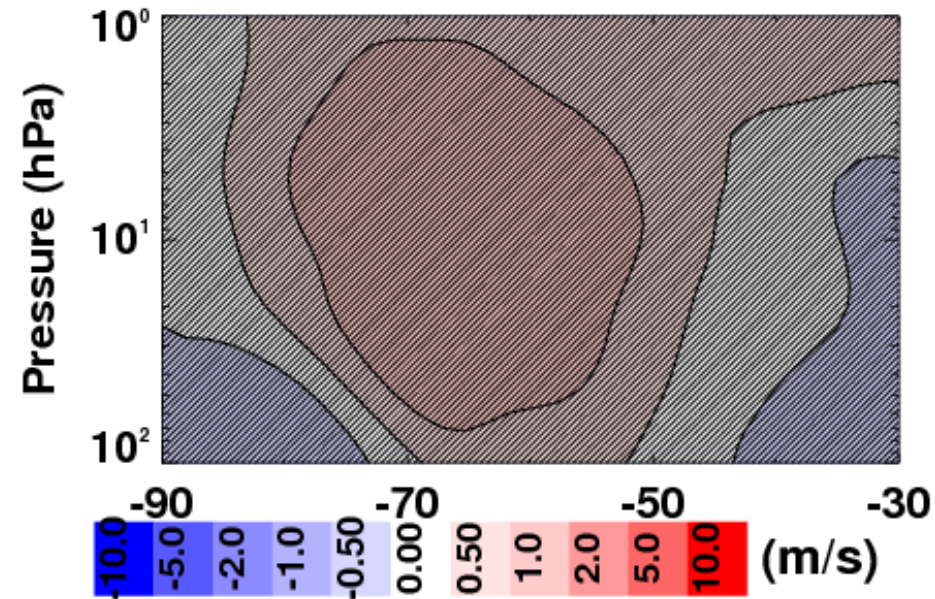


High EPP–Low EPP (December)

ΔU High Solar
High EPP - Low EPP

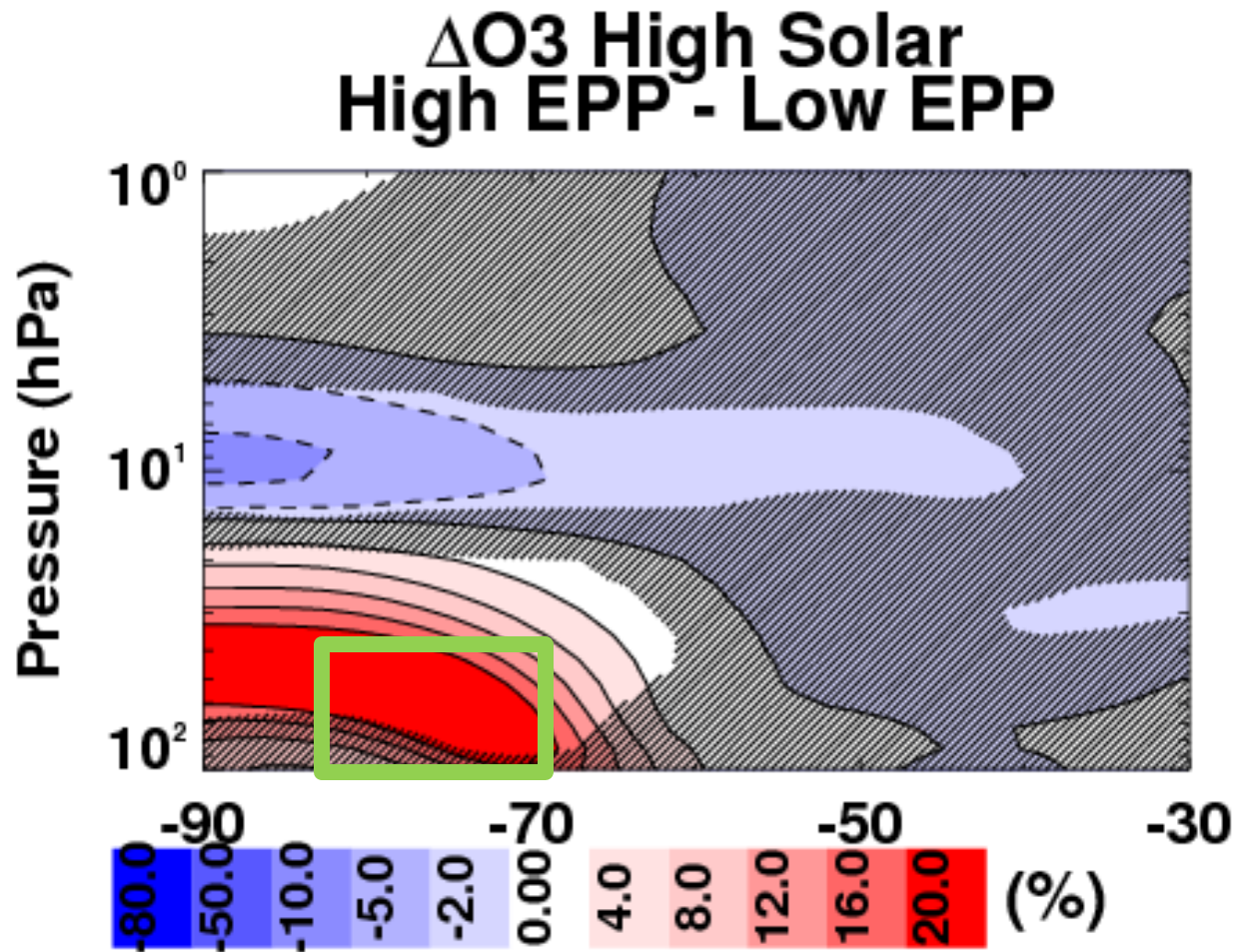


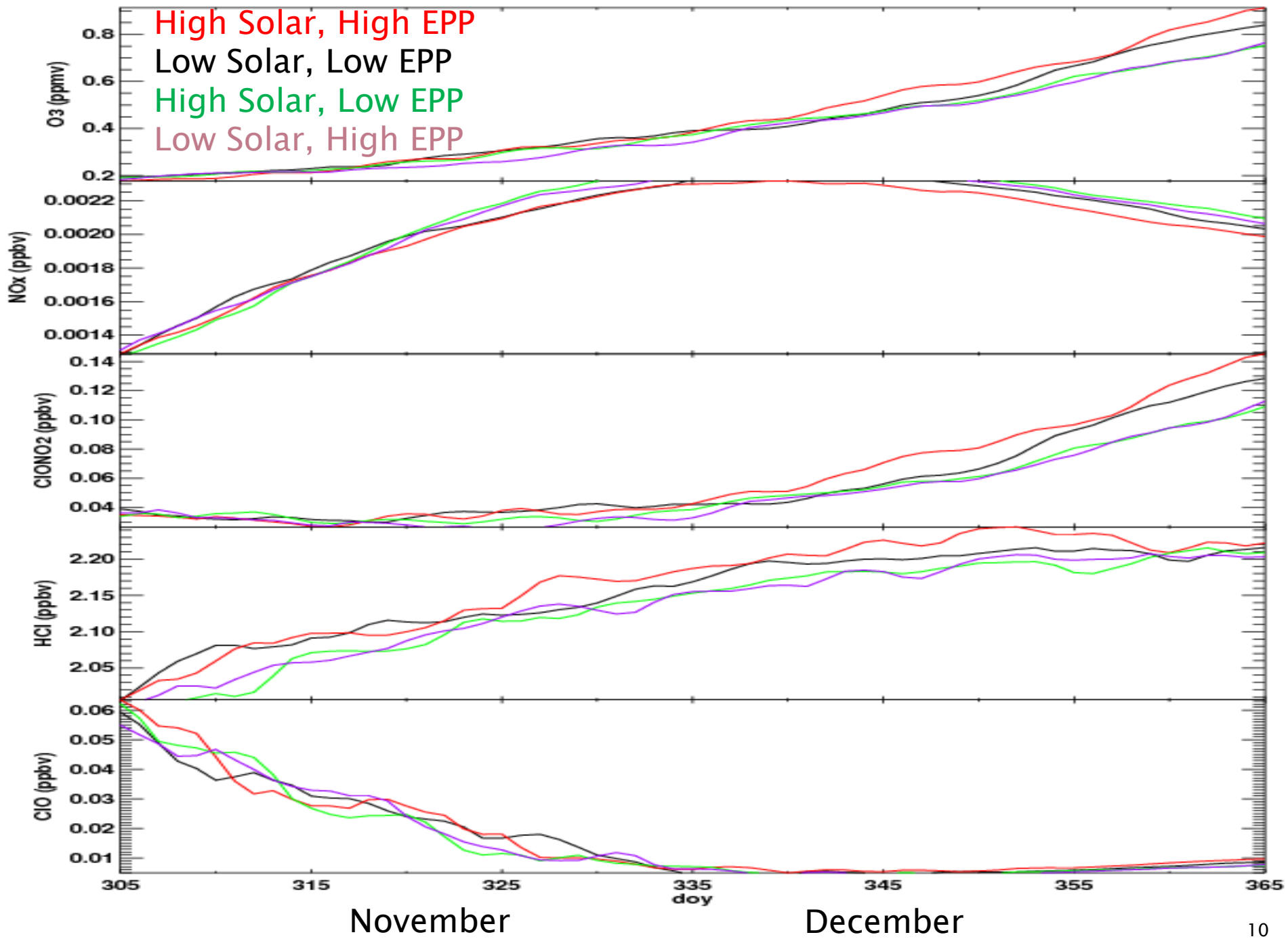
ΔU Low Solar
High EPP - Low EPP



Binning Data

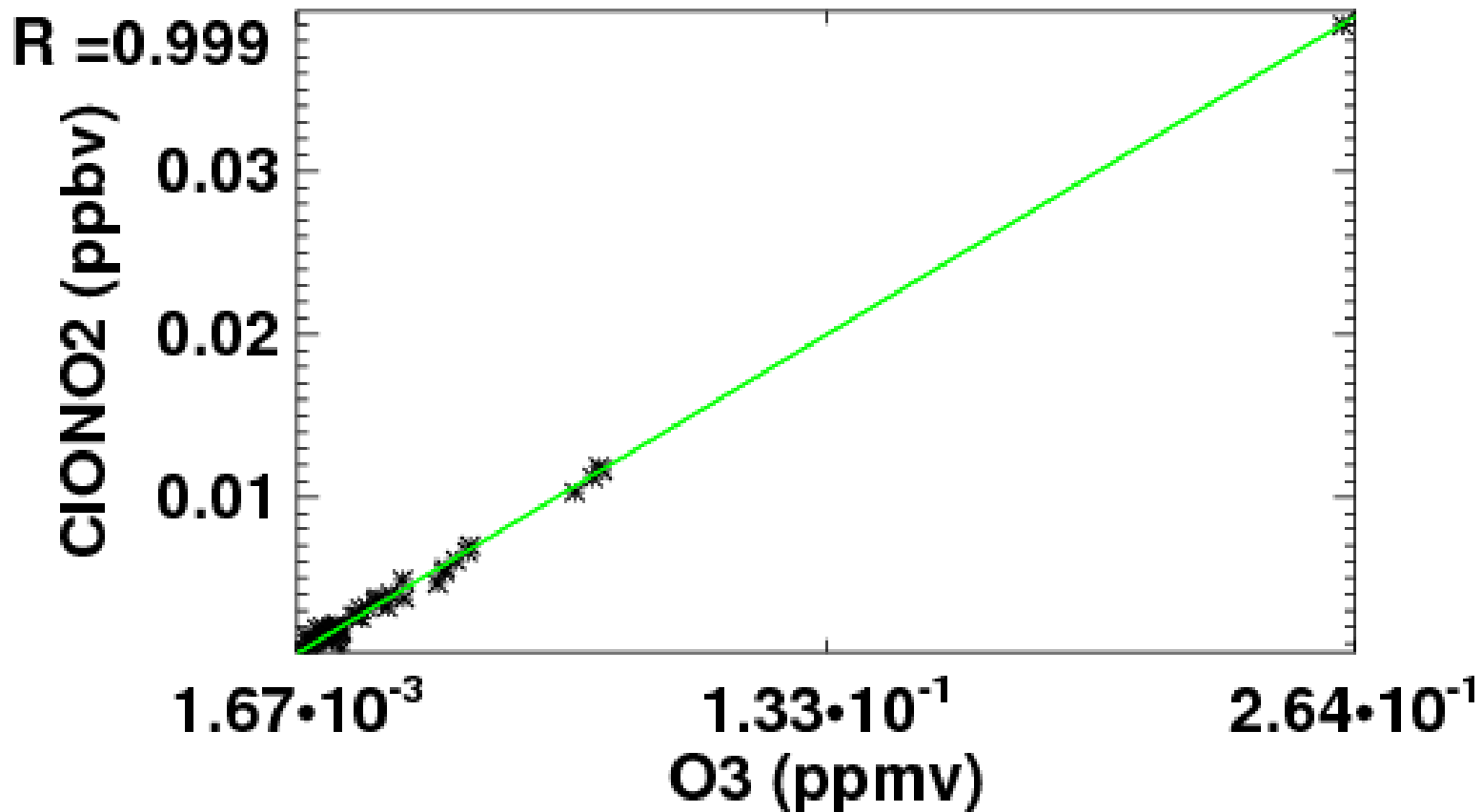
- ▶ Bin average daily data:
 - 80S–70S latitude
 - 50–100 hPa





High Pass Filtered Data Extreme Correlation

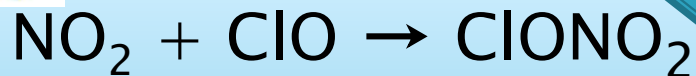
WACCM4 High Solar, High EPP High Pass Data Correlation



EPP and Solar Irradiance to O₃ Mechanism



EPP-NO_x descends through atmosphere.



Less O₃ destruction.

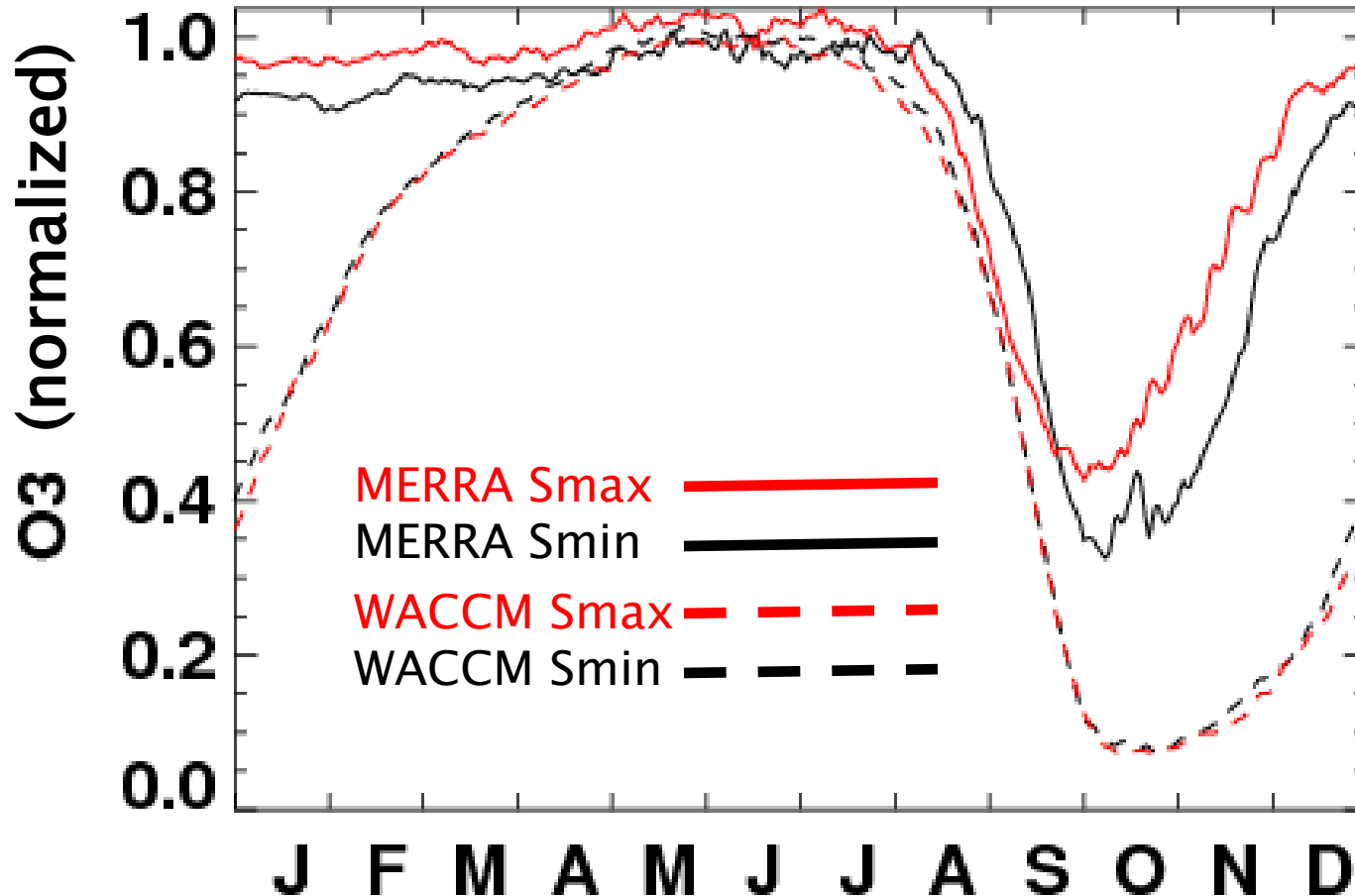
The “real” world.

- ▶ Results are compared to NASA Modern-ERa Retrospective Analysis (MERRA).
- ▶ MERRA dataset goes back to 1979.
- ▶ Split into 2 datasets:
 - **Solar Maximum**: Take peak solar max years and surrounding years (+/- 1 year). (8 years total)
 - **Solar Minimum**: Same as solar maximum but for solar minimum. (9 years total).
- ▶ These results are preliminary! (I plan on investigating more robust cutoffs for separation of MERRA data.)



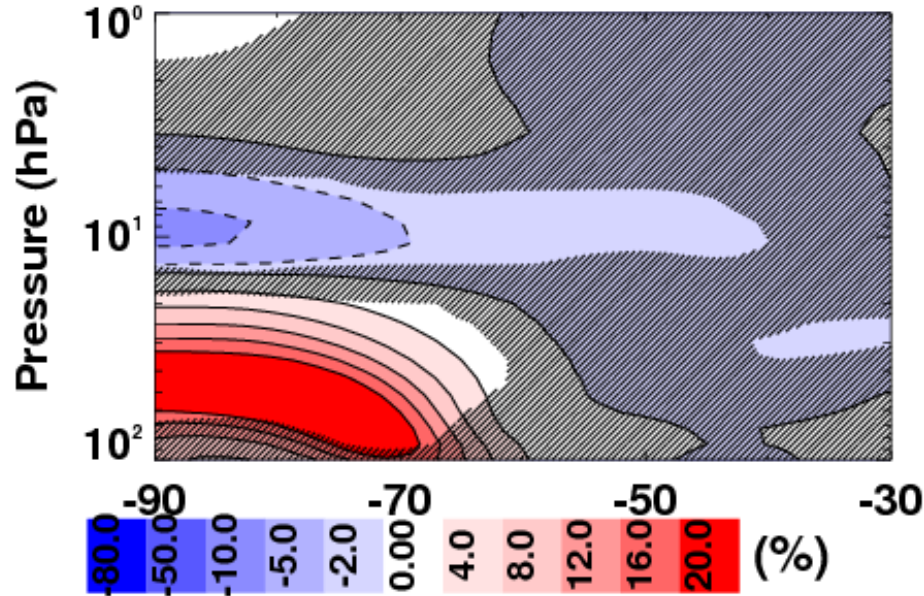
WACCM and MERRA

O3 MERRA and WACCM



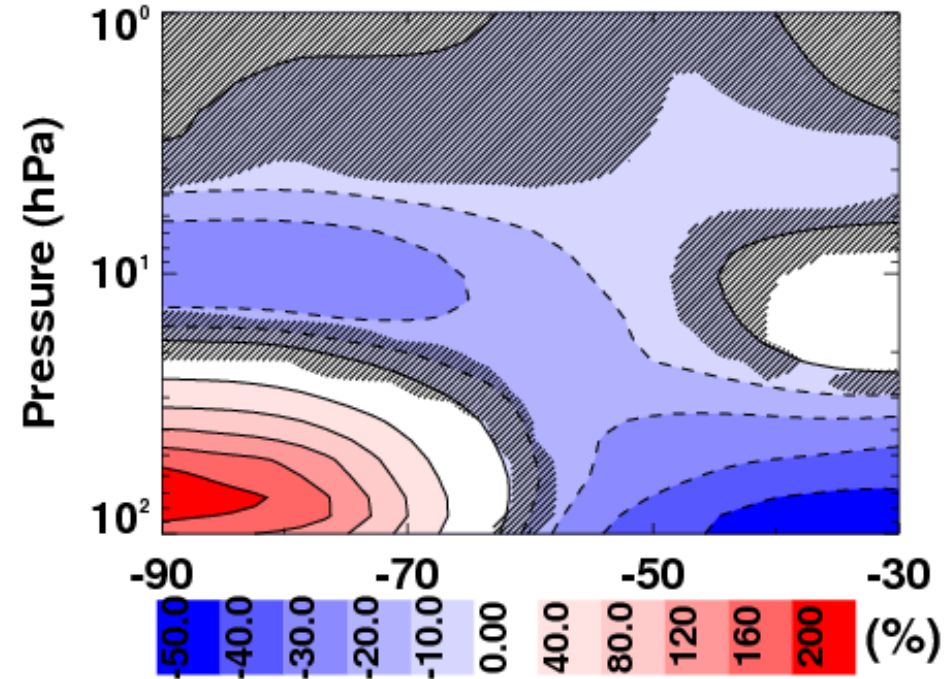
WACCM (Dec)

ΔO_3 High Solar
High EPP - Low EPP



MERRA (Oct)

MERRA ΔO_3 Solar Max - Solar Min



Conclusions

- ▶ WACCM and MERRA show significant changes in annual O₃ recovery caused by solar cycle (EPP and Solar irradiance).
- ▶ In the lower stratosphere, WACCM O₃ changes (20% increase) are smaller than MERRA (200% increase) and occur at different times (early December for WACCM, late October for MERRA).
 - This can be explained by persisting vortex (known issue) in WACCM.
- ▶ Results suggest that EPP and Solar irradiance changes associated with the solar cycle can create large differences in annual SH polar O₃ recovery.

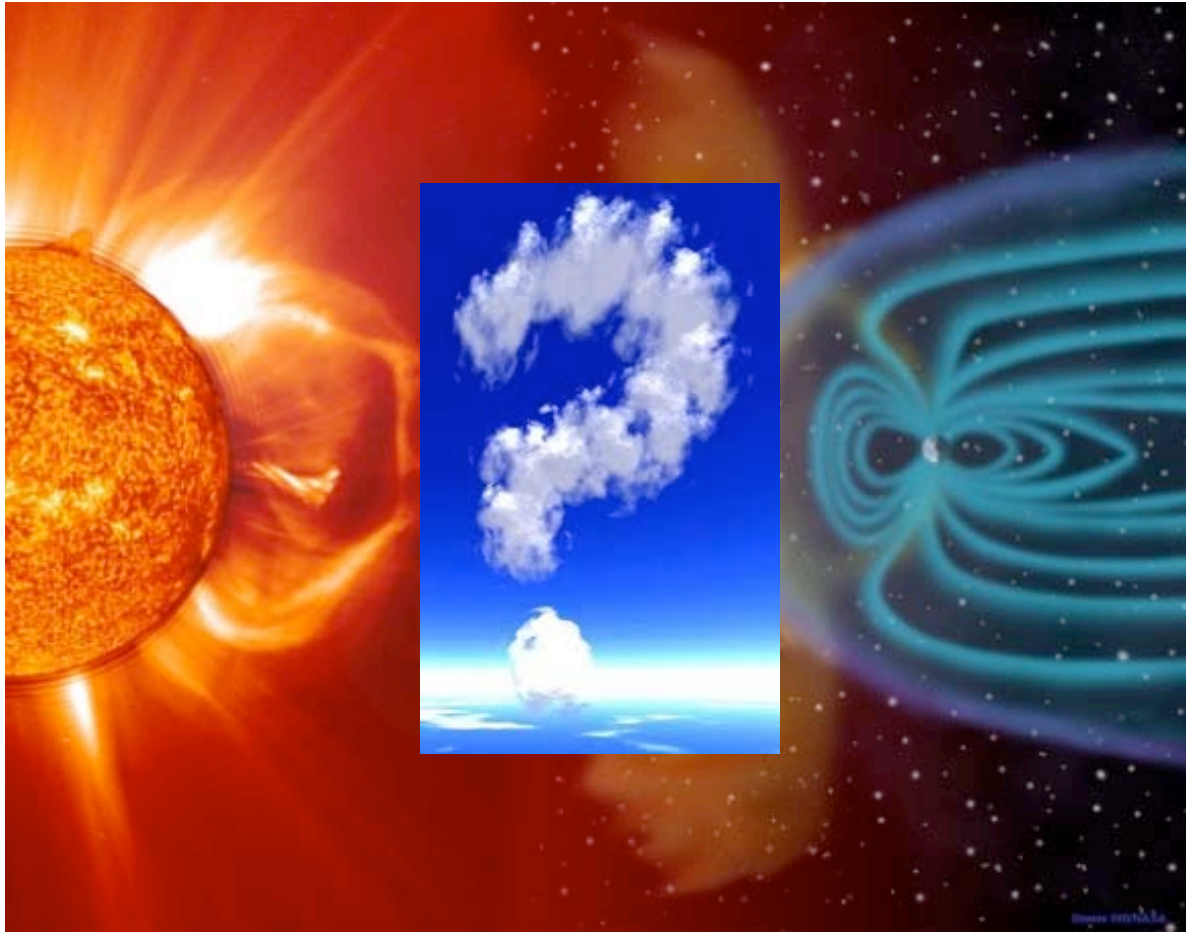
What's Next?

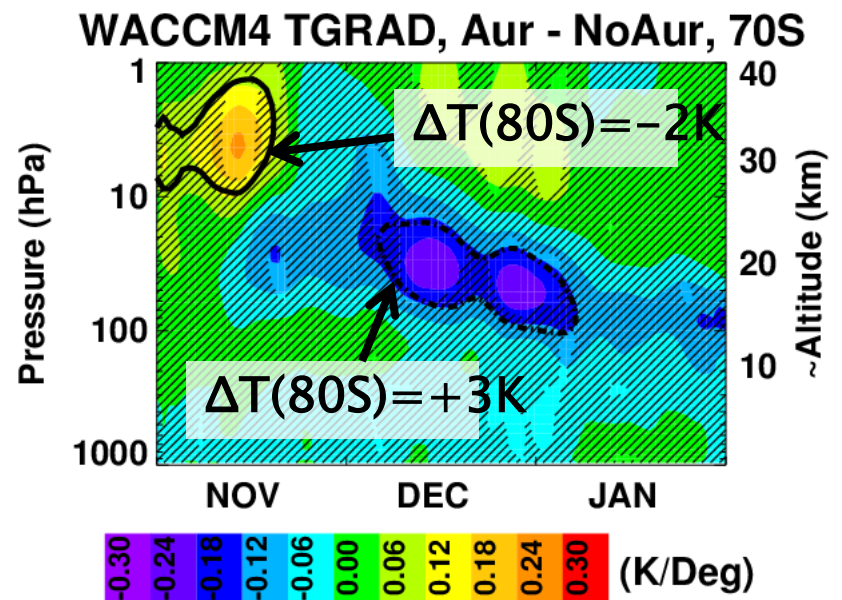
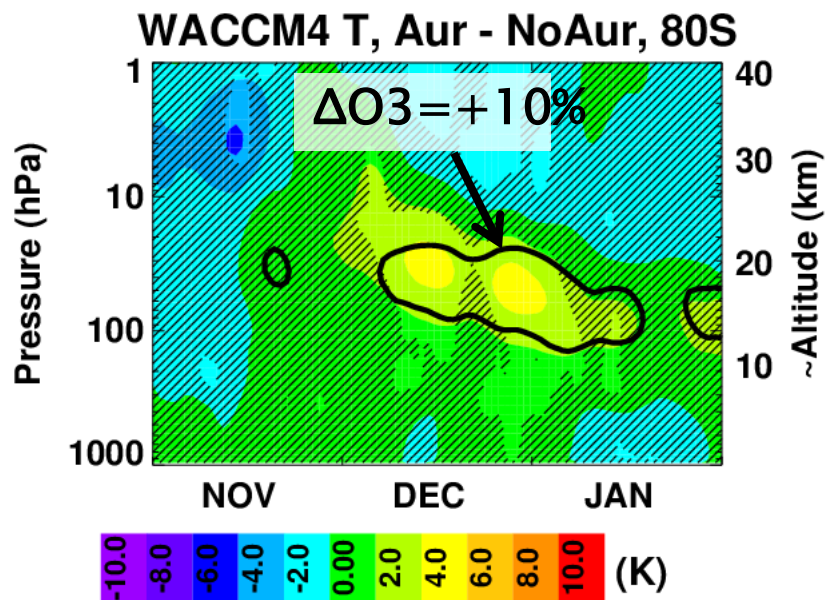
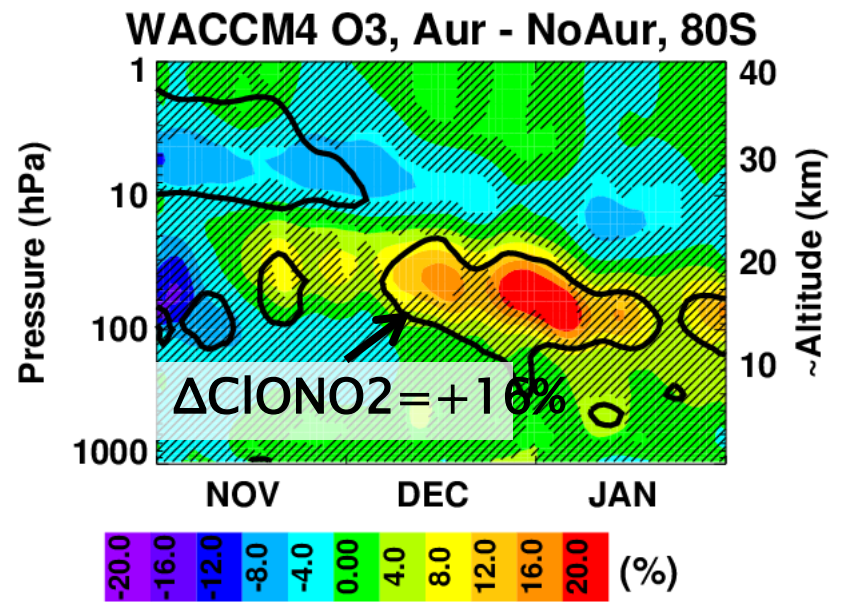
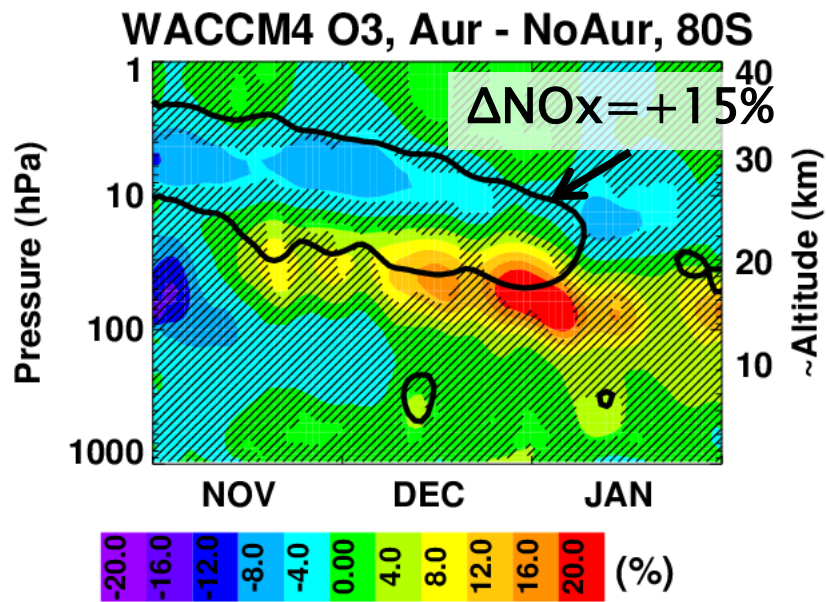
- ▶ Look for MERRA signal directly in satellite data.
 - Likely will use the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) instrument on board ENVISAT.
 - This will allow analysis of mechanism as well, since MIPAS observes O_3 and $ClONO_2$.
- ▶ Include other types of EPP.



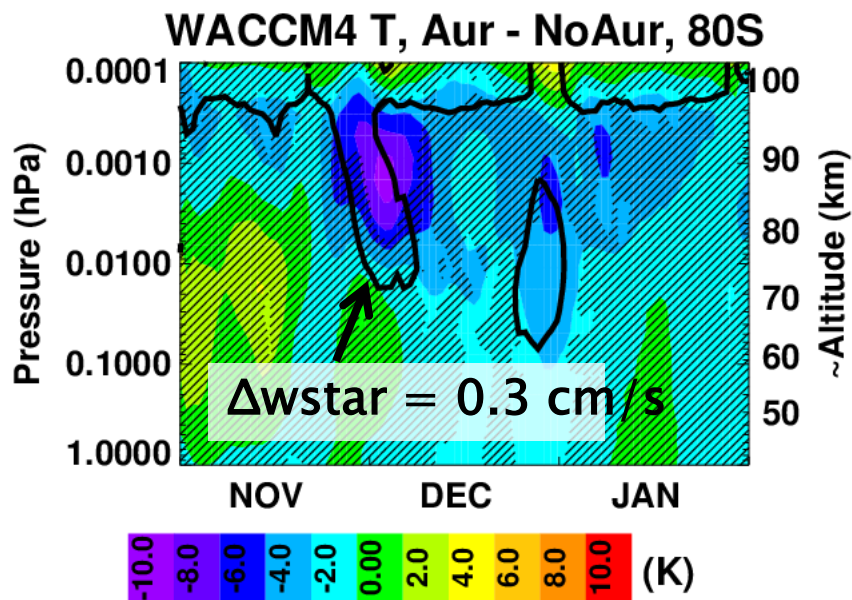
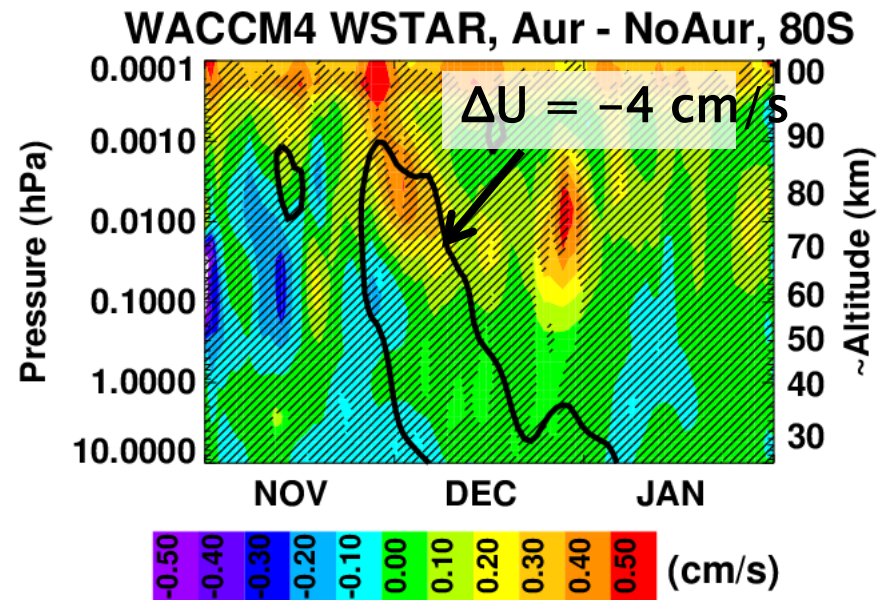
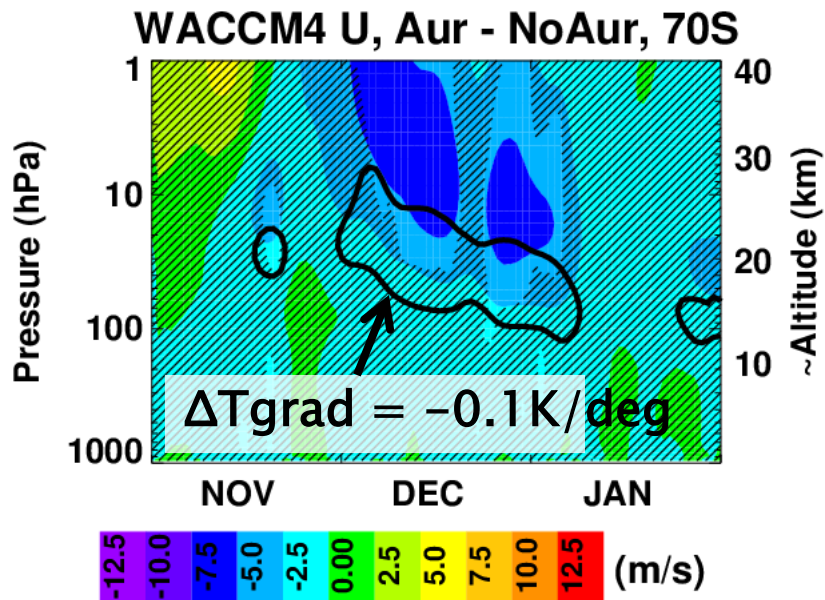
Thank You!

Questions





Upper left: Increased EPP-NO_x causes O₃ depletion and increased ClONO₂. Upper right: Increased ClONO₂ mitigates O₃ loss. Lower left: Increased O₃ causes temperature to increase. Lower right: Change in polar T leads to change in latitudinal T gradient at 70S.



Upper left: Changing latitudinal gradient in T causes U at 70S above to become more negative (easterly/summer) .

Upper right: Changing U at 70S affects GW filtering so that w_{bar}^* above, at 80S, is increased (i.e., stratospheric zonal wind more summer-like, leading to more ascent in the mesosphere) .

Lower left: Increase in w_{bar}^* leads to mesospheric cooling at 80S.

