

# Fast-J & Cloud-J -- an update

Michael Prather, Juno Hsu, Alex Nicolau, Alex Veidenbaum (UC Irvine)  
Philip Cameron-Smith (LLNL) & Michael Prather

**CESM Chemistry-Climate WG, June 2015, Breckenridge**

Geosci. Model Dev. Discuss., 8, 4051-4073, 2015  
www.geosci-model-dev-discuss.net/8/4051/2015/  
doi:10.5194/gmdd-8-4051-2015  
© Author(s) 2015. This work is distributed  
under the Creative Commons Attribution 3.0 License.



Geoscientific Model Development  
An interactive open-access journal of the European Geosciences Union

| EGU.eu |

**Model Description Paper**

27 May 2015

## Photolysis rates in correlated overlapping cloud fields: Cloud-J 7.3

**M. J. Prather**

Earth System Science Department, UC Irvine California, USA

### Review Status

This discussion paper is under review for the journal Geoscientific Model Development (GMD).

Received: 29 April 2015 – Accepted: 06 May 2015 – Published: 27 May 2015

**Abstract.** A new approach for modeling photolysis rates ( $J$  values) in atmospheres with fractional cloud cover has been developed and implemented as Cloud-J – a multi-scattering eight-stream radiative transfer model for solar radiation based on Fast-J. Using observed statistics for the vertical correlation of cloud layers, Cloud-J 7.3 provides a practical and accurate method for modeling atmospheric chemistry. The combination of the new maximum-correlated cloud groups with the integration over all cloud combinations represented by four quadrature atmospheres produces mean  $J$  values in an atmospheric column with root-mean-square errors of 4% or less compared with 10–20% errors using simpler approximations. Cloud-J is practical for chemistry-climate models, requiring only an average of 2.8 Fast-J calls per atmosphere, vs. hundreds of calls with the correlated cloud groups, or 1 call with the simplest cloud approximations. Another improvement in modeling  $J$  values, the treatment of volatile organic compounds with pressure-dependent cross sections is also incorporated into Cloud-J.

**Citation:** Prather, M. J.: Photolysis rates in correlated overlapping cloud fields: Cloud-J 7.3, Geosci. Model Dev. Discuss., 8, 4051-4073, doi:10.5194/gmdd-8-4051-2015, 2015.

fast-JX ver-7.3 standalone CTM code  
 UCI FJX v7.3+ JPL10+IUPAC 2014 fixes - same data as FJX\_spec\_6.8d.dat

18

w-eff wavelength (nm) notes: flux-weighted average over all intervals in bins  
 solflx solar #/cm2/s notes: SUSIM average 11Nov94(low) + 29Mar92(med-high)  
 solarw solarheat W/m2 notes: from SUSIM, checks with other refs.  
 Y-PAR photosyn act rad notes: Mccree 1972ab PAR spectrum  
 Raylay Rayleigh Scatter notes: flux weighted mean cross sections (cm^2)

GlyAld HOCH2CHO > notes: Glycol Aldehyde => CH2OH+HCO[0.83] CH3OH+CO[0.10] OH+CH2CHO[0.07]  
 JPL10  
 MEKeto CH3COC2H5 > notes: Methylene Ketone => C2H5+CH3CO[0.85] CH3+C2H5CO[0.15] X67  
 MEKeto notes:  
 PrAld C2H5CHO >C2H5+ notes: Propionaldehyde(propanal) => C2H5+HCO JPL10

X-sects	1	O2	O2=O+O	3	180.0	260.0	300.0	298.0
X-sects	2	O3	O3-total	3	218.0	258.0	298.0	298.0
X-sects	3	O3(1D)	Qyl d O3=0(1D)+O2	3	200.0	260.0	320.0	298.0
X-sects	4	NO	NO=N+O	x 1	298.0			
X-sects	5	H2COa	H2CO=>H+HCO	2	223.0	298.0		
X-sects	6	H2COb	H2CO=>H2+CO	2	223.0	298.0		
X-sects	7	H2O2	H2O2=>OH+OH	2	200.0	300.0		
X-sects	8	CH3OOH	CH3OOH=>CH3O+OH	1	298.0			
X-sects	9	N02	N02=>N0+O	2	200.0	294.0		
X-sects	10	N03	N03=N02+O/N0+O2	2	190.0	298.0		
X-sects	11	N2O5	N2O5=>N02+N03	2	233.0	300.0		
X-sects	12	HN02	HNO2=>OH+N0	1	300.0			
X-sects	13	HN03	HNO2=>OH+N02	2	200.0	300.0		
X-sects	14	HN04	H02N02=>H02+N02	1	300.0			
X-sects	15	ClN03a	ClN03=Cl+N03	x 2	200.0	300.0		
X-sects	16	ClN03b	ClN03=ClO+N02	x 2	200.0	300.0		
X-sects	17	Cl2	Cl2=>Cl+Cl	x 2	200.0	300.0		
X-sects	18	HOCl	HOCl=>OH+Cl	x 1	300.0			
X-sects	19	OClO	OClO=>O+ClO	x 1	204.0			
X-sects	20	Cl2O2	ClOCl=>ClO+ClO	x 1	220.0			
X-sects	21	ClO	ClO=>Cl+O	x 1	300.0			
X-sects	22	BrO	BrO=>Br+O	x 1	300.0			
X-sects	23	BrN03	BrON02=>BrO+N02	x 2	200.0	300.0		
X-sects	24	HOBr	HOBr=>OH+Br	x 1	300.0			
X-sects	25	BrCl	BrCl=>Br+Cl	x 2	200.0	300.0		
X-sects	26	N2O	N2O=>N2+O	x 2	200.0	300.0		
X-sects	27	CFCl3	CFCl3=>CFCl2+Cl	x 2	220.0	298.0		
X-sects	28	CF2Cl2	CF2Cl2=>CF2Cl+	x 2	220.0	300.0		
X-sects	29	F113	CF3CCl3=>	x 2	210.0	300.0		
X-sects	30	F114	CF3CFCl2=>	x 2	210.0	300.0		
X-sects	31	F115	CF3CF2Cl=>	x 1	300.0			
X-sects	32	CCl4	CCl4=>	x 2	200.0	300.0		

298.0

X-sects	32	CCl4	CCl4=>	x	2	200.0	300.0
X-sects	33	CH3Cl	CH3Cl=>CH3+Cl	x	2	200.0	300.0
X-sects	34	MeCCl3	CH3CCl3=>CH3CCl2	x	2	200.0	300.0
X-sects	35	CH2Cl2	CH2Cl2=>CH2Cl+Cl	x	2	200.0	300.0
X-sects	36	CHF2Cl	CHF2Cl=>CHF2+Cl	x	2	200.0	300.0
X-sects	37	F123	CF3CCHCl2=>	x	2	210.0	295.0
X-sects	38	F141b	CH3CFCl2=>	x	2	200.0	300.0
X-sects	39	F142b	CH3CF2Cl=>	x	2	210.0	298.0
X-sects	40	CH3Br	CH3Br=>CH3+Br	x	2	200.0	300.0
X-sects	41	H1211	CF2ClBr=>	x	2	200.0	300.0
X-sects	42	H1301	CF3Br=>	x	2	200.0	300.0
X-sects	43	H2402	CF2BrCF2Br=>	x	2	200.0	300.0
X-sects	44	CH2Br2	CH2Br2=>		2	200.0	300.0
X-sects	45	CHBr3	CHBr3=>		2	210.0	300.0
X-sects	46	CH3I	CH3I=>CH3+I		2	243.0	300.0
X-sects	47	CF3I	CF3I=>CF3+I		2	243.0	300.0
X-sects	48	OCS	OCS=>CO+S		2	200.0	300.0
X-sects	49	PAN	CH3C(O)COON02=>		2	250.0	298.0
X-sects	50	CH3N03	CH3ON02=>CH3O+N0		2	200.0	300.0
X-sects	51	ActAl d	Acetal dhyde	p	3	177.0	566.0 999.0
X-sects	52	MeVK	CH3C(O)CH=CH2 >	p	3	177.0	566.0 999.0
X-sects	53	MeAcr	CH2C(CH3)CHO >		1	298.0	
X-sects	54	GlyAl d	HOCH2CHO >		1	298.0	
X-sects	55	MEKeto	CH3COC2H5 >	p	2	177.0	999.0
X-sects	56	PrAl d	C2H5CHO >C2H5+		1	298.0	
X-sects	57	MGl yxl	CH3COC(O)H >HCO+	p	3	177.0	566.0 999.0
X-sects	58	Gl yxl a	CHOCHO=>HCO+HCO	p	2	177.0	999.0
X-sects	59	Gl yxl b	CHOCHO=>HCO+HCO	p	2	177.0	999.0
X-sects	60	Gl yxl c	CHOCHO=>HCO+HCO	p	2	177.0	999.0
X-sects	61	Acet- a	Acetn=CH3CO+CH3	p	3	177.0	566.0 999.0
X-sects	62	Acet- b	Acetn=CH3+CH3+CO		3	235.0	260.0 298.0

fast-JX ver-7.3 standalone CTM code  
UCI FJX v7.3+ JPL10+IUPAC 2014 fixes - same data as FJX\_spec\_6.8d.dat

18

w-eff wavelength (nm) notes: flux-we  
solflx solar #/cm2/s notes: SUSIM a  
solarw solarheat W/m2 notes: from SU  
Y-PAR photosyn act rad notes: Mccree  
Raylay Rayleigh Scatter notes: flux we

GlyAl d HOCH2CHO > notes: Glycol  
JPL10  
MEKeto CH3COC2H5 > notes: Methyl e  
MEKeto notes:  
PrAld C2H5CHO >C2H5+ notes: Propion

X-sects 1 O2 O2=O+O  
X-sects 2 O3 O3-total  
X-sects 3 O3(1D) Qy1 d O3=O(1D)+O2  
X-sects 4 NO NO=N+O  
X-sects 5 H2COa H2CO=>H+HCO  
X-sects 6 H2COb H2CO=>H2+CO  
X-sects 7 H2O2 H2O2=>OH+OH  
X-sects 8 CH3OOH CH3OOH=>CH3O+OH  
X-sects 9 NO2 NO2=>NO+O  
X-sects 10 NO3 NO3=NO2+O/NO+O2  
X-sects 11 N2O5 N2O5=>NO2+NO3  
X-sects 12 HNO2 HONO=>OH+NO  
X-sects 13 HNO3 HONO2=>OH+NO2  
X-sects 14 HNO4 HO2NO2=>HO2+NO2  
X-sects 15 ClNO3a ClNO3=Cl+NO3  
X-sects 16 ClNO3b ClNO3=ClO+NO2  
X-sects 17 Cl2 Cl2=>Cl+Cl  
X-sects 18 HOCl HOCl=>OH+Cl  
X-sects 19 OClO OClO=>O+ClO  
X-sects 20 Cl2O2 ClOOCl=>ClO+ClO  
X-sects 21 ClO ClO=>Cl+O  
X-sects 22 BrO BrO=>Br+O  
X-sects 23 BrNO3 BrONO2=>BrO+NO2  
X-sects 24 HOBr HOBr=>OH+Br  
X-sects 25 BrCl BrCl=>Br+Cl  
X-sects 26 N2O N2O=>N2+O  
X-sects 27 CFC13 CFC13=>CFC12+Cl  
X-sects 28 CF2Cl2 CF2Cl2=>CF2Cl+  
X-sects 29 F113 CF3CCl3=>  
X-sects 30 F114 CF3CFC12=>  
X-sects 31 F115 CF3CF2Cl=>  
X-sects 32 CCl4 CCl4=>

## Fast-JX v6.8d

Last F77 heritage version of Fast-J

## Fast-JX v7.0d

F90 CAM5

## Fast-JX v7.1c

F90 WACCM (allows WACCM-J <200 nm)

## Fast-JX v7.2

Cloud-J for UCI CTM (Neu 2007 MAX-RAN)

## Cloud-J v7.3

New correlated cloud overlap (Prather 2015)

## Solar-J v7.3

Extend FJX bins, clouds & molecules to 5 μm

298.0

x 2 220.0 300.0  
x 2 210.0 300.0  
x 2 210.0 300.0  
x 1 300.0  
x 2 200.0 300.0

X-sects 58 Glyxl a CHOCHO=>HCO+HCO p 2 177.0 999.0  
X-sects 59 Glyxl b CHOCHO=>HCO+HCO p 2 177.0 999.0  
X-sects 60 Glyxl c CHOCHO=>HCO+HCO p 2 177.0 999.0  
X-sects 61 Acet- a Acetn=CH3CO+CH3 p 3 177.0 566.0 999.0  
X-sects 62 Acet- b Acetn=CH3+CH3+CO 3 235.0 260.0 298.0

## Cloud-J v7.3 is really new and presents a breakthrough in Scale Independence

A problem with deterministic cloud overlap algorithms using correlation lengths is that they fail the scale-independence test in that the number of atmospheres to be averaged over **grows as  $2^{NL}$** , where NL is the number of layers.

Cloud-J eliminates this problem with two quantizations that bring the number of cloud combinations to a fixed maximum, independent of NL:

- 1) Group the cloudy layers within one correlation length into one maximum overlap group (*new to Prather 2015*), creating  $6\frac{1}{2}$  MAX groups (the  $\frac{1}{2}$  includes the cirrus shield separation). Clouds close together are MAX overlapped.
- 2) Quantize the fractional cloud cover in any to the nearest 10<sup>th</sup> percentile (0%, 10%, 20%, ... 100%), ensuring that no matter how many layers are in each MAX group, there are at most 10 different ICAs (*from Neu 2007*)
- 3) Assume that each MAX group is correlated with the one above at a level corresponding to about one e-fold (*algorithm new to Prather 2015*)
- 4) The maximum # of ICAs to characterize is less than  $5 \times 10^6$ , independent of NL. With coding in Cloud-J, this computational time is inconsequential. The # of ICAs for NL=40 and a straightforward correlation algorithm is as large as  $10^{12}$ .

## Cloud-J

limited data passed to CLOUD\_JX (only interface to rest of model)

```
call CLOUD_JX (U0, SZA, REFLB, SOLF, FGO, LPRTJ, PPP, ZZZ, TTT, &  
DDD, RRR, OOO, LWP, IWP, REFFL, REFFI, CLF, CWC, &  
AERSP, NDXAER, L1_, AN_, VALJXX, JVN_, &  
CLDFLAG, NRANDO, IRAN, L3RG, NICA, JCOUNT)
```

***Only needs from CAM are profiles of 5 cloud quantities***

```
LWP/IWP = Liquid/Ice water path (g/m2)  
REFFL/REFFI = R-effective(microns) in liquid/ice cloud  
CLF = cloud fraction (0.0 to 1.0)
```

***and profiles of aerosol quantities***

```
AERSP = aerosol path (g/m2)  
NDXAER = aerosol index type
```

## Photolysis rates in correlated overlapping cloud fields: Cloud-J 7.3

A new approach for modeling the observed vertical correlation of fractional cloud layers is presented.

The vertical correlation length  $L$  (km) is based on observations:

*1.5 km near surface – to – 3 km in upper troposphere.*

For practicality, 6 maximally overlapped (MAX) groups are based on  $L$ :

1 <sup>st</sup>	0 – 1.5 km	2 <sup>nd</sup>	1.5 – 3.5 km	3 <sup>rd</sup>	3.5 – 6 km
4 <sup>th</sup>	6 – 9 km	5 <sup>th</sup>	9 – 13 km	6 <sup>th</sup>	>13 km

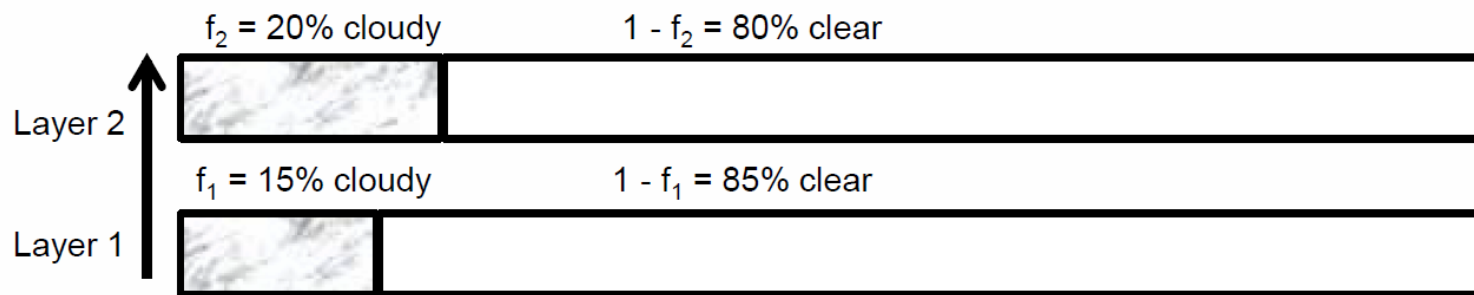
a 7<sup>th</sup> MAX group is split from the top group, if cirrus shields present.

A new algorithm for correlated overlap of the MAX groups is presented:

MAX-COR has limits of      MAX-RAN for zero correlation (cc=0)  
one MAX group for full correlation (cc=1)

Correlation coefficient = 0.33 ( $\sim 1/e$ ) between MAX groups is set

## Cloud overlap algorithms:

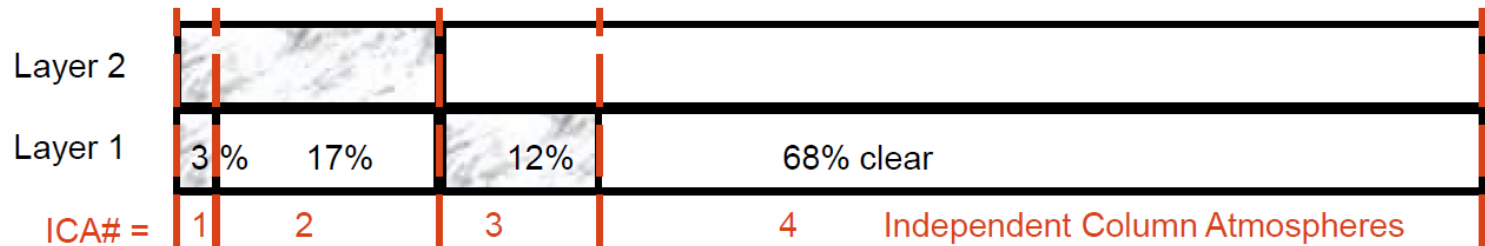


Layer 2: cloudy fraction =  $f^{L2} = 20\%$  clear fraction =  $1 - f^{L2} = 80\%$

Layer 1: cloudy fraction =  $f^{L1} = 15\%$  clear fraction =  $1 - f^{L1} = 85\%$

## Random overlap (RAN)

RANdom overlap (between layers here, or between MAX-Groups)



ICAs #1 and #2 fall beneath cloudy fraction in Layer 2:

$$w^{L1}(\#1) = f^{L1} \quad (1)$$

$$w^{L1}(\#2) = 1 - f^{L1} \quad (2)$$

$$w^{L2}(\#1) = w^{L2}(\#2) = f^{L2} \quad (3)$$

$$W^{L1-L2}(\#1) = w^{L1}(\#1)w^{L2}(\#1) = f^{L1}f^{L2} = 0.15 \times 0.20 = 3\% \text{ (from Fig. 1)} \quad (4)$$

$$W^{L1-L2}(\#2) = w^{L1}(\#2)w^{L2}(\#2) = (1 - f^{L1})f^{L2} = 0.85 \times 0.20 = 17\% \quad (5)$$

ICAs #3 and #4 fall beneath clear fraction in Layer 2:

$$w^{L1}(\#3) = f^{L1} \quad \text{and} \quad w^{L2}(\#3) = 1 - f^{L2} \quad (6)$$

$$w^{L1}(\#4) = 1 - f^{L1} \quad \text{and} \quad w^{L2}(\#4) = 1 - f^{L2}. \quad (7)$$

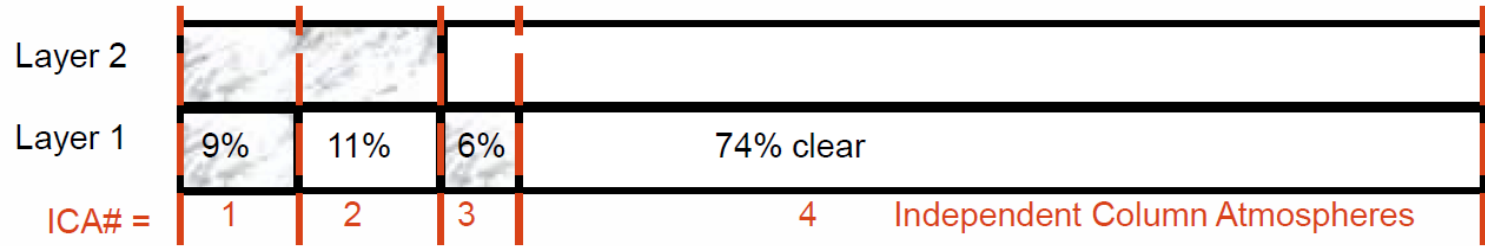
$$W^{L1-L2}(\#3) = w^{L1}(\#3)w^{L2}(\#3) = f^{L1}(1 - f^{L2}) = 0.15 \times 0.80 = 12\% \text{ (Fig. 1)} \quad (8)$$

$$W^{L1-L2}(\#4) = w^{L1}(\#4)w^{L2}(\#4) = (1 - f^{L1})(1 - f^{L2}) = 0.85 \times 0.80 = 68\% \quad (9)$$



## Correlated overlap (COR)

CORrelated overlap, with  $cc = 1/2$  (between layers here, or between MAX-Groups)



$$g = 1 + cc(1/f^{L2} - 1), \text{ subject to } g \leq 1/f^{L1} \text{ and } g \leq 1/f^{L2} \quad (10)$$

$$w^{L1}(\#1) = gf^{L1} = 3 \times 0.15 = 45\% \quad (11)$$

$$w^{L1}(\#2) = 1 - gf^{L1} = 1 - 0.45 = 55\% \quad (12)$$

$$w^{L1}(\#3) = f^{L1}(1 - gf^{L2})/(1 - f^{L2}) = 0.15 \times (1 - 3 \times 0.20)/0.80 = 7.5\% \quad (13)$$

$$w^{L1}(\#4) = 1 - w^{L1}(\#3) = 1 - 0.075 = 92.5\% \quad (14)$$

$$w^{L2}(\#1) = w^{L2}(\#2) = f^{L2} \quad (15)$$

$$w^{L2}(\#3) = w^{L2}(\#4) = (1 - f^{L2}). \quad (16)$$

ICAs #1 has more L1 cloudy fraction below L2 cloud (9%) than in RAN (3%)

$$W^{L1-L2}(\#1) = w^{L1}(\#1)w^{L2}(\#1) = gf^{L1}f^{L2} = 3 \times 0.15 \times 0.20 = 9\% \quad (17)$$

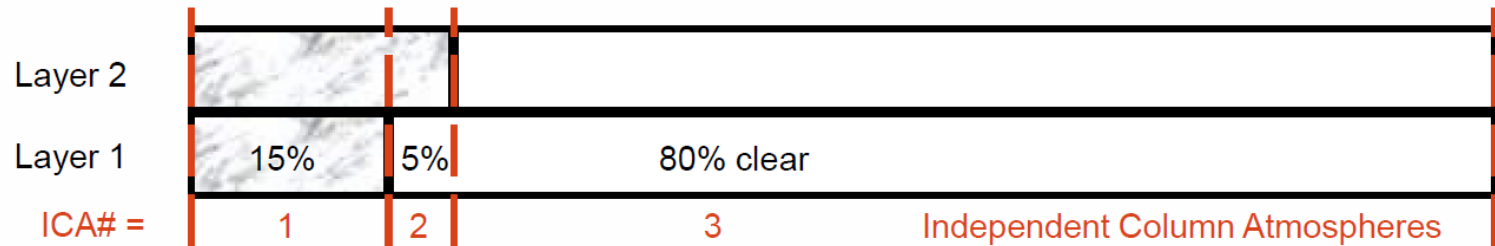
$$W^{L1-L2}(\#2) = w^{L1}(\#2)w^{L2}(\#2) = (1 - gf^{L1})f^{L2} = 0.55 \times 0.20 = 11\% \quad (18)$$

$$W^{L1-L2}(\#3) = w^{L1}(\#3)w^{L2}(\#3) = f^{L1}(1 - gf^{L2}) = 0.15 \times 0.40 = 6\% \quad (19)$$

$$W^{L1-L2}(\#4) = w^{L1}(\#4)w^{L2}(\#4) = 1 - f^{L2} - f^{L1}(1 - gf^{L2}) = 1 - 0.20 - 0.06 = 74\% \quad (20)$$

## Maximal overlap (MAX)

MAXimum overlap (connected layers become a MAX-Group)



ICAs #1 has more L1 cloudy fraction below L2 cloud (15%) than in COR (9%)

There are only 3 ICAs

A MAX group corresponds to set of  $(N1+1)$  ICAs:

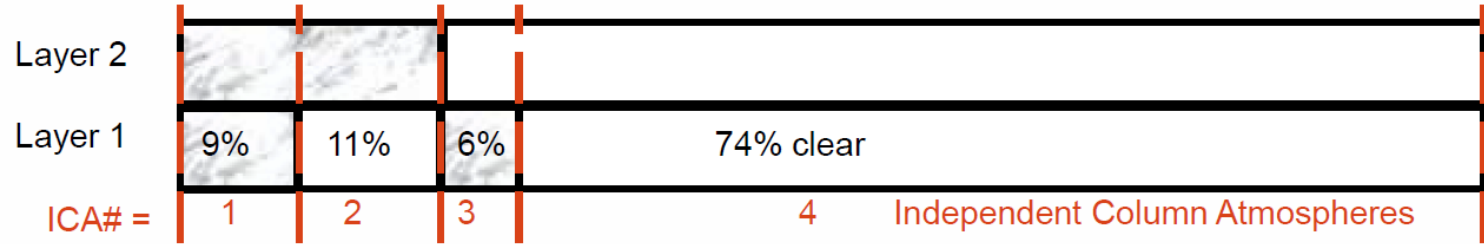
fractions  $(f_1, f_2, f_3, \dots)$  and their weights  $(w_1, w_2, w_3, \dots)$

$$F = \sum f_i \leq 1$$

$f_{J1=1:N1}^{G1}$  totaling  $F^{G1}$  and a clear-sky with fraction  $1 - F^{G1}$

## Maximal groups with correlated overlap (MAX-COR)

CORrelated overlap, with  $cc = \frac{1}{2}$  (between layers here, or between MAX-Groups)



For 2 MAX-COR groups ( $G1$  &  $G2$ ), each with a number of ICAs ( $N1+1$  &  $N2+1$ ), the numbering scheme is:

$$(1, 1), (2, 1), (3, 1), \dots, (N1 + 1, 1), (1, 2), (2, 2), (3, 2), \dots, (N1 + 1, N2 + 1) \quad (21)$$

$$J1 = (M - 1) \bmod (N1 + 1) + 1 \quad (22)$$

$$J2 = \text{integer}((M - 1)/(N1 + 1)) \bmod (N2 + 1) + 1 \quad (23)$$

The correlation enhancement factor  $g$  is based on the sum of cloudy fractions in each group

$$g = 1 + cc (1/F^{G2} - 1), \text{ subject to } g \leq 1/F^{G1} \text{ and } g \leq 1/F^{G2} \quad (24)$$

The weights for each level depend on the cloudy-clear combinations:

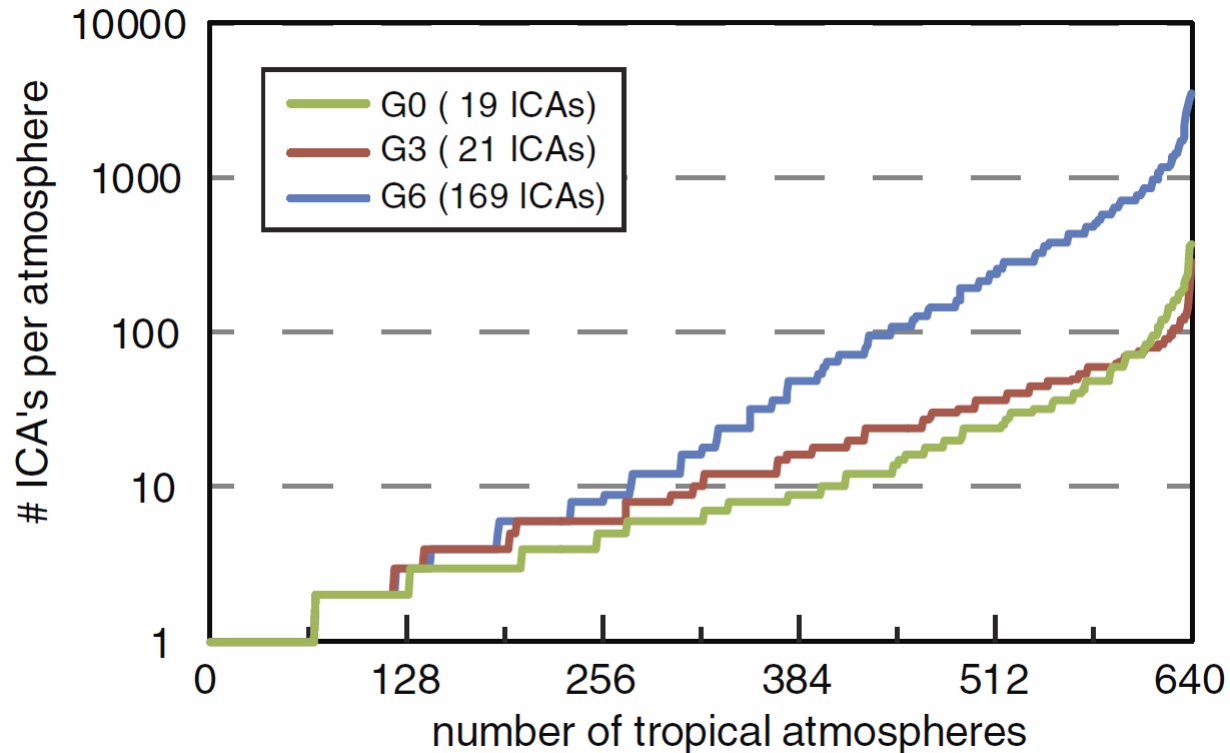
$$w^{G1}(\text{cloudy}^{G1}, \text{cloudy}^{G2}) = g f_{J1}^{G1} \quad (25)$$

$$w^{G1}(\text{clear}^{G1}, \text{cloudy}^{G2}) = 1 - \sum (\text{over cloudy } G1) g f_{J1}^{G1} = 1 - g F^{G1} \quad (26)$$

$$w^{G1}(\text{cloudy}^{G1}, \text{clear}^{G2}) = f_{J1}^{G1} (1 - g F^{G2}) / (1 - F^{G2}) \quad (27)$$

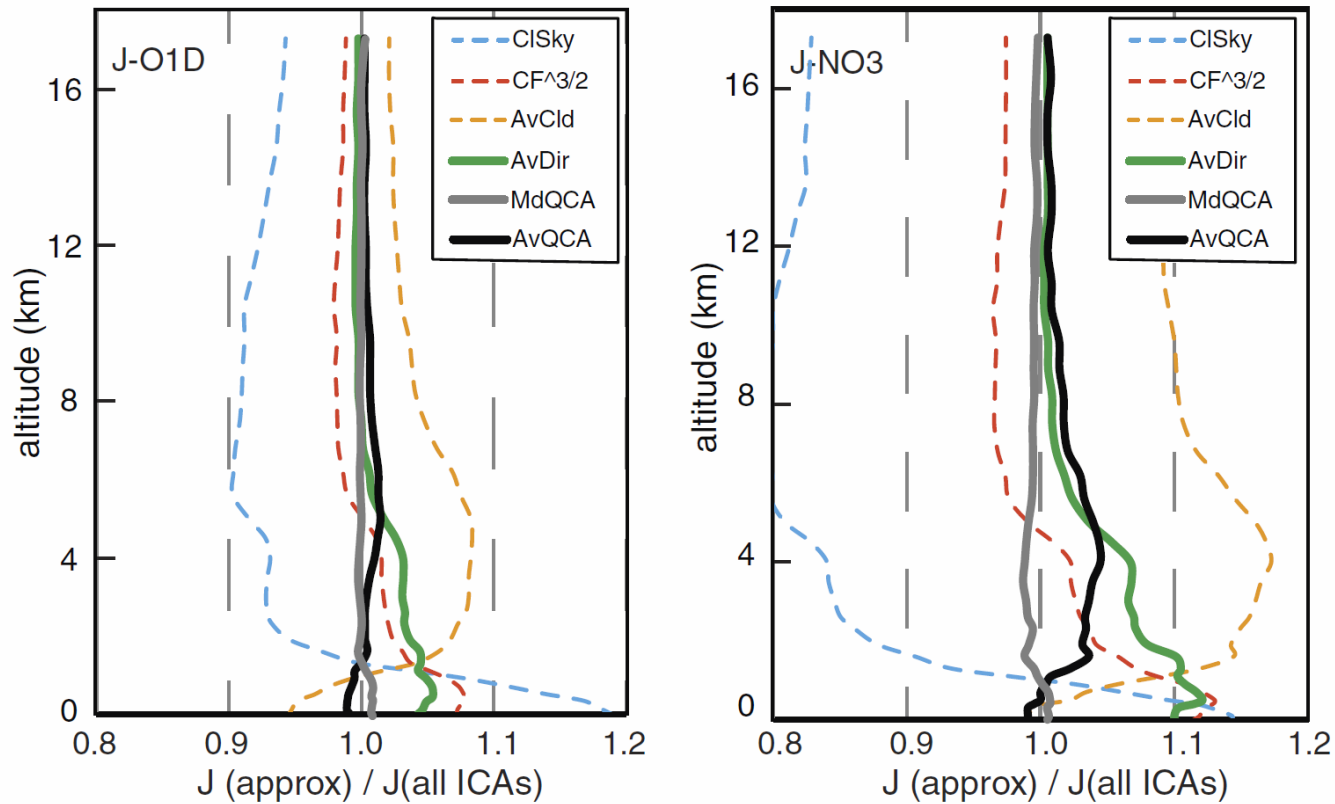
$$w^{G1}(\text{clear}^{G1}, \text{clear}^{G2}) = 1 - F^{G1} (1 - g F^{G2}) / (1 - F^{G2}) \quad (28)$$

Comparisons of # ICAs generated by: G6 = new MAX-COR model  
G3 = fixed 3-level MAX-RAN model  
G0 = MAX-RAN with breaks at every  $f = 0$



**Figure.** Number of Independent Column Atmospheres (ICAs) generated by three different cloud overlap models (G0, G3, G6) from 640 different tropical fractionally cloudy atmospheres (FCAs) and sorted in order of increasing ICA number.

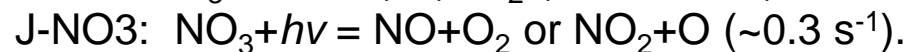
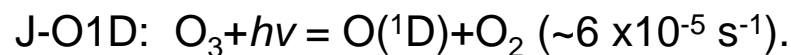
The different correlation coefficients used in the G6 model do not change the number of ICAs, only their weightings. The average number of ICAs per FCA is given in the legend.



**Figure.** J-value mean error (640 tropical T319L60 atmospheres) for various approximations used to average over all ICAs (total here = 108125). Reference model in G6 w/cc=0.33.

Three simple cloud methods (ClearSky, CldFr to 3/2 power, AvgCloud) do not use cloud-overlap model.

Three approximations for the ICAs are Median Quad Colm Atmos (Neu 2007), Average QCA and Average Direct Beam (both Prather 2015).



**Table 1.** J-value average and root-mean-square error for cloud overlap model (all ICAs)

Cloud overlap models to generate ICAs		ICAs <sup>a</sup>	avg err 0–1 km		rms err 0–1 km		rms err 0–16 km	
			J-O <sup>1</sup> D	J-NO <sub>3</sub>	J-O <sup>1</sup> D	J-NO <sub>3</sub>	J-O <sup>1</sup> D	J-NO <sub>3</sub>
G0	MAX-RAN with MAX groups bounded by layers with CF = 0	19	+2%	+2%	21%	17%	6%	11%
G3	3 MAX-RAN groups split at 1 km and at the ice-only cloud level	21	+2%	+2%	15%	15%	5%	7%
G6/.00	6 MAX-COR groups, cc = 0.00	169	-1%	-1%	5%	4%	2%	3%
G6/.33	6 MAX-COR groups, cc = 0.33 <sup>b</sup>	169	-	-	-	-	-	-
G6/.99	6 MAX-COR groups, cc = 0.99	169	+2%	+1%	11%	8%	4%	7%

All models are OK for large-scale averages  
(e.g., 640 atmospheres and 108,125 ICAs).

RMS errors are large for the small number MAX-RAN groups  
(G0, G3, with G6/.99 ~ 1 MAX group).

RMS errors for 6 MAX-COR with cc = 0.00 (effectively MAX-RAN) are OK.

Recognizing the correlation across the 6 MAX-COR groups makes a difference!

**Table 2.** J-value average and root-mean-square error for approximating all ICAs

			avg err 0–1 km		rms err 0–1 km		rms err 0–16 km	
			J-O <sup>1</sup> D	J-NO <sub>3</sub>	J-O <sup>1</sup> D	J-NO <sub>3</sub>	J-O <sup>1</sup> D	J-NO <sub>3</sub>
Simple cloud models			ICAs					
CISky	clear sky, ignore clouds	1	+14%	+10%	24%	20%	14%	23%
AvCld	average fractional cloud across layer	1	-5%	+1%	11%	11%	8%	15%
CF3/2	increase CF to CF <sup>3/2</sup> and average over layer	1	+7%	+11%	10%	15%	5%	8%
ICA approximations			J calls					
AvDir <sup>c</sup>	average direct beam from all ICAs	1	+5%	+11%	6%	13%	3%	7%
MdQCA <sup>c</sup>	Quadrature Column Atmospheres uses mid-point in each QCA	2.8	+1%	0%	4%	4%	4%	5%
AvQCA <sup>c</sup>	QCAs, uses average in each QCA	2.8	-1%	0%	3%	2%	2%	4%
Ran-3	Select 3 ICAs at random	3	+2%	+1%	12%	12%	9%	12%

Simple cloud models all have large errors.

Either QCA models are best.

Averaging the direct beam over all ICAs did not work as well as expected.

Picking random ICAs (instead of the QCAs) is poor choice for rms error.

## PREVIEW: Solar-J

Solar-J is Cloud-J extended beyond 800 nm with RRTMG-SW bands and thus will do solar heating

Solar-J includes direct and diffuse PAR (with 4-angle diffuse field)

Solar-J is funded by DOE BER and will be implemented as a drop-in replacement for Cloud-J, including cloud overlap and ice/liquid water absorption.

First ***off-line*** tests of Solar-J vs RRTMG-SW are now underway.



## PREVIEW: Solar-J

fast-J solar bins			
#	Wavelength (nm)	Solar (W/m <sup>2</sup> )	PAR (μE)
1	187	0.01	
2	191	0.02	
3	193	0.02	
4	196	0.01	
5	202	0.08	
6	208	0.04	
7	211	0.09	
8	214	0.11	
9	261	4.84	
10	267	2.97	
11	277	2.23	
12	295	3.97	
13	303	5.03	
14	310	3.23	
15	316	5.59	
16	333	22.98	3
17	380	80.45	125
18	412-850	696.40	2026

RRTM-SW solar bins			
Wavelength #	Wavelength range (nm)	Solar (W/m <sup>2</sup> )	
28	200 - 263	3.12	
27	263 - 345	50.15	
26	345 - 441	129.5	
25	441 - 625	347.2	
24	625 - 778	218.1	
23	778 - 1242	345.7	
22	1242 - 1299	24.29	
21	1299 - 1626	102.9	
20	1626 - 1942	55.63	
19	1942 - 2151	22.43	
18	2151 - 2500	23.73	
17	2500 - 3077	20.36	
16	3077 - 3846	12.11	
29	3846 - 12195	12.79	

*Fast-J would need to add 9 super-bins which amounts to ~40 added calculations*

*Overlap bin 18 vs 24-25 (14 bins) will become Fast-J bins 18a,b,c,d,e.*

## Preview: Solar-J

Juno Hsu has completed the first cut and adding RRTMG-SW to the Fast-J short-wavelength bins. The key overlap region is the last Fast-J bin #18 (442 – 778 nm) which overlaps 2 RRTM bins.

Fast-J now runs with the RRTM absorption longward of 442 nm.

Clear sky, high-sun comparison shown here:

